



# *Testing Einstein with Orbiting Gyroscopes*

*[+ some remarks on STEP]*

Symposium on Geometry  
& the Universe

Stony Brook

20-21 October 2005

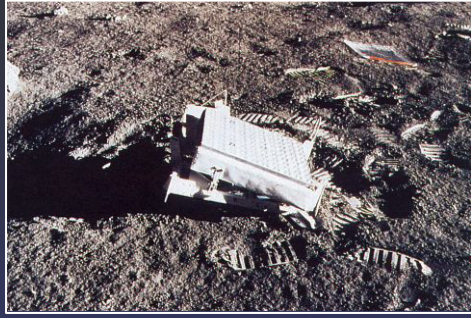
***Francis Everitt***



# Roadmap for a Gravity Probe

- GP-B & the drama of launch
- Basic experiment concept
- Near Zeroes & why we need them
- On-orbit
- Dither & aberration: 2 secrets of GP-B
- Calibration/verification
- Wider lessons, with some remarks on STEP

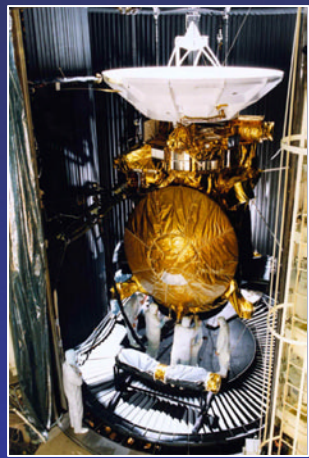
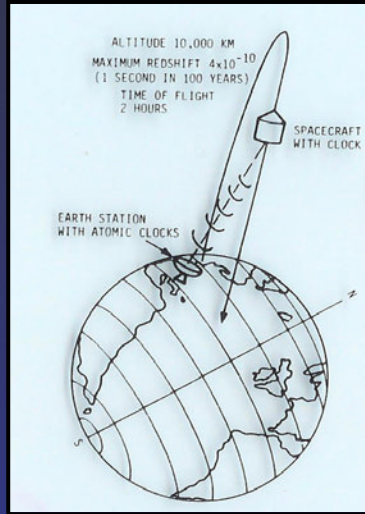
# Testing Einstein – NASA’s Contributions



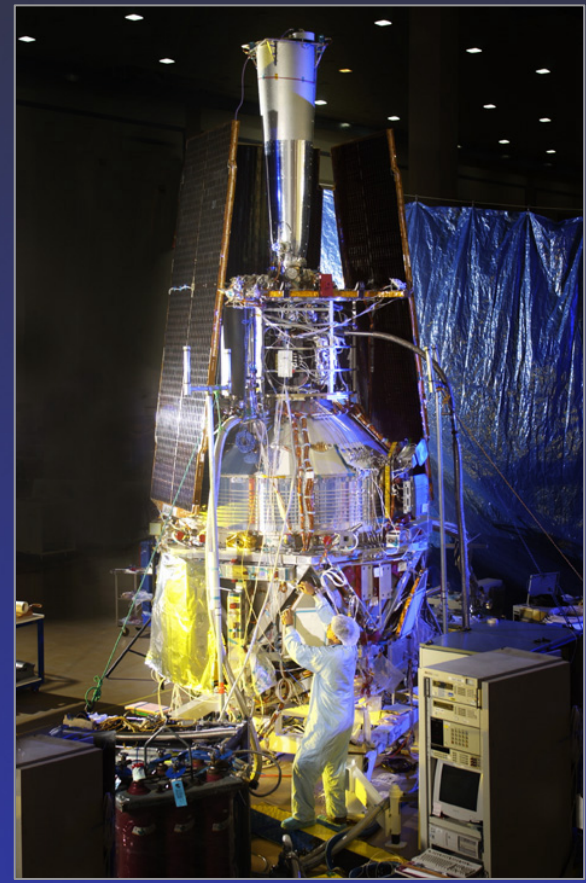
Laser Ranging: to reflectors on Moon (1968+)

Gravity Probe B  
Two new effects with ultra-accurate gyroscopes

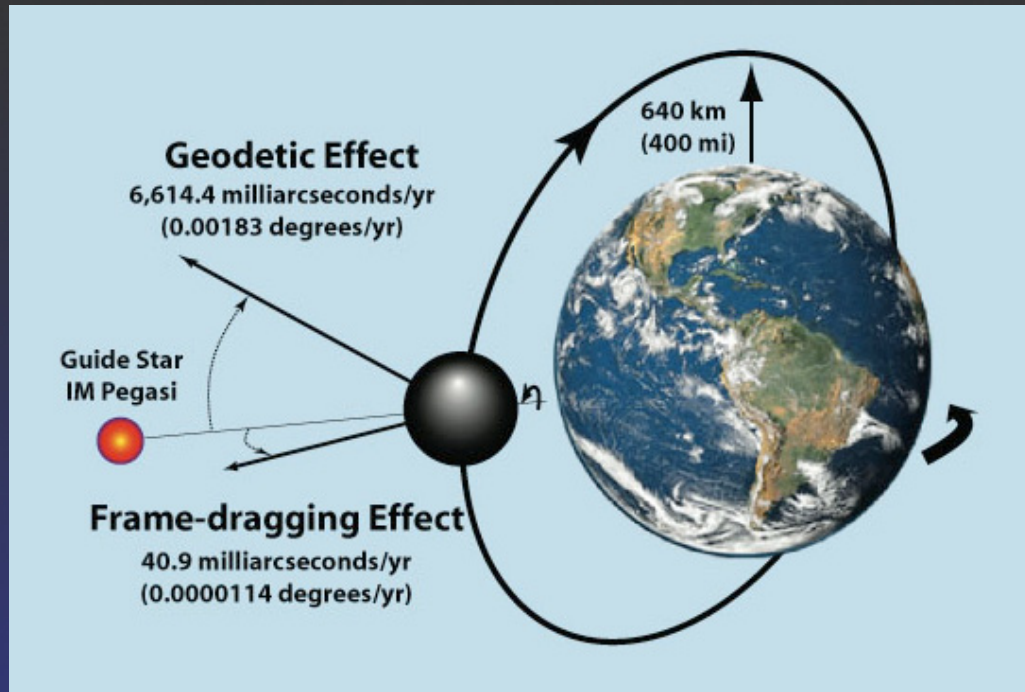
The Gravity Probe A clock experiment (1976)



Radar Time Delay:  
to Viking Lander on Mars (1976)  
to Cassini spacecraft around Saturn (1999+)



# The Relativity Mission Concept



"If at first the idea is not absurd,  
then there is no hope for it."  
-- A. Einstein

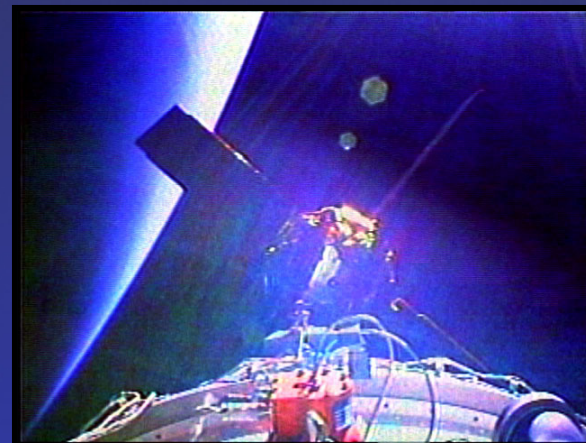


- Basic formula: *Leonard Schiff*

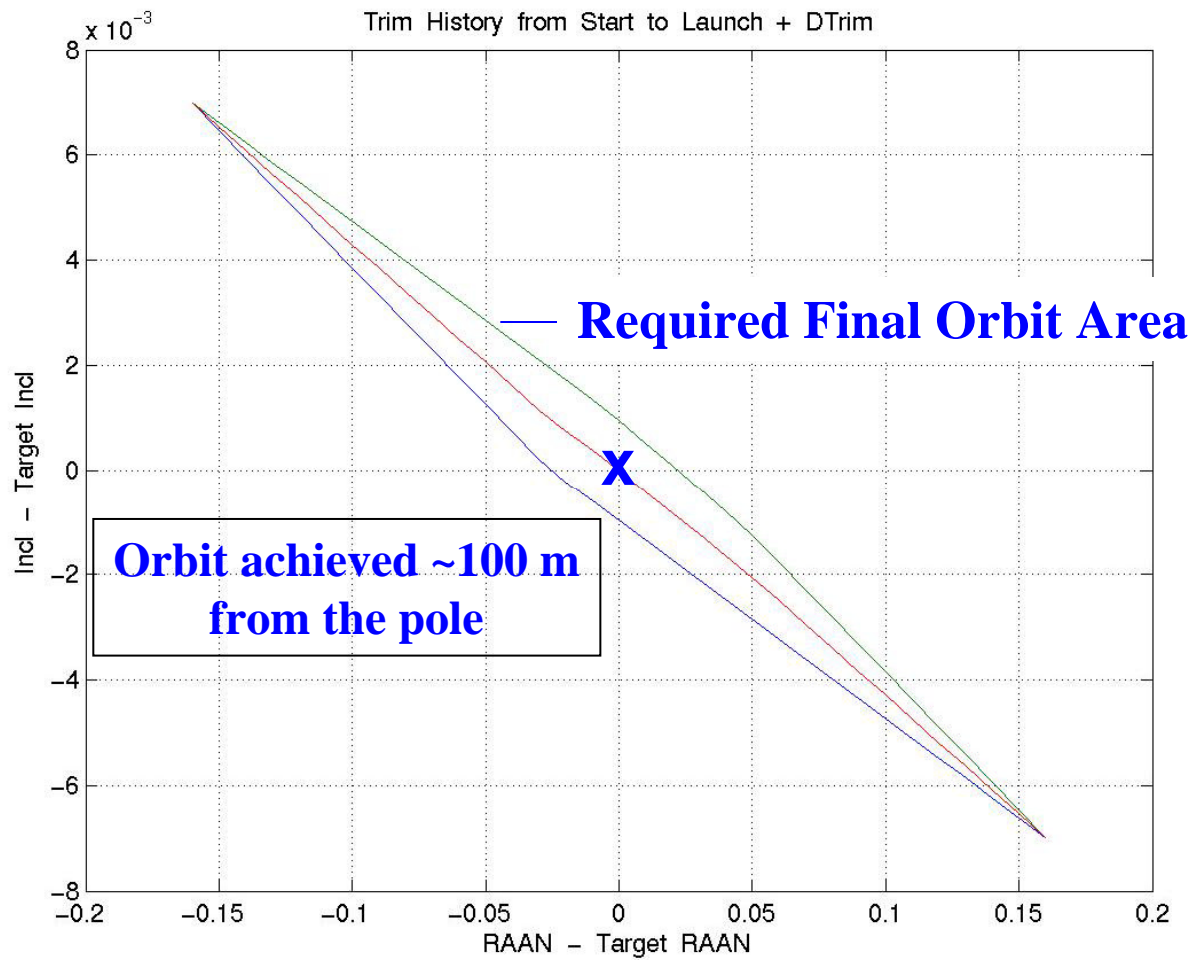
$$\Omega = \frac{3GM}{2c^2 R^3} (\mathbf{R} \times \mathbf{v}) + \frac{GI}{c^2 R^3} \left[ \frac{3\mathbf{R}}{R^2} (\boldsymbol{\omega} \cdot \mathbf{R}) - \boldsymbol{\omega} \right]$$

- Oblateness correction: *\* Dan Wilkins (Physics), John Breakwell (Aero/Astro)*

# Launch: April 20, 2004 – 09:57:24

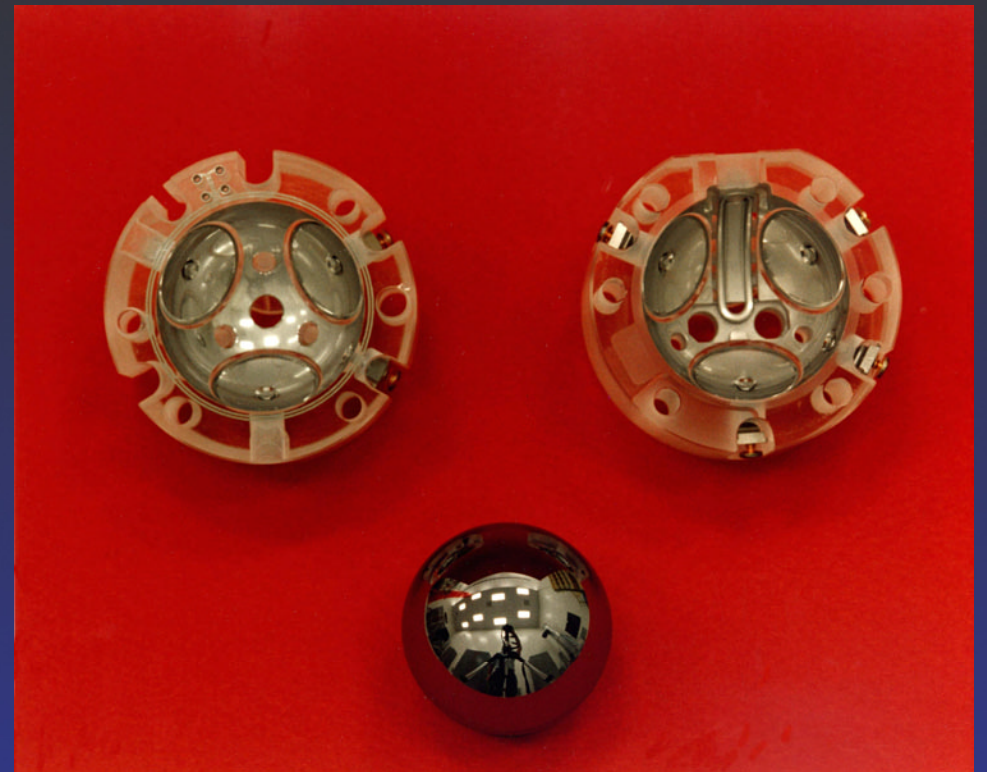
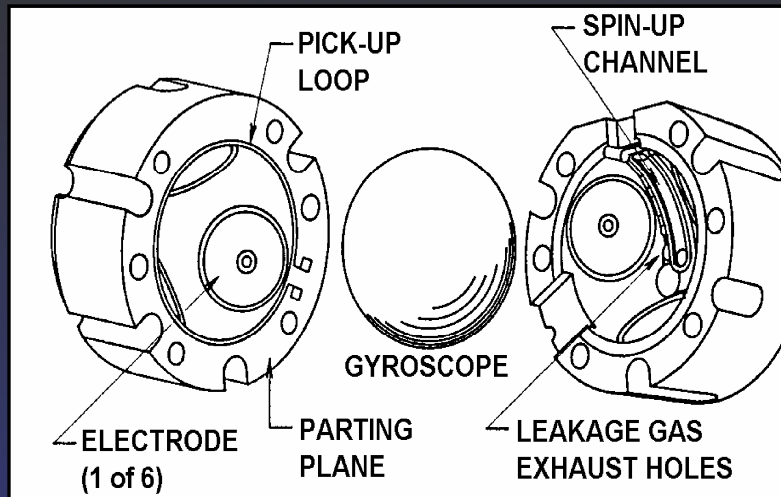


# Boeing & Luck -- A Near Perfect Orbit



**Delta II Nominal Accuracy**

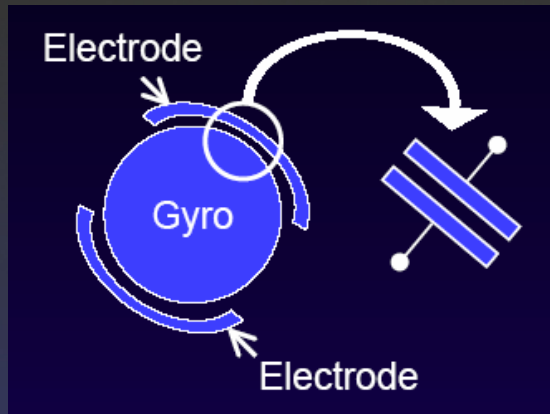
# Gyro I: Overview



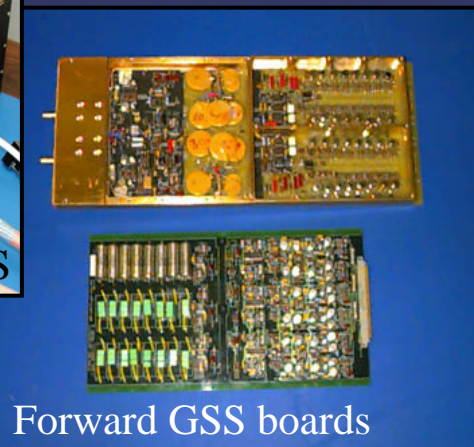
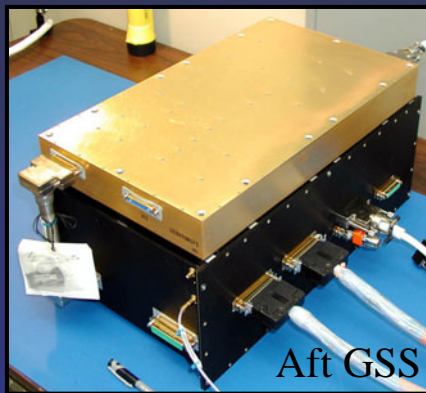
- **Electrical Suspension**
- **Gas Spin-up**
- **Magnetic Readout**

*"Everything should be made as simple as possible, but not simpler."*  
-- A. Einstein

# Gyro II: Suspension Characteristics



- Operates over 9 orders of magnitude of g levels
- Range of motion within cavity (15,000 nm) for:
  - science (centered in cavity)
  - spinup (offset to spin channel ~ 11,000 nm)
  - calibration (offset, 200 nm increments)
- Alignment (roll phased voltage variation)



Analog ground-based version:  
John Nikirk, Dick Van Patten &  
John Gill (Aero/Astro)

Digital flight version:

- \* Bill Bencze (EE) & joint Stanford-Lockheed Martin team, including 3 Aero/Astro, 2 EE PhDs & 6 undergraduates (4 departments)

Nanometer references {

- thickness of sheet of paper ~ 100,000 nm
- diameter of atom ~ .1 to .5 nm

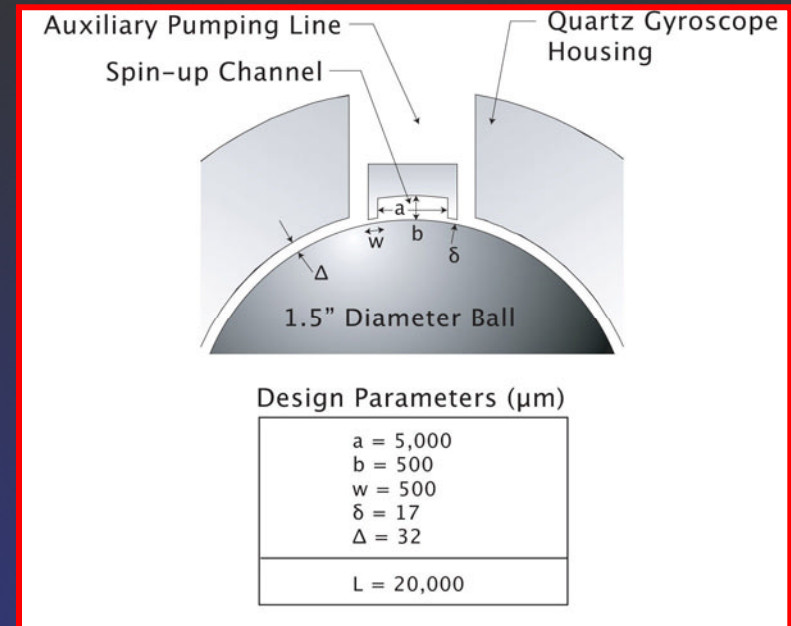


# Gyro III: The Spin-up Problem(s)

## Torque Switching Requirement

$$T_r/T_s < \Omega_0 t_s \sim 10^{-14}$$

$T_s, T_r$  - spin & residual cross-track torques  
 $t_s$  - spin time;  $\Omega_0$  - drift requirement



## Differential Pumping Requirement

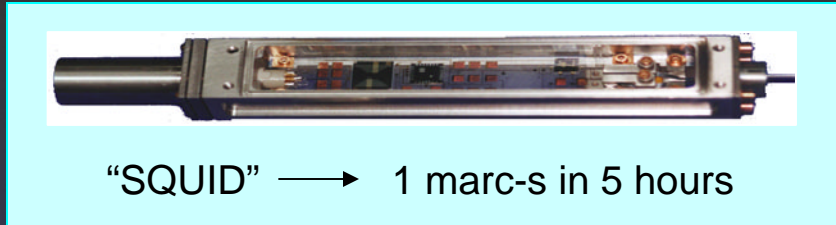
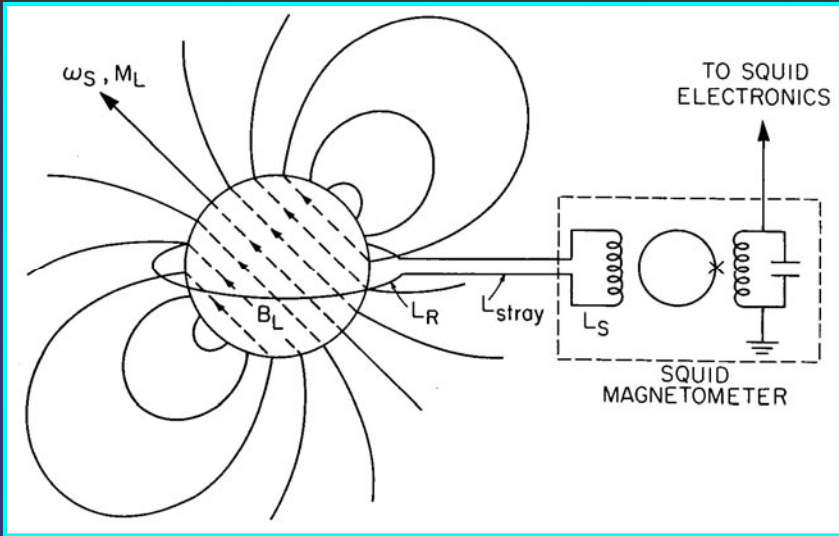
spin channel ~ 10 torr (sonic velocity)  
 electrode area <  $10^{-3}$  torr



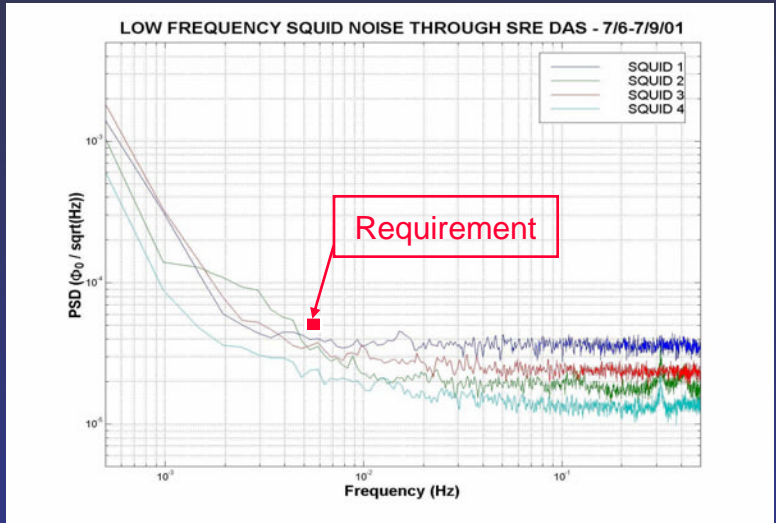
\* Dan Bracken (Physics)  
 Don Baganoff (Aero/Astro)  
 + refinements by John Lipa, John Turneure & several students

*"Any fool can get the steam into the cylinders; it takes a clever man to get it out again afterwards." -- G. J. Churchward, ~ 1895*

# Gyro IV: London Moment Readout

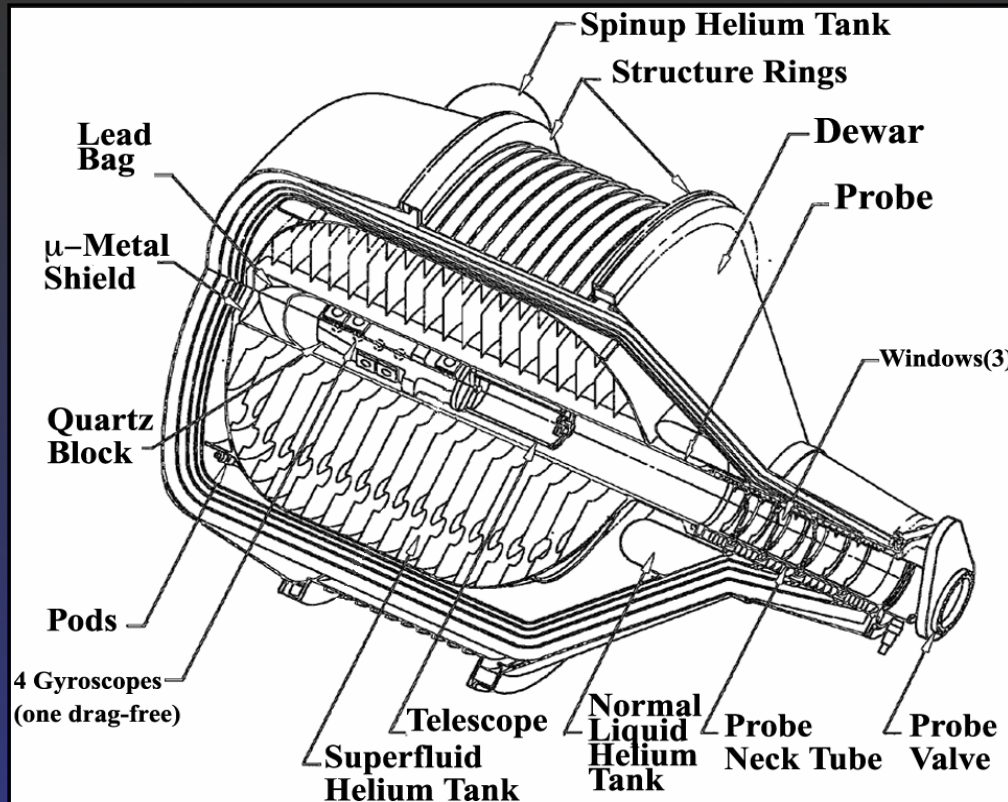


Jim Lockhart (\* Physics & SFSU)  
 Barry Muhlfelder (HEPL)  
 \* Greg Gutt & \* Ming Luo (EE)  
 Bruce Clarke (HEPL)  
 Terry McGinnis (Lockheed)  
 + many more



- ◆ Noise performance
- ◆ DC trapped flux <  $10^{-6}$  gauss
- ◆ AC shielding >  $10^{12}$
- ◆ Centering stability < 50 nm

# The GP-B Cryogenic Payload



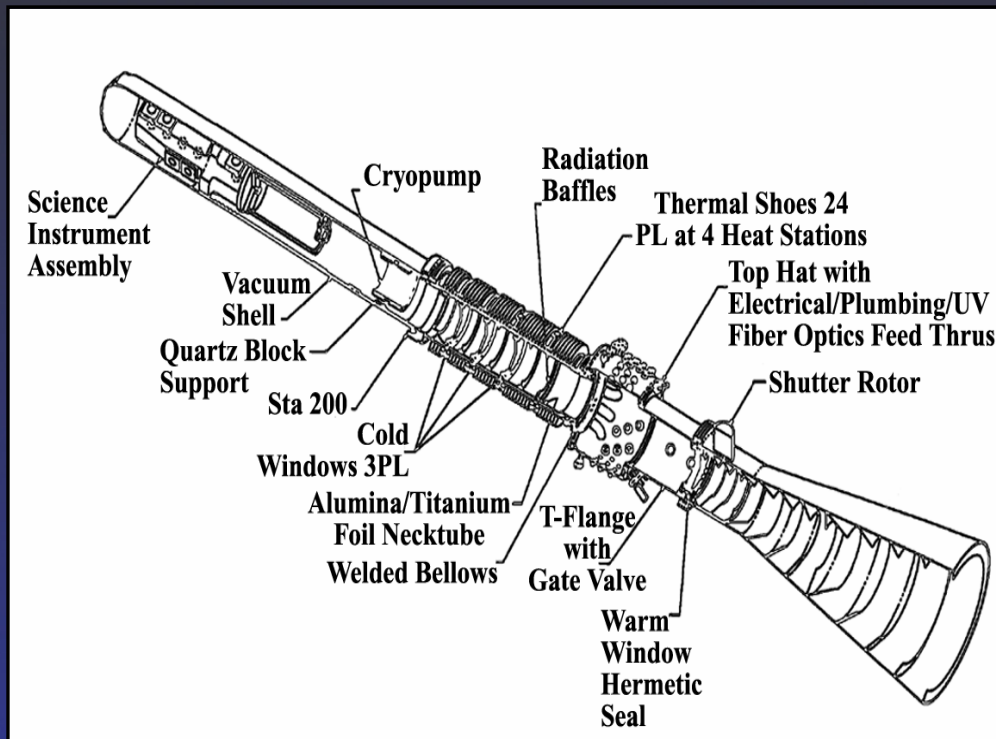
*Two notable doctoral dissertations:*

- \* Peter Selzer (Physics) - porous plug for space
- \* John McCuan (Math) - helium tidal studies

Payload in ground testing at Stanford, August 2002

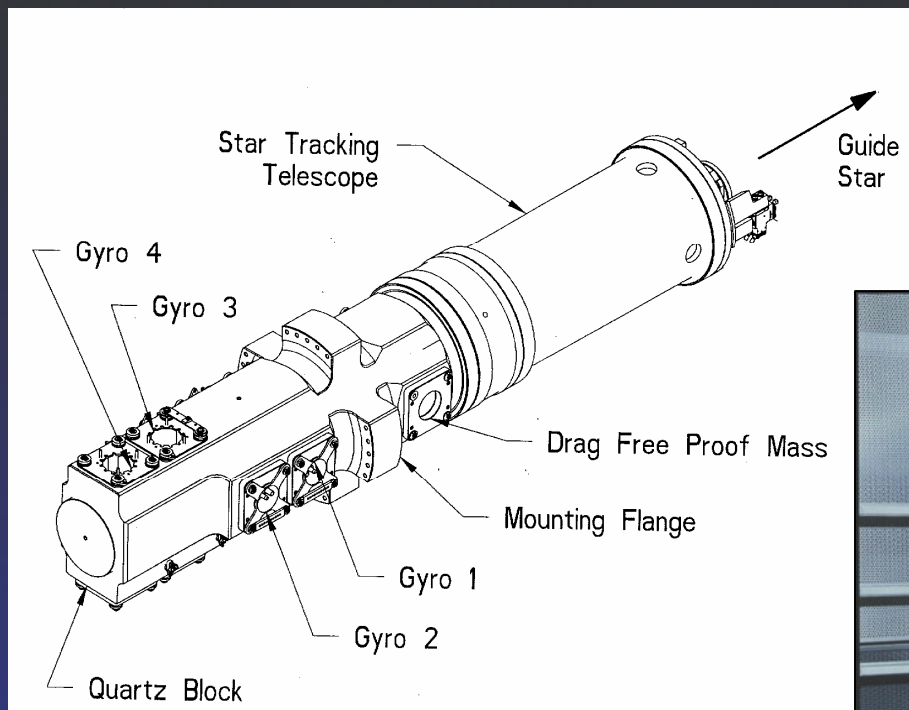
# The GP-B Flight Probe

Lockheed Martin Lead: Gary Reynolds



Assembled probe at Lockheed prior to shipment to Stanford

# Alignment, Bonding & Cryogenic Stability



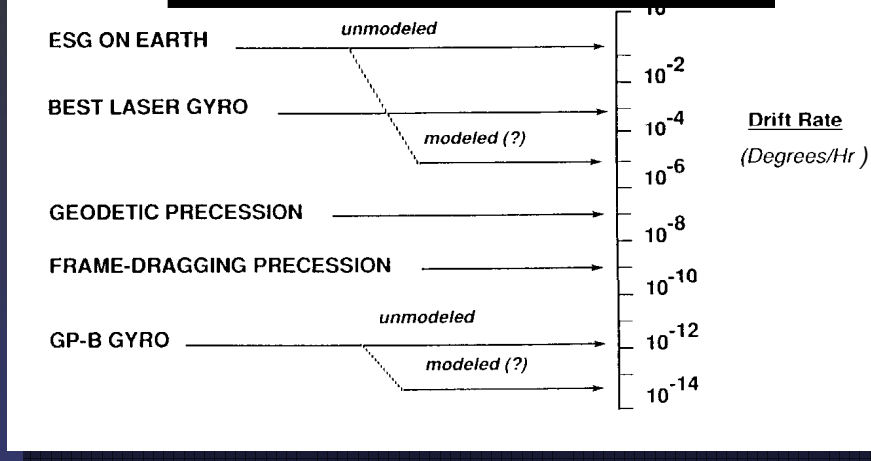
*"Design a precision apparatus as if it were made of jello – if it is stable then, it may just work."*  
-- H.A. Rowland, ~1900



Assembly & alignment: *Doron Bardas (Physics), \* Robert Brumley (EE!)*  
Silicate bonding: *Jason Gwo (Berkeley Chemistry!)*

# Near Zeros & Why We Need Them

## Why go to space?



$1 \text{ marcsec/yr} = 3.2 \times 10^{-11} \text{ deg/hr}$

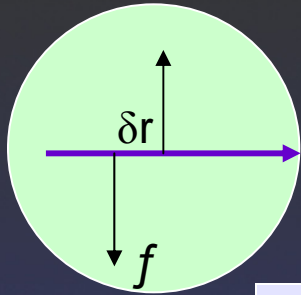
- ### Six Crucial Near Zeros
- 1) rotor inhomogeneities
  - 2) "drag-free"
  - 3) rotor asphericity
  - 4) magnetic field
  - 5) pressure
  - 6) electric charge

GP-B Co-PIs:  
 Brad Parkinson  
 Dan DeBra  
 John Turneaure

0.1 marcsec is the width of a human hair seen from 100 miles

- Gyroscope drift  $\leq 0.05 \text{ marcsec/yr}$
- Readout error effect  $\leq 0.08 \text{ marcsec/yr}$
- Guide star proper motion uncertainty  $\leq 0.09 \text{ marcsec/yr}$

# Mass-Unbalance, Drag-Free: 1st & 2nd Near Zeros



Drift-rate  $\Omega = T / I\omega_s$

Torque  $T = M f \delta r$

Moment of Inertia  $I = 2Mr^2 / 5$

**requirement**  $\Omega < \Omega_0 \sim 0.1 \text{ marc-s/yr}$  ( $1.54 \times 10^{-17} \text{ rad/s}$ )

$$f \frac{\delta r}{r} < \frac{2}{5} v_s \Omega_0$$

$$v_s = \omega_s r = 950 \text{ cm/s} \quad (80 \text{ Hz})$$

On Earth ( $f = g$ )

$$\frac{\delta r}{r} < 5.8 \times 10^{-18}$$

(ridiculous)

Standard satellite ( $f \sim 10^{-8} g$ )

$$\frac{\delta r}{r} < 5.8 \times 10^{-10}$$

(unlikely)

GP-B drag-free ( $f \sim 10^{-12} g$   
cross-track average)

$$\frac{\delta r}{r} < 5.8 \times 10^{-6}$$

(straightforward)

GP-B rotor  $\frac{\delta r}{r} \sim 3 \times 10^{-7}$

drift-rate for the  
drag-free GP-B  
< 0.01 marc-s/yr

Drag-free eliminates mass-unbalance torque -- and key to understanding/quantification of other support torques

# Asphericity: 3<sup>rd</sup> Near Zero -- Making

- Self-aligning laps
- Uniform rotation-rate, pressure
- 6 combinations of directions, reversed 2 & 2 every 6 seconds
- Continuous-feed lapping compound
- Controlled pH
- Interested, skilled operators!

## MSFC

Wilhelm Angele

John Rasquin

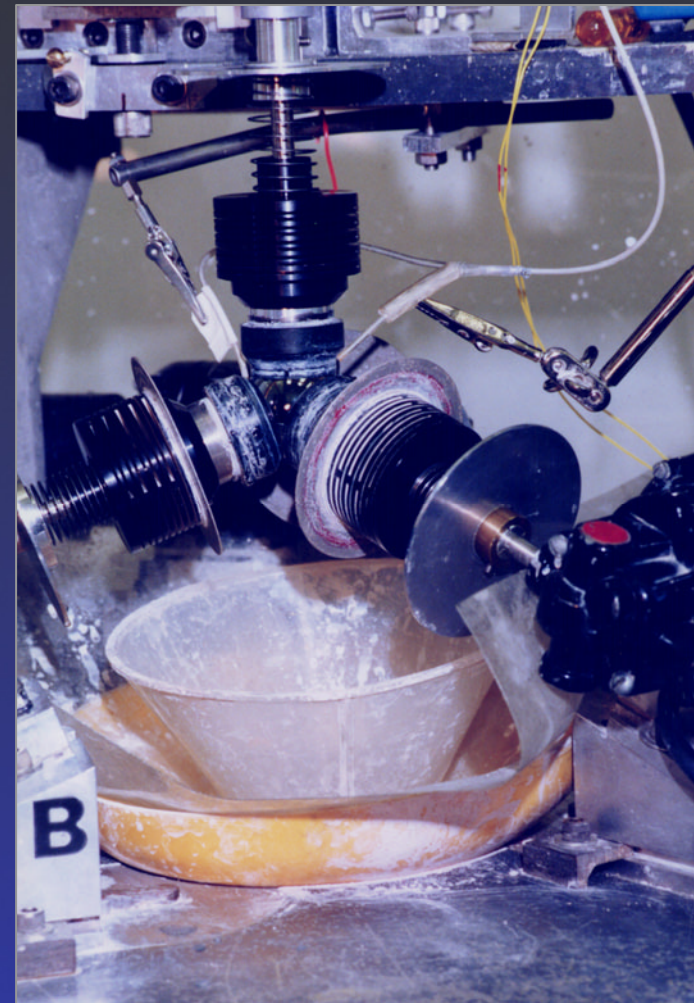
Ed White

## STANFORD

Thorwald van Hooydonk

Frane Marcelja

Victor Graham (visitor)

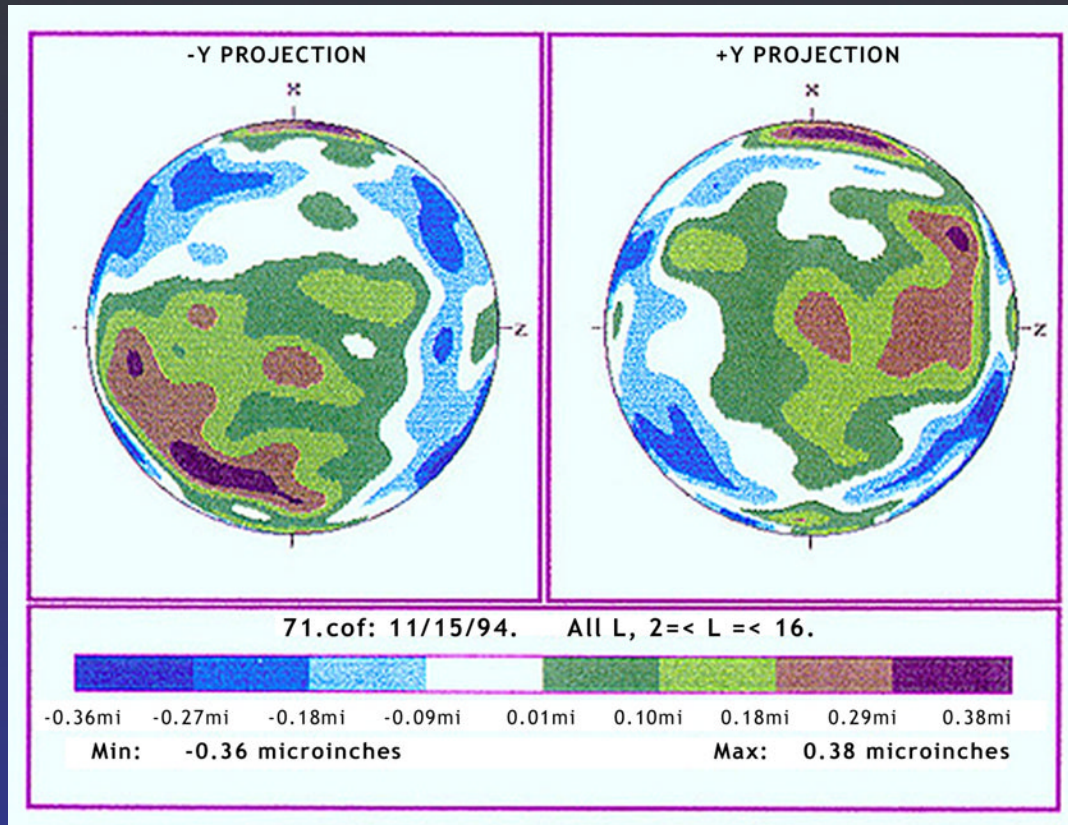




# Asphericity: 3<sup>rd</sup> Near Zero -- Measuring

## Students 1988 - 1992

- \* Grace Chang (A/A)
- \* Rebecca Eades (Math)
- \* Benjamin Lutch (undeclared)
- \* Dave Schleicher (Comp Sci)
- \* Dieter Schwarz (EE)
- \* Michael Bleckman (Hamburg)
- \* Christoph Willsch (Göttingen)



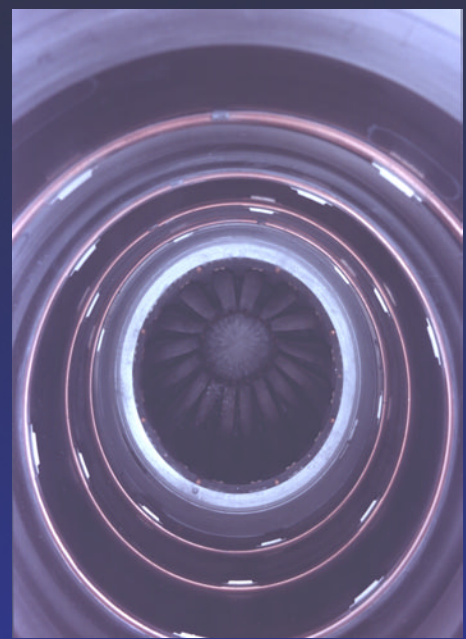
Roundness Measurement to ~ 1 nm →



## Superconducting Lead Bag Technology

- flux = field x area
- successive expansions  
    ➔ stable field levels  $\sim 10^{-7}$  gauss
- $10^{-12}$  [ =120 dB! ] ac shielding through combination of cryoperm, lead bag, local superconducting shields & symmetry

\* Blas Cabrera (1976 Physics PhD)  
*Dewar bag*  
Jim Lockhart (\* Physics, SFSU),  
Mike Taber (\* Physics, HEPL),  
Chuck Warren, Dave Murray  
*Magnetic material testing*  
John Mester, Grace Brauer





# On-Orbit: GP-B Mission Operations

Program Manager – Gaylord Green

## Mission Operations Center



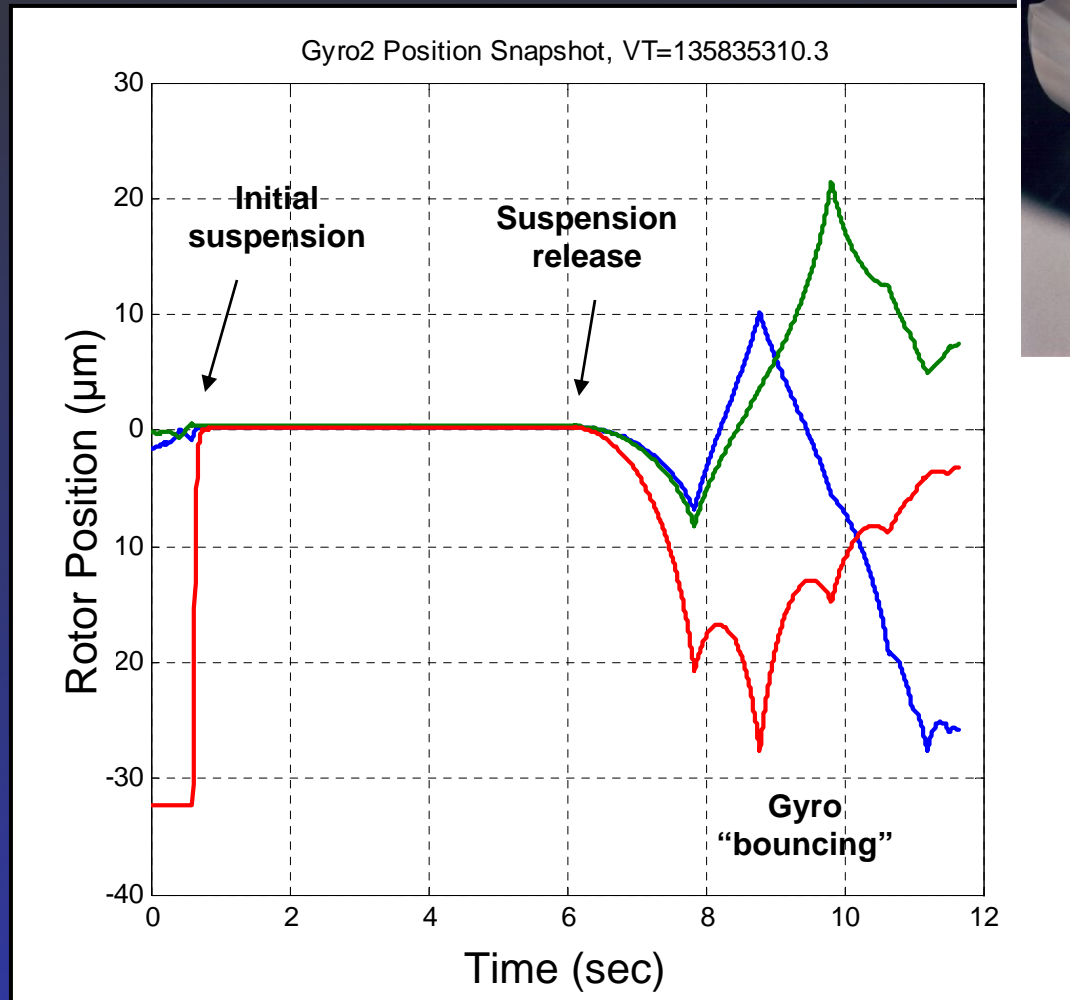
### Anomaly Room

- Marcie Smith (NASA Ames)
- Kim Nevitt (NASA MSFC)
- Rob Nevitt (NavAstro)
- Brett Stroozas (NavAstro)
- Lewis Wooten (NASA MSFC)
- Ric Campo (Lockheed Martin)
- Jerry Aguinado (LM)
- + many more



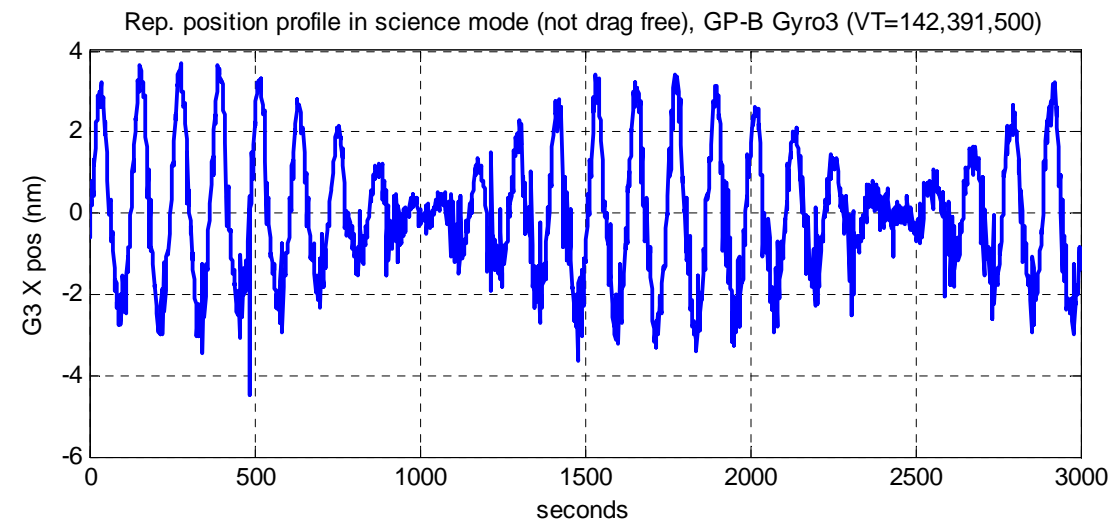
# GP-B Gyro On-Orbit Initial Liftoff

Initial Gyro Levitation and De-levitation using analog backup system

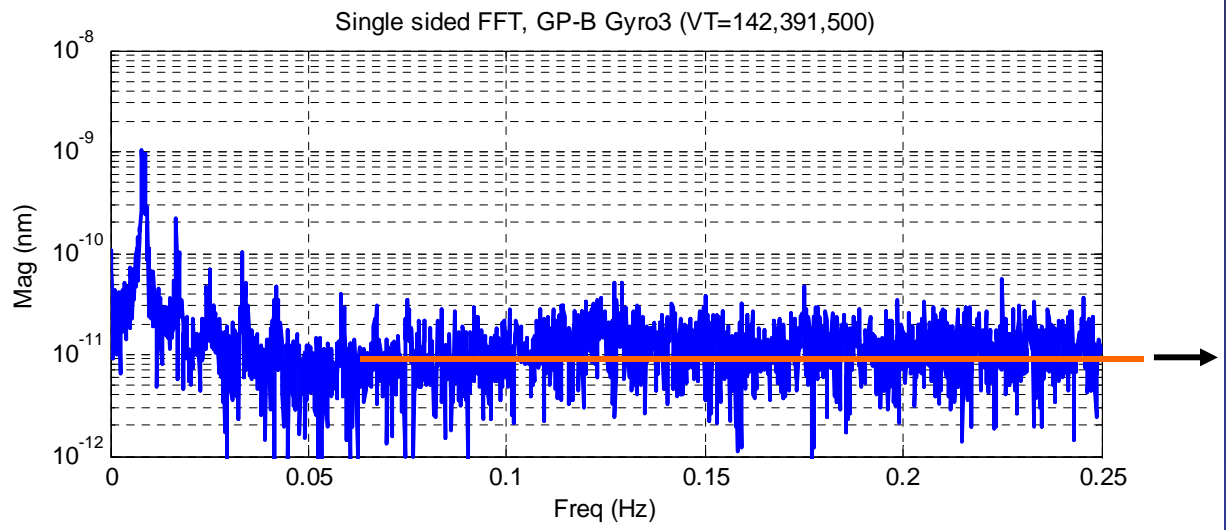


David Hipkins (HEPL)  
 \* Yoshimi Ohshima (A/A)  
 Steve Larsen (LM)  
 Colin Perry (LM)  
 + many more!

# Suspension Performance On-Orbit



Gyro position –  
non drag-free gravity  
gradient effects in  
Science Mission Mode

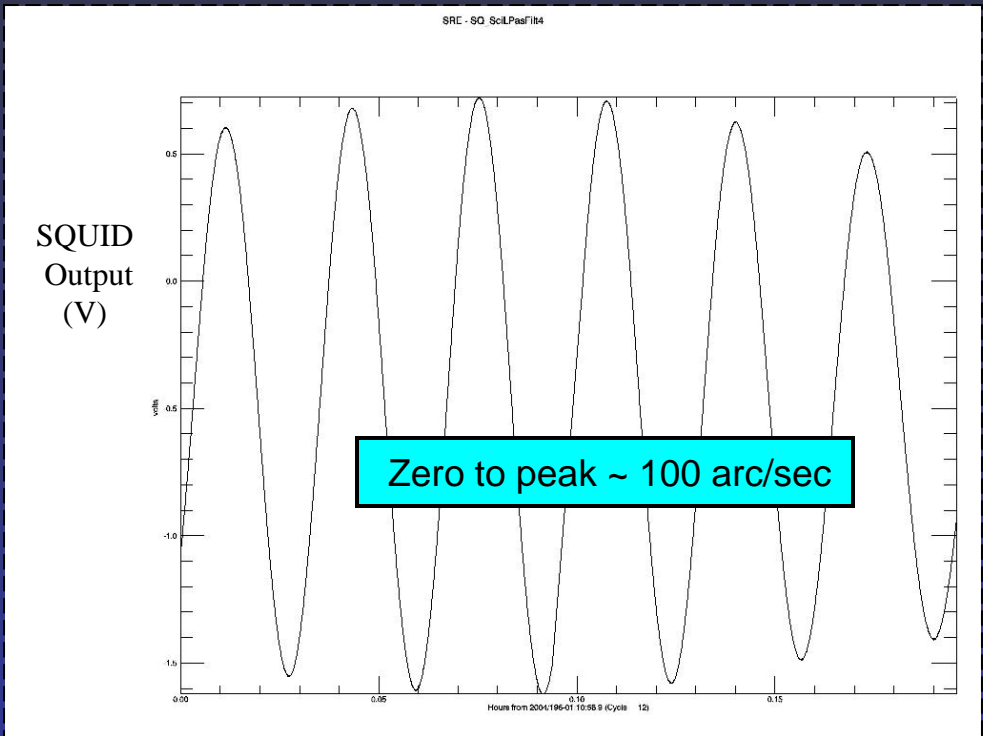
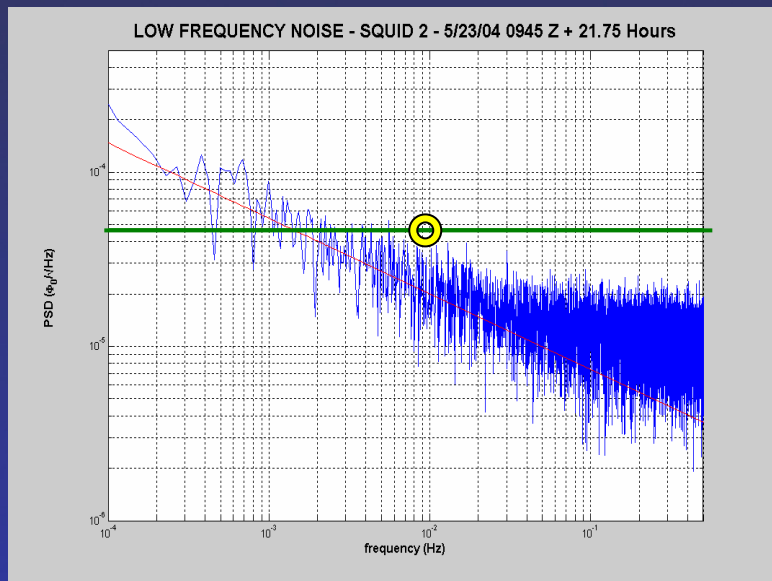
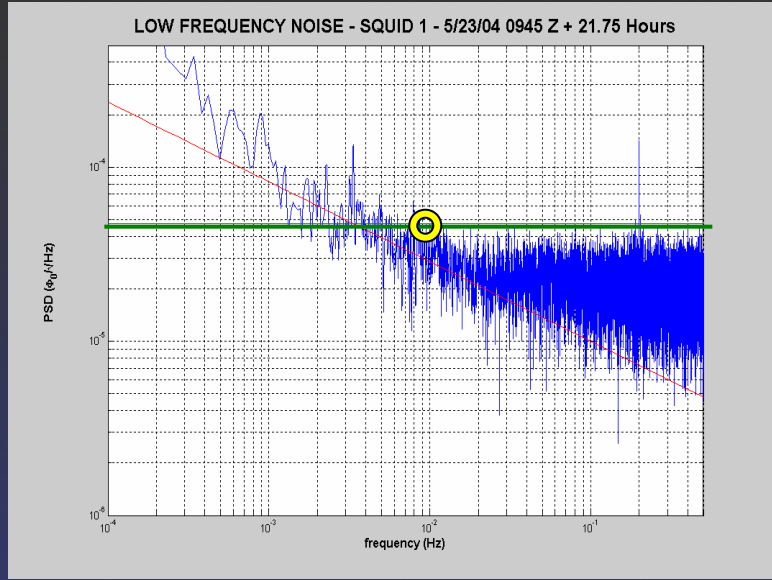


Measurement noise –  
0.45 nm rms

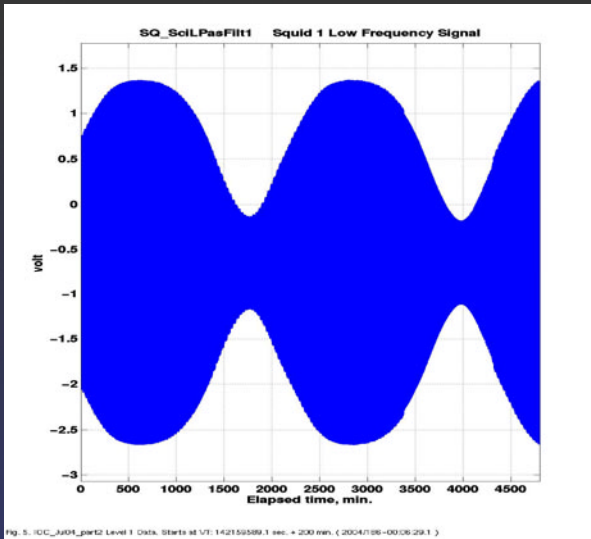
Noise floor

# Gyro Readout Performance On-Orbit I

Bruce Clarke, Barry Muhlfelder + the team



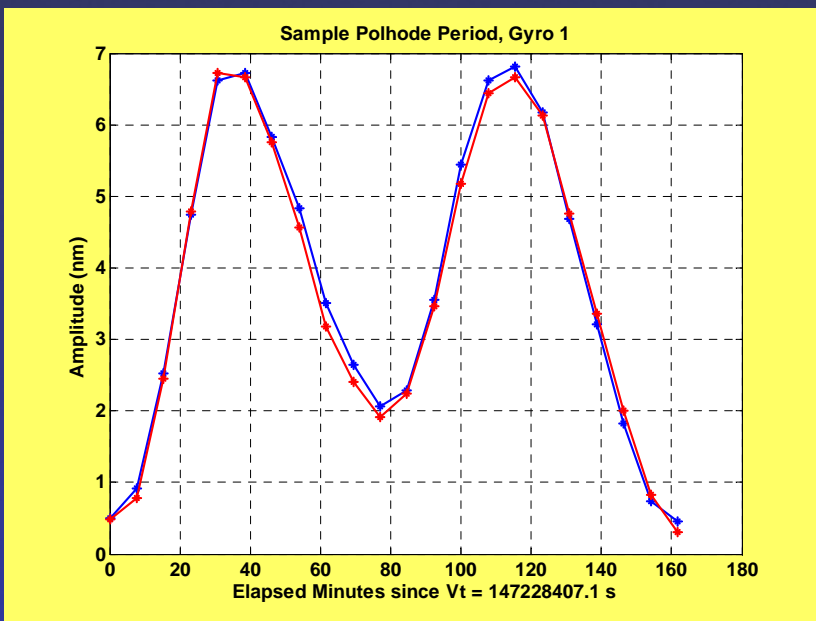
# Mass Unbalance (& $\frac{\Delta I}{I}$ ): 1<sup>st</sup> Near Zero



Gyro # 1 @ 3 Hz  
36-hour Polhode Period

$$\frac{\Delta I}{I} < 2 \times 10^{-6}$$

Mac Keiser & Paul Shestople + the team



Gyro # 1 @ 79.3858 Hz

Mass Unbalance (nm)

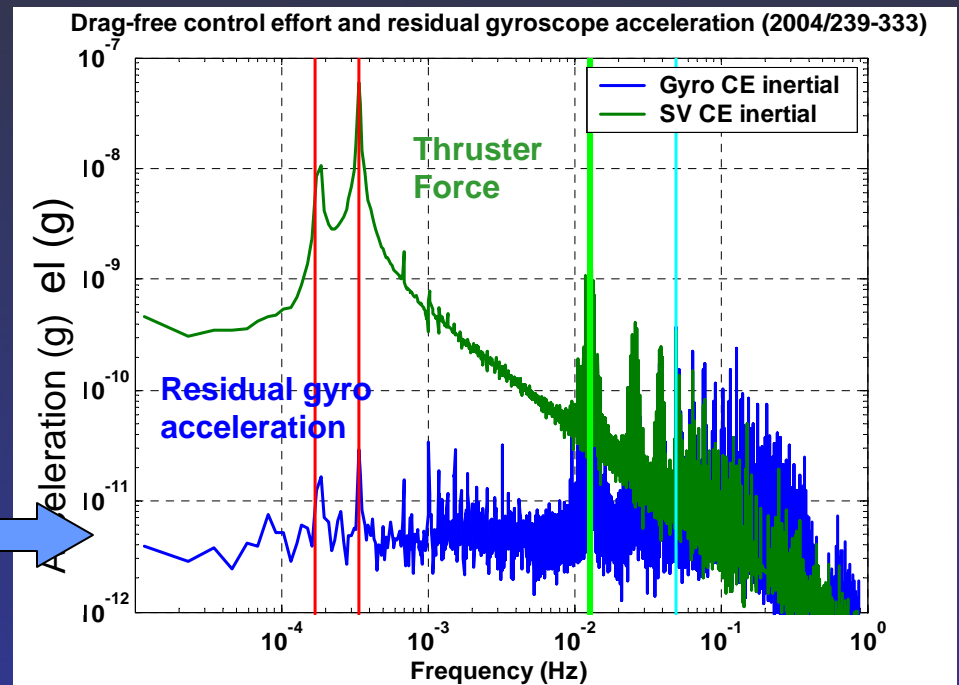
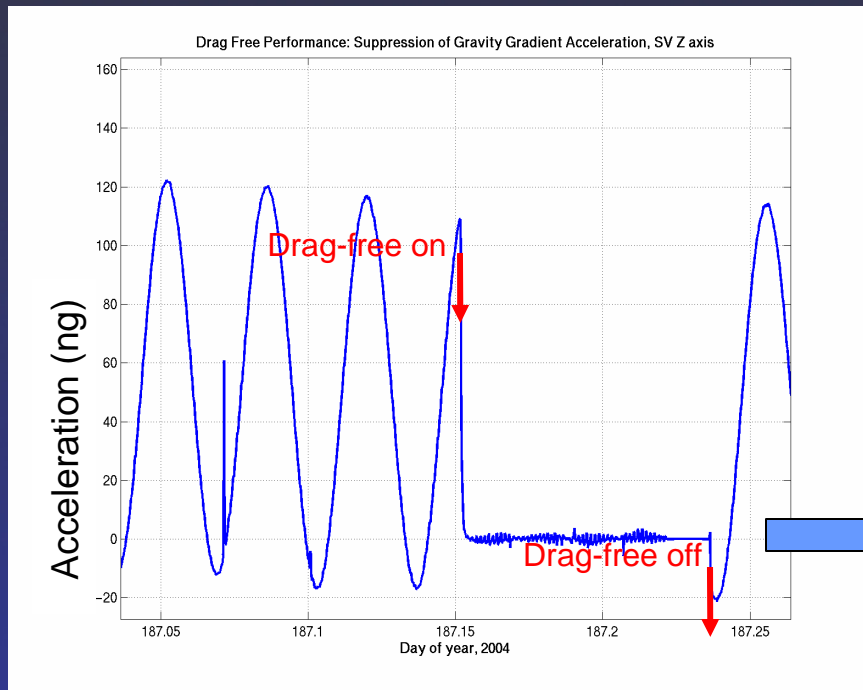
Gyro #	1	2	3	4
Prelaunch estimate	18.8	14.5	16.8	13.5
On-orbit data	6.9	4.4	3.3	6.0

# Drag-Free: 2<sup>nd</sup> Near Zero

Lockheed Martin Attitude/Translational Control  
 Design Lead: **Jon Kirschenbaum**

Toward guide star

Cross track



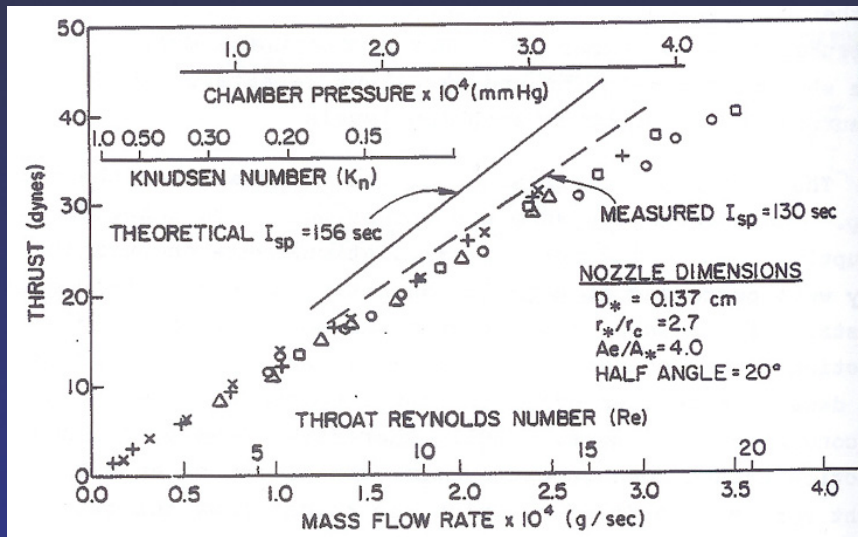
Demonstrated accelerometer (drag free) performance better than  $10^{-11}$  g DC to 1 Hz



# Boil-off, Altitude & Thrust -- A Subtle Combination

- A very different control system
  - Continuous flow → proportional thrusters
  - Reynolds' #  $\rho v l / \eta \sim 10!!$  -- flowing like honey

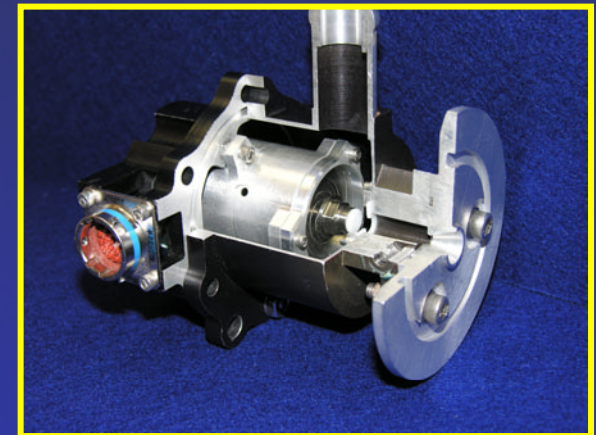
• Thrust calibration: \* John Bull + \* Jen Heng Chen (A/A)



← He specific impulse vs. mass flow rate

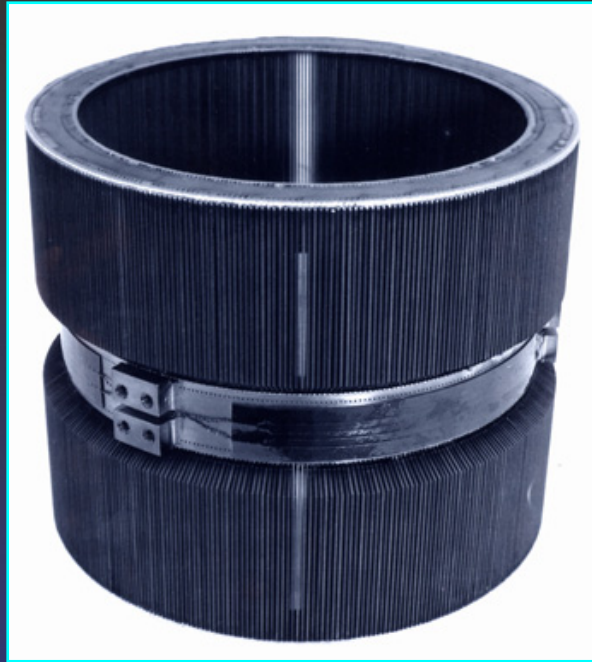
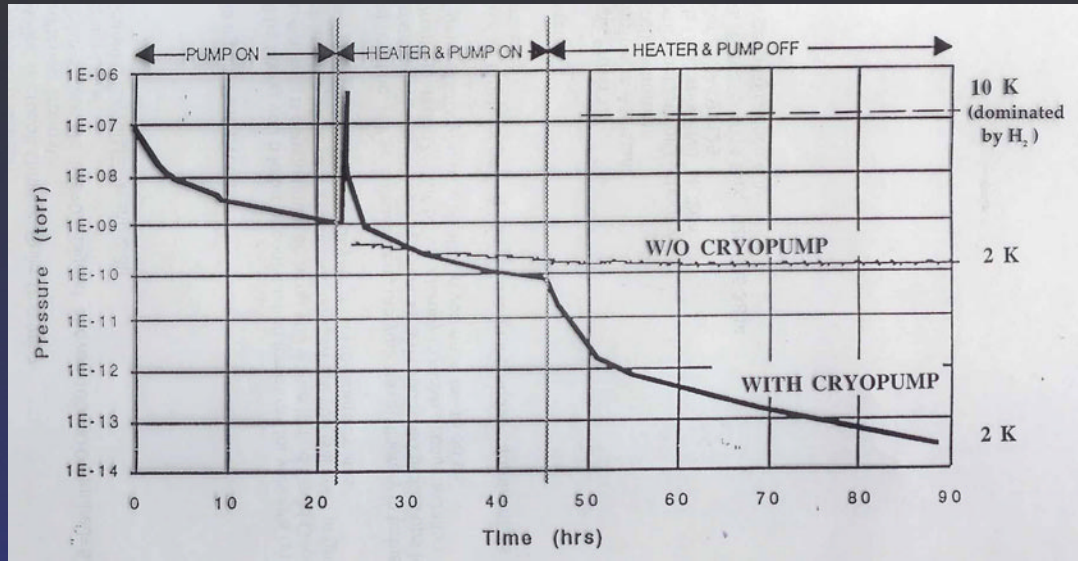
• Lockheed Martin thrusters: Jeff Vanden Beukel

\* Yusuf Jafry (A/A) with LM team



# Ultra-low Pressure: 5<sup>th</sup> Near Zero

## Low Temperature Bakeout (ground demonstration)



The Cryopump

## Gyro spindown periods on-orbit (years)

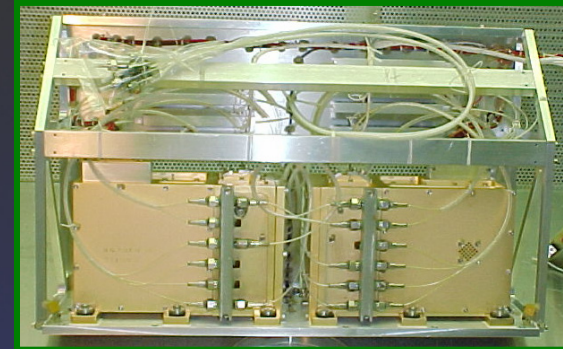
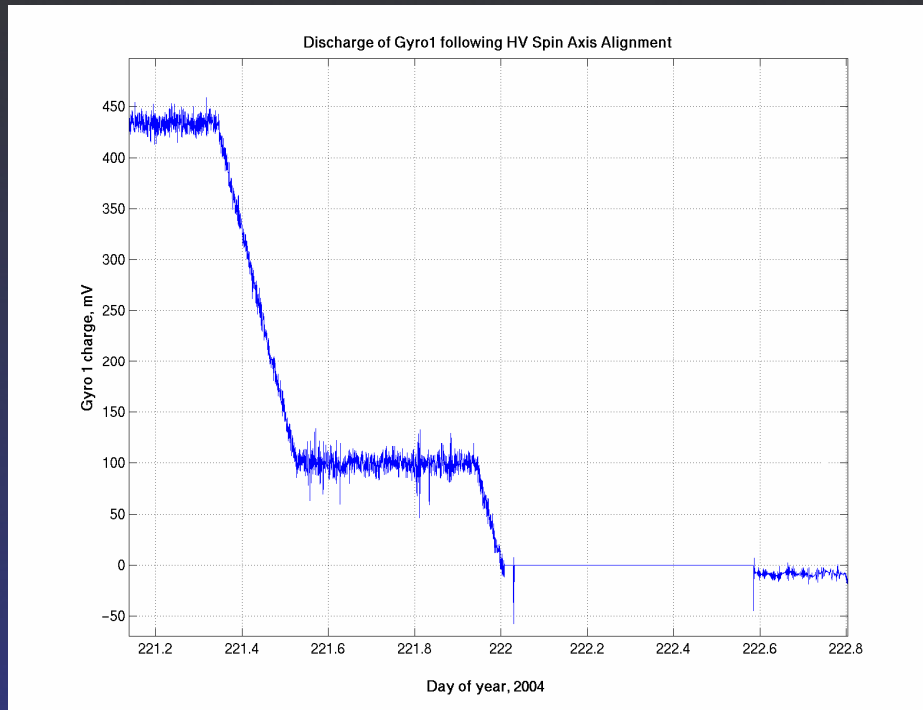
	before bakeout	after bakeout
Gyro #1	~ 50	15,800
Gyro #2	~ 40	13,400
Gyro #3	~ 40	7,000
Gyro #4	~ 40	25,700

John Lipa, John Turneaure (Physics) + students; adsorption isotherms for He at low temperature,\* Eric Cornell, (undergraduate honors thesis)

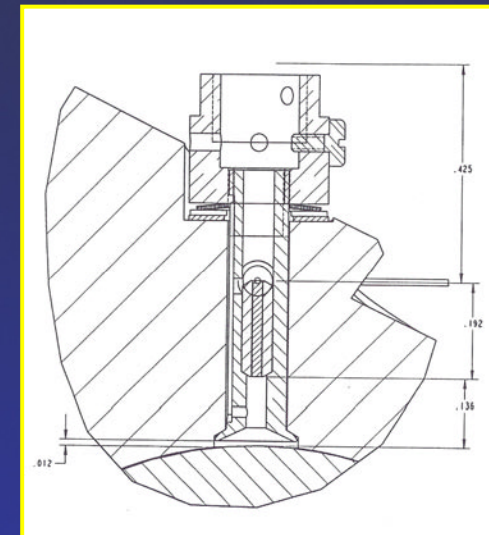
pressure < 1.5 x 10<sup>-11</sup> torr  
(+ minute eddy-current damping effects?)

# Rotor Electric Charge: 6<sup>th</sup> Near Zero

## Discharge of Gyro #1



## Ti Steering Electrode



Saps Buchman, Dale Gill, Bruce Clarke (Physics, HEPL)  
+ \* Brian DiDonna & \* Ted Quinn (Physics)

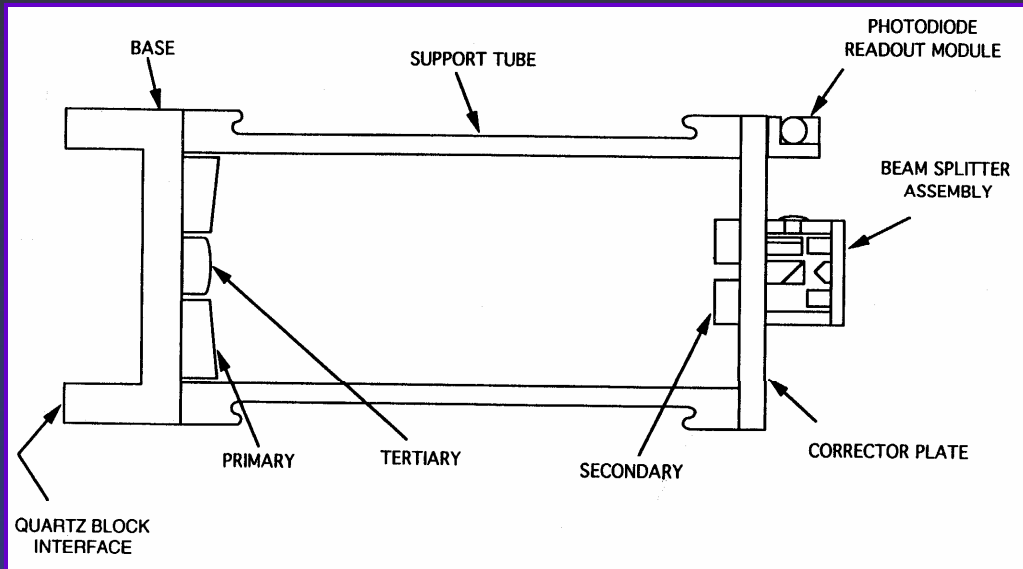
Typical charge rates ~ 0.1 mV/day

# Star Tracker I: Concept

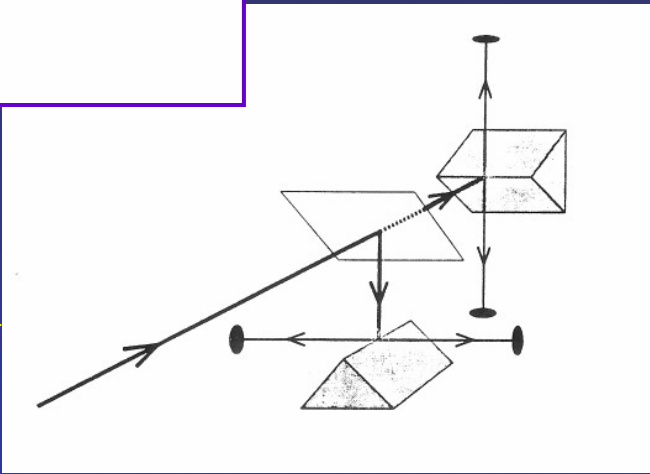
Design Lead: Don Davidson, Davidson Optronics, Inc. & OID

## Some dimensions

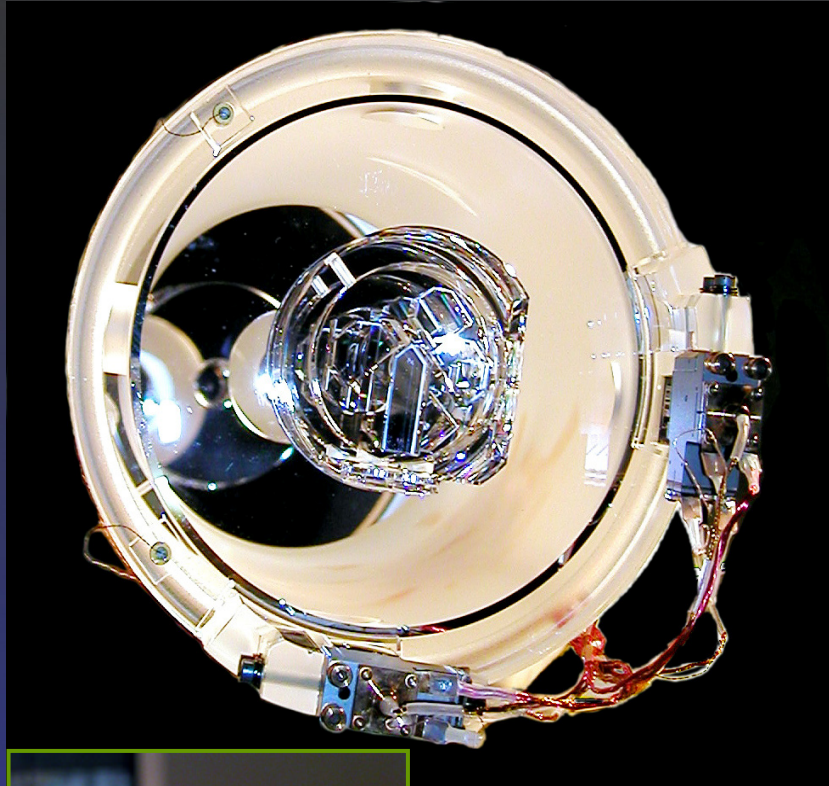
Physical length	0.33 m
Focal length	3.81 m
Aperture	0.14 m
<u>At focal plane</u>	
Image dia.	50 $\mu$ m
0.1 marc-s	0.18 nm



Beam splitter assembly (detail) →



# Star Tracker II: Under Test



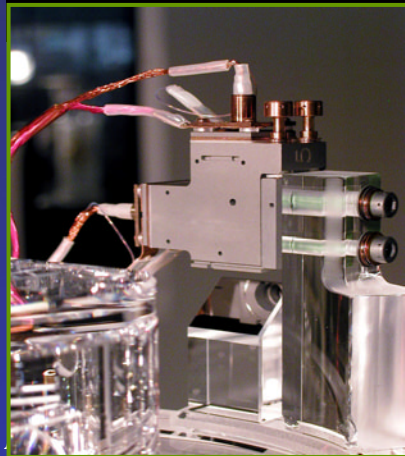
John Lipa, Jason Gwo, Suwen Wang  
(Physics, HEPL), Bob Farley (Lockheed),  
John Goebel (NASA Ames)

*Telescope development*

\* Mo Badi (Ap Phys), \* Dana Clark (ME),  
\* Chris Cumbermack (Pre-med!),  
\* Howard Shen (EE) + 6 others

*Artificial Star #3*

\* Ted Acworth, \* Rob Bernier



Detector  
Package



Si Diode Detector

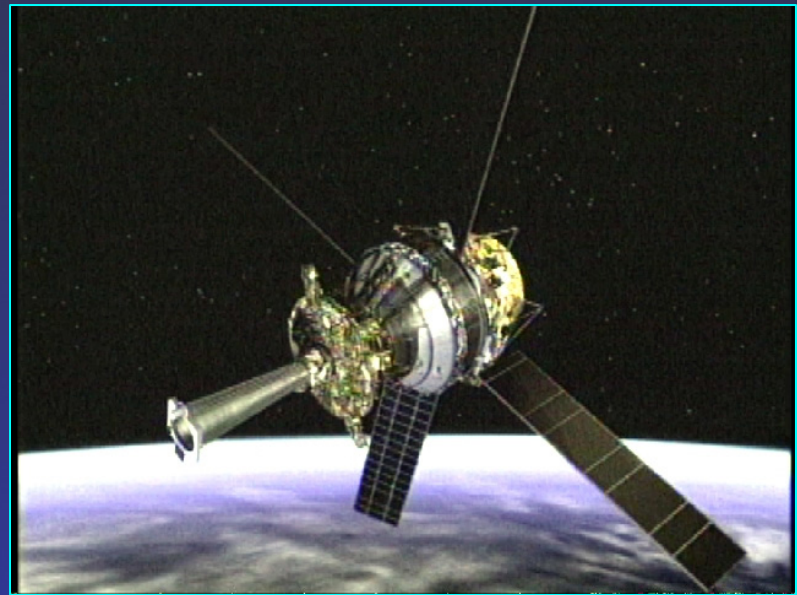
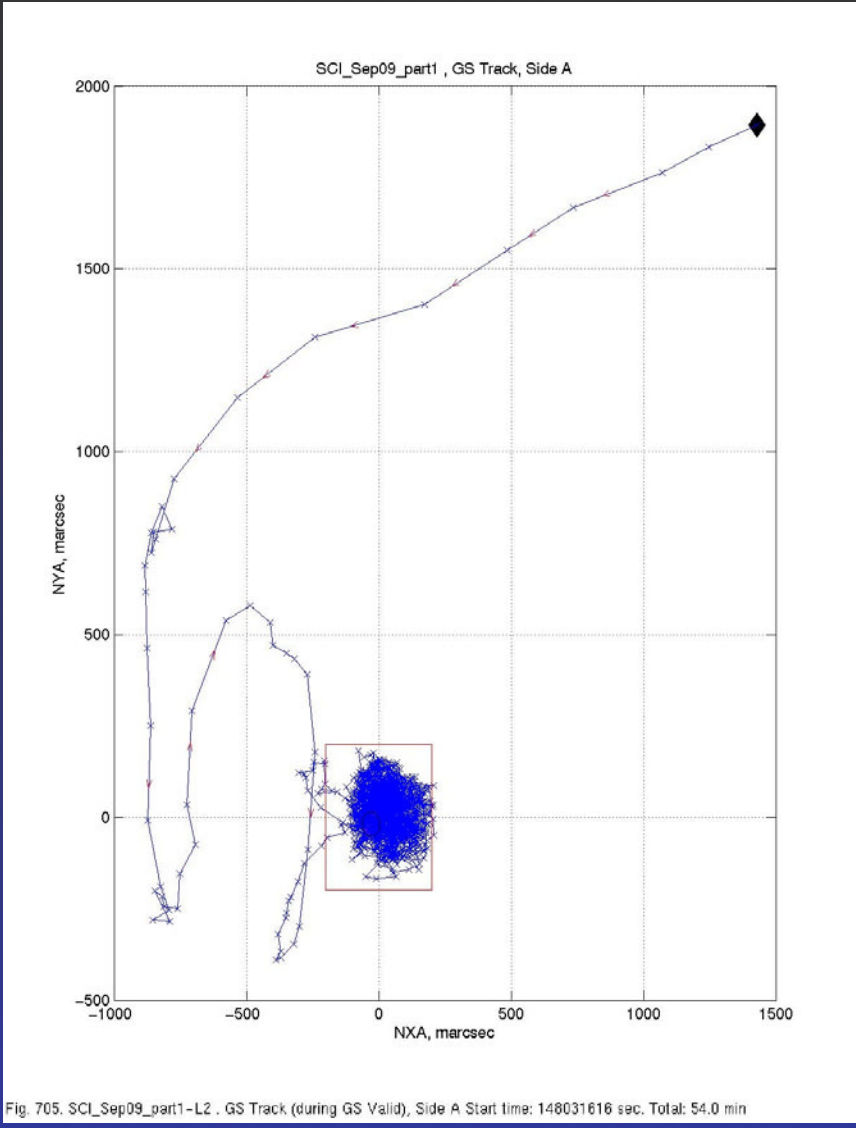
Artificial  
Star #3



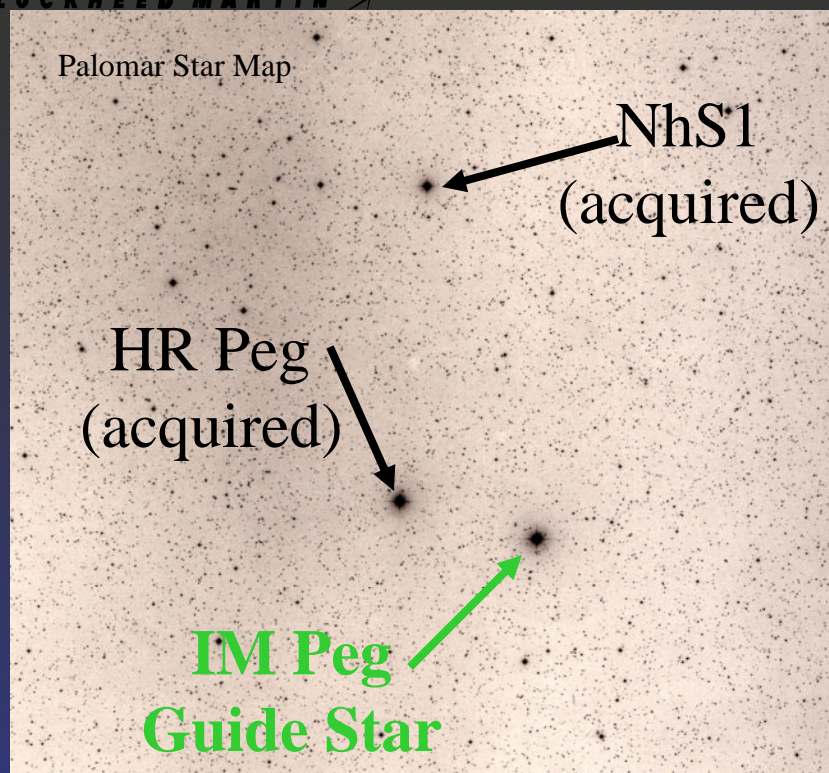
# Star Tracker III: Acquiring Star

Drive-in time ~ 110 s

RMS pointing ~ 90 marc-s



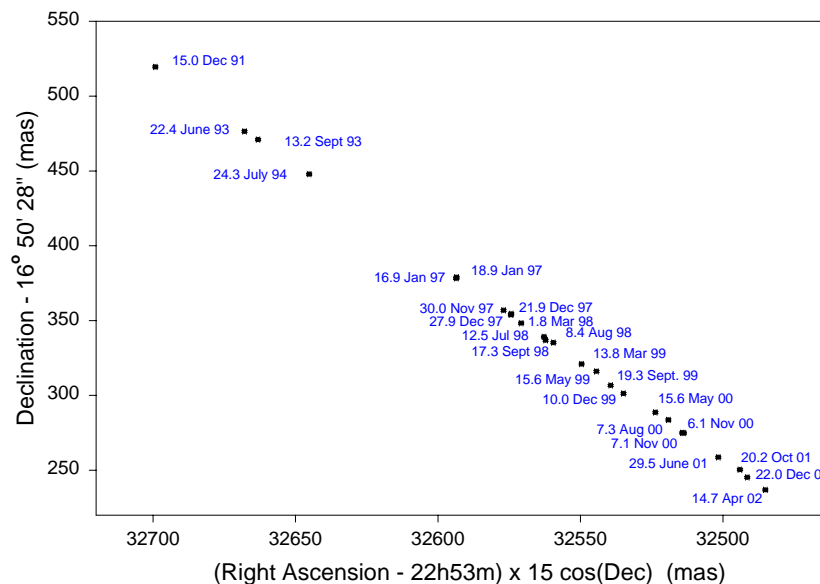
# IM Peg (HR 8703) Guide Star Identification



*Very Large Array, Socorro, New Mexico*

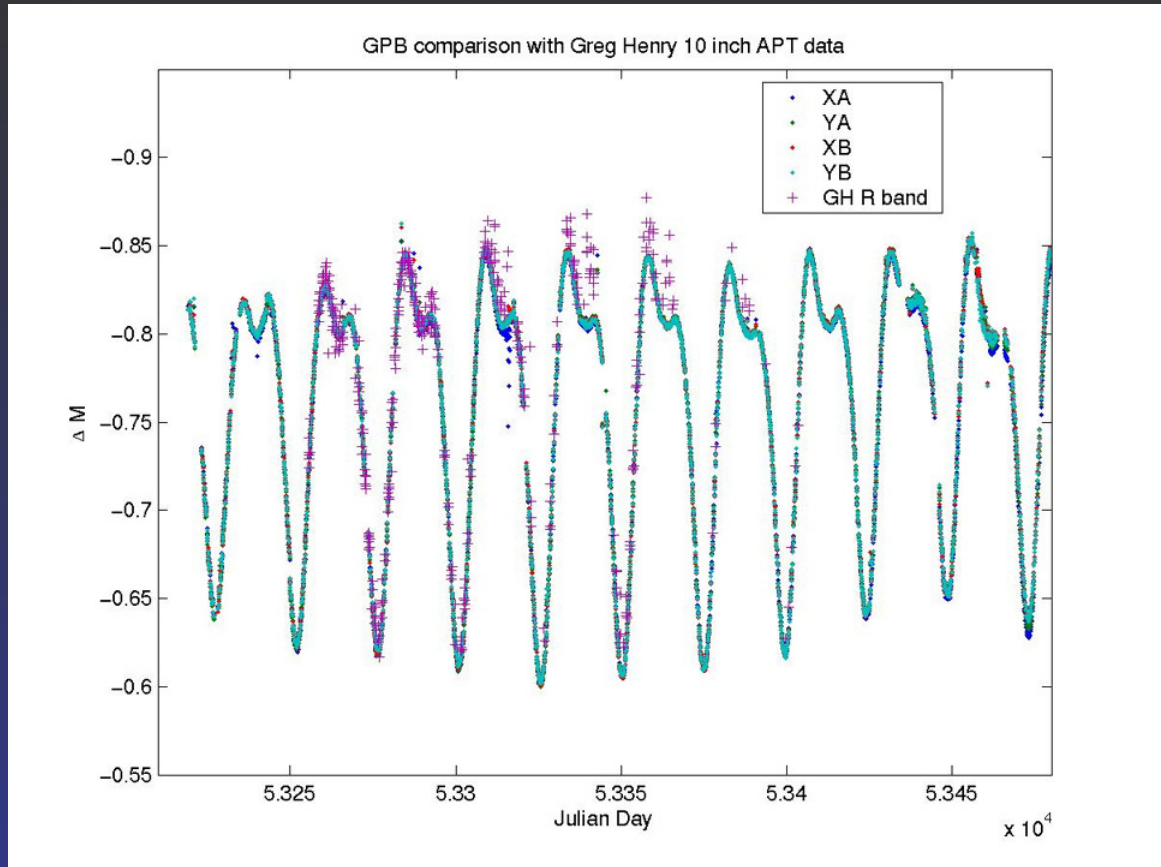
- Optical & radio binary star
- Magnitude - 5.7 (variable)
- Declination - 16.84 deg
- Proper motion measured by SAO using VLBI

Preliminary HR 8703 Positions for Peak of Radio Brightness  
 Solar System Barycentric, J2000 Coordinate System



# Ground-based & Space Observations of IM Peg

John Goebel (NASA Ames)  
 Suwen Wang (Stanford)  
 Michael Ratner (SAO)  
 Greg Henry (U of Tenn.)  
 Jeff Kolodziejczak  
 (NASA Marshall Center)  
 Svetlana Berdyugina  
 (ETH, Switzerland)



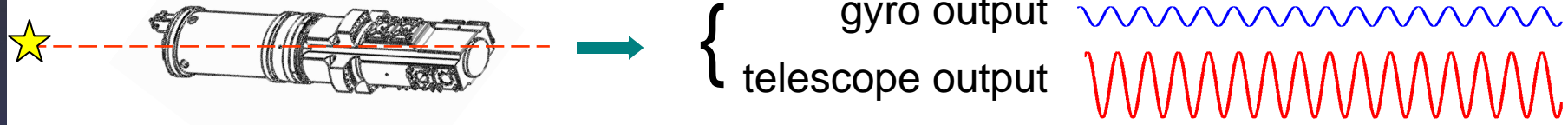
+++ GP-B flight data (*peaked toward red*)  
 ■ G. Henry's ground-based data (*visible light*)


Through GP-B, IM Peg will be most completely characterized star in the entire heavens!



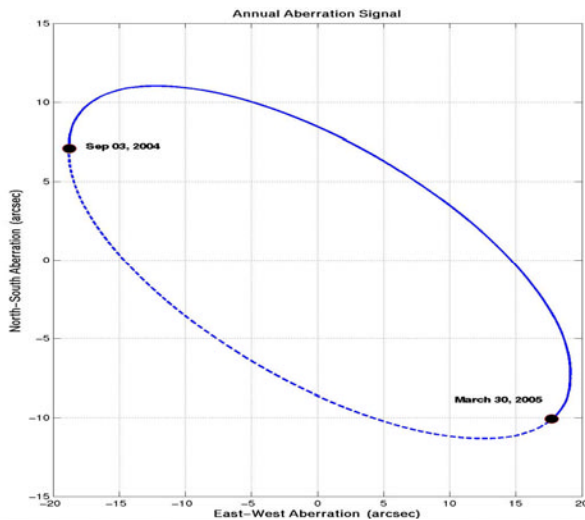
# Dither & Aberration: Two Secrets of GP-B


Dither -- Slow 30 marc-s oscillations injected into pointing system



 *scale factors matched for accurate subtraction*

Aberration (Bradley 1729) -- Nature's calibrating signal for gyro readout



Orbital motion  varying apparent position of star  
 $(v_{\text{orbit}}/c + \text{special relativity correction})$

Earth around Sun -- 20.4958 arc-s @ 1 year period  
 S/V around Earth -- 5.1856 arc-s @ 97.5 min period

 Continuous accurate calibration of GP-B experiment

# 3 Phases of In-flight Verification

## A. Initial orbit checkout (128 days)

- ◆ re-verification of all ground calibrations [scale factors, tempco's etc.]
- ◆ disturbance measurements on gyros at low spin speed

## B. Science Phase (353 days)

- ◆ exploiting the built-in checks [Nature's helpful variations]

## C. Post-experiment tests (46 days)

- ◆ refined calibrations through deliberate enhancement of disturbances, etc. [...learning the lesson from Cavendish]

### *Data Reduction Team:*

Mac Keiser, \* Ed Fei (undeclared), Michael Heifetz, Jie Li, Yoshimi Ohshima (\* A/A), \* Michael Salomon (A/A), David Santiago (\* Physics), Alex Silbergleit, \* Sara Smoot (A/A), Vladimir Solomonik, Karl Stahl (\* ME) + Bill Bencze, Peter Boretsky, Bruce Clarke, Dan DeBra, Barry Muhlfelder, Paul Shestople, John Turneaure, Suwen Wang, Paul Worden

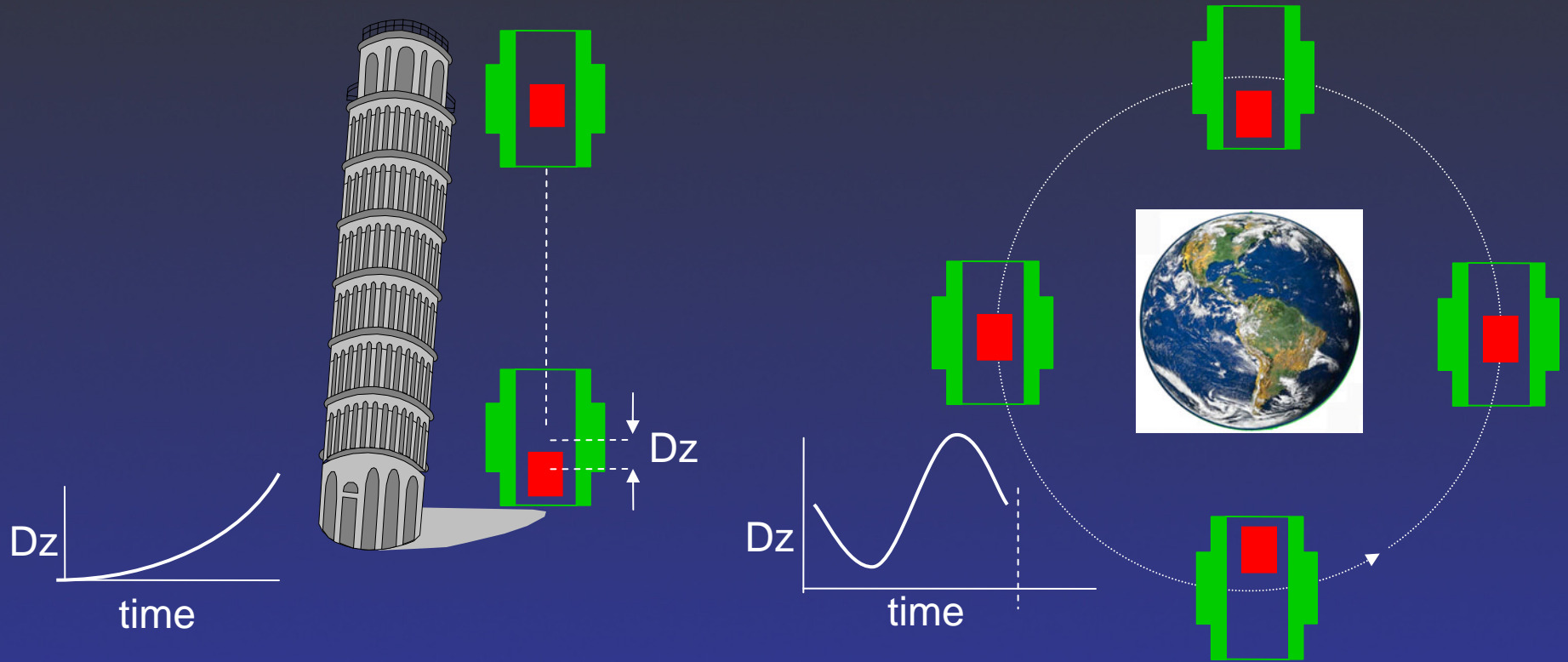
## Wider Significance of the GP-B Experience

- Physics-Aero/Astro collaboration from the start
- 'Near Zero' in space
- Some technologies
  - ◆ *Drag-free*
  - ◆ *Pointing*
  - ◆ *Cryogenics*
- Integrated Science/Operations team
- University/industry/NASA collaboration



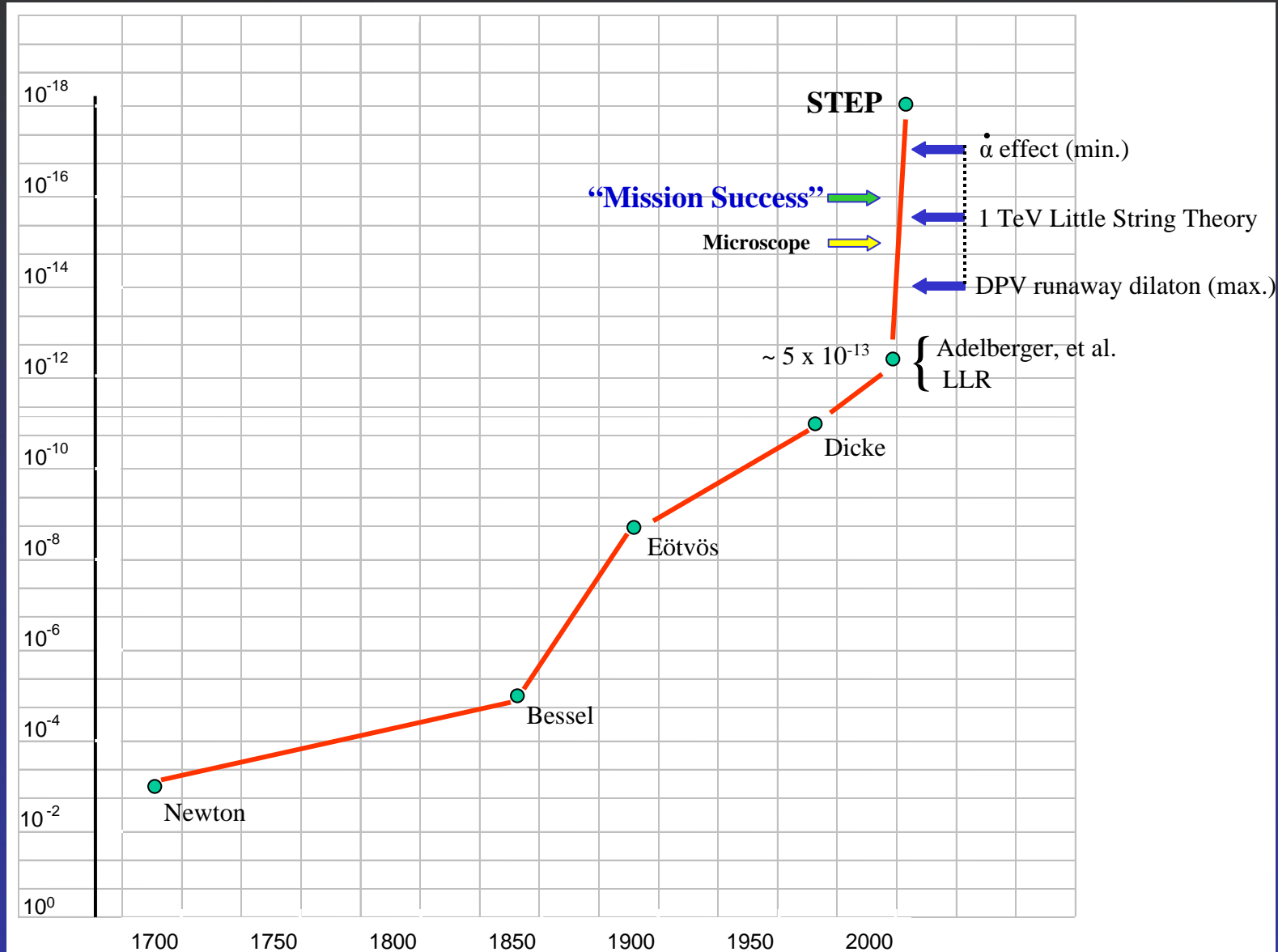
# Satellite Test of the Equivalence Principle

Newton's Mystery {  $F = ma$  mass - the receptacle of inertia  
 $F = GMm/r^2$  mass - the source of gravitation



Orbiting drop tower experiment { \* More time for separation to build  
 \* Periodic signal

# Space > 5 Orders of Magnitude Leap



# STEP International Collaboration

## Research Center Partners

Stanford University

University of Birmingham, UK

ESTEC

FCS Universität, Jena, Germany

Imperial College, London, UK

Institut des Hautes Études Scientifiques, Paris

ONERA, Paris, France

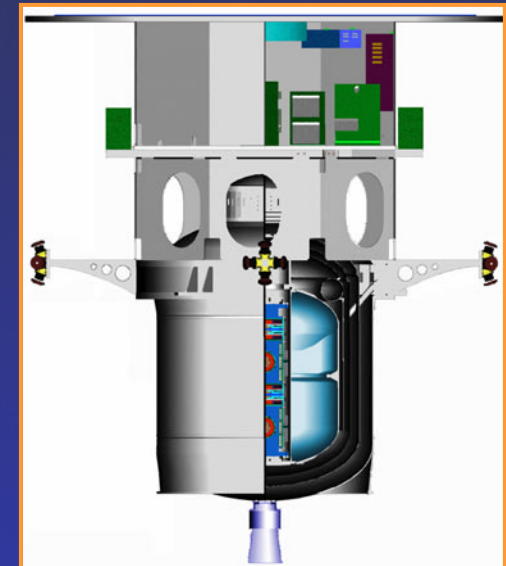
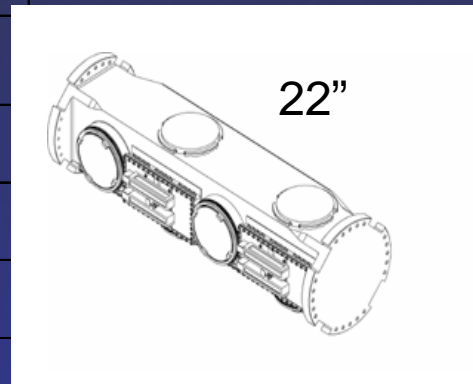
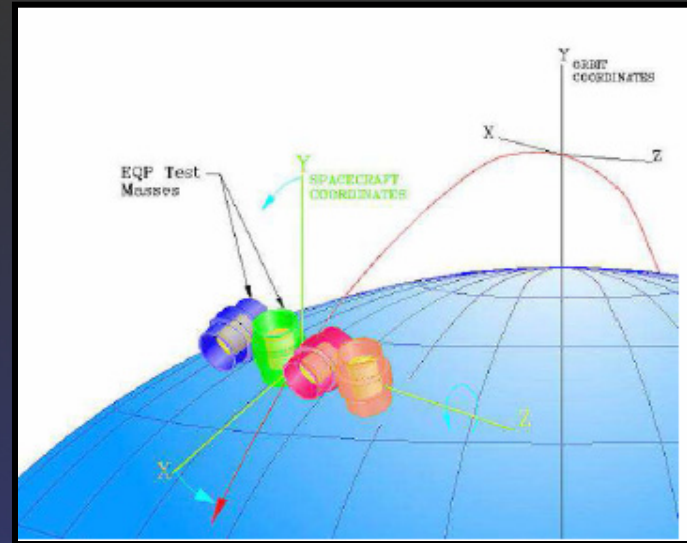
PTB, Braunschweig, Germany

Rutherford Appleton Laboratory, UK

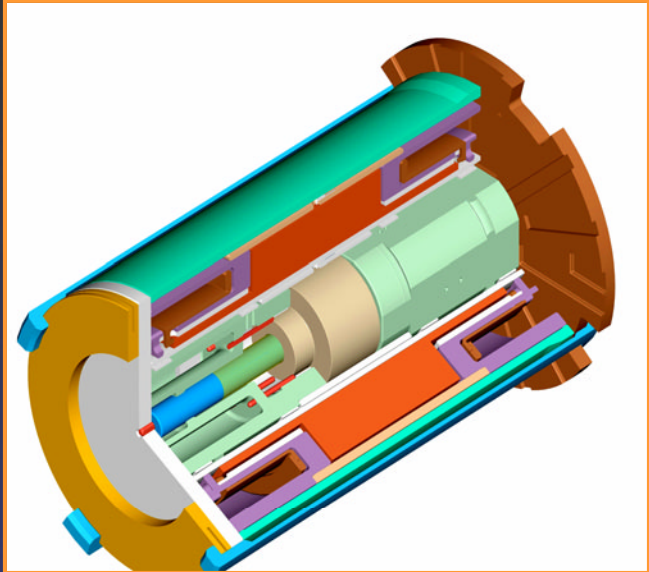
University of Strathclyde, UK

Università di Trento, Italy

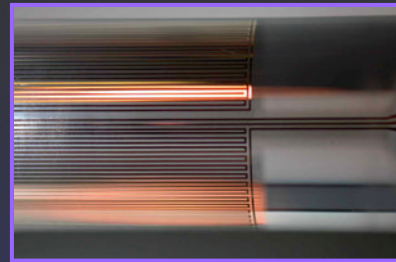
ZARM, Universität Bremen, Germany



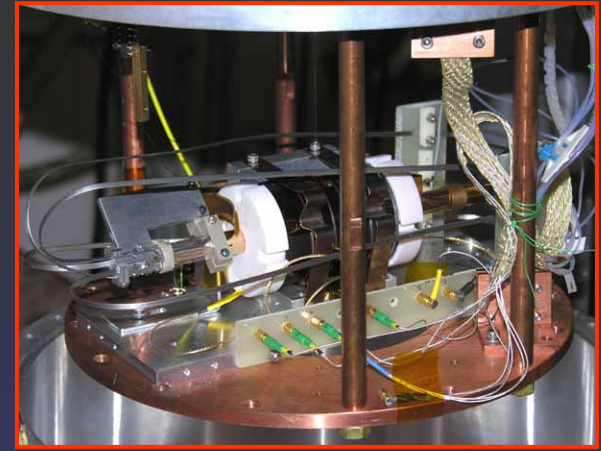
# Some Elements of the STEP Mission



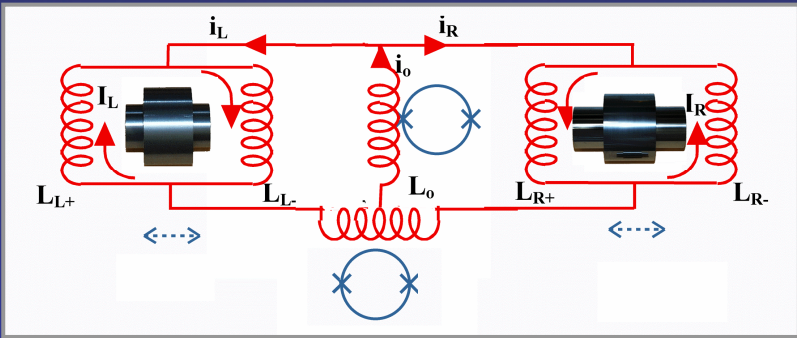
assembled flight instrument



magnetic bearing



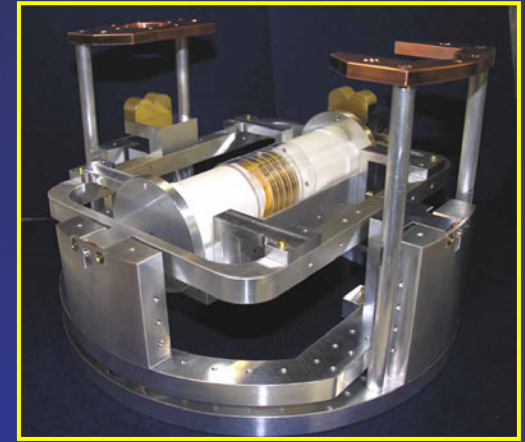
bearing under test



differential SQUID readout

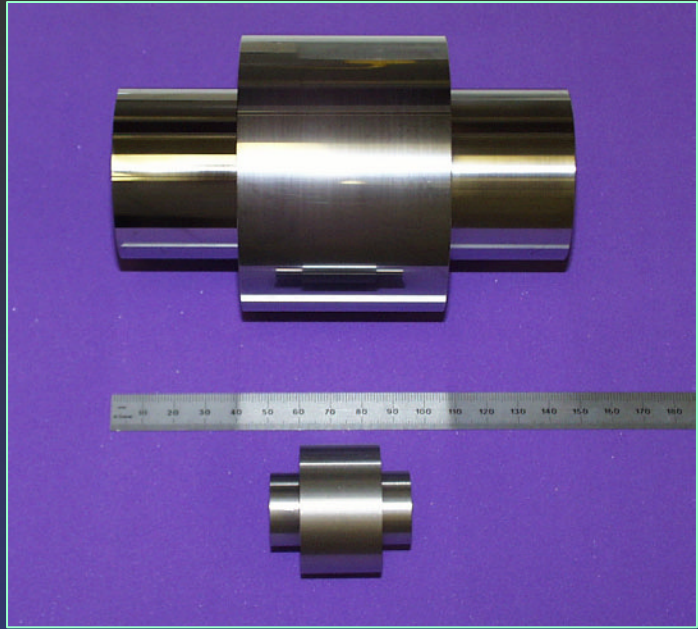


SQUID assembly



2-axis tilt platform

# Test Mass Shape & Composition



Dimensions give 6<sup>th</sup> order insensitivity to gravity gradient disturbances from spacecraft --  $\mu\text{m}$  tolerances

## Test masses as 'different' as possible

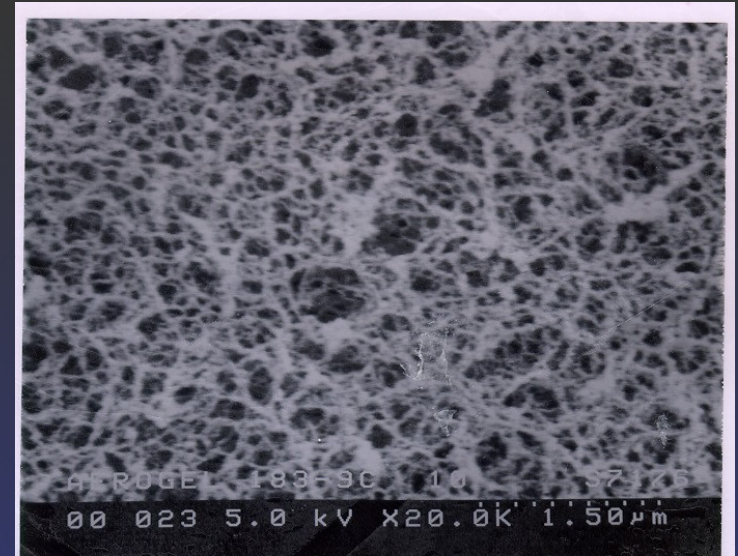
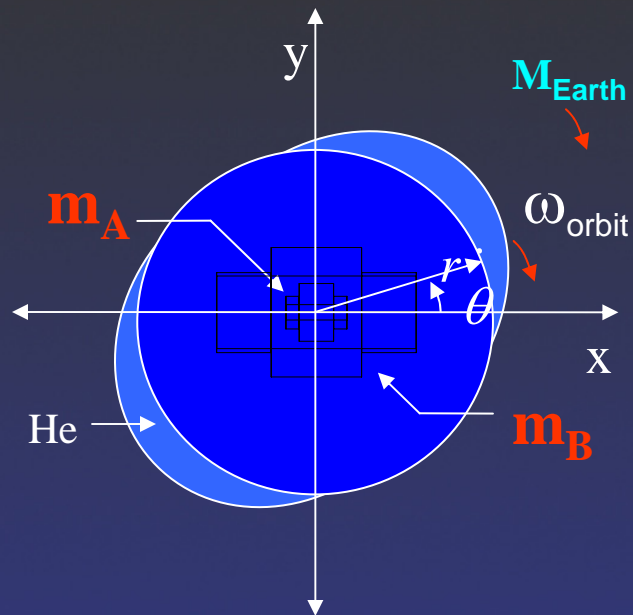
Material	Z	N	$\left(\frac{N + Z}{\mu} - 1\right)10^3$ Baryon Number	$\frac{N - Z}{\mu}$ Lepton Number	$\frac{Z(Z - 1)}{\mu(N + Z)^3}$ Coulomb Parameter
Be	4	5	-1.3518	0.11096	0.64013
Si	14	14.1	0.8257	0.00387	2.1313
Nb	41	52	1.0075	0.11840	3.8462
Pt	78	117.116	0.18295	0.20051	5.3081

Damour C&QG 13 A33 (1996)

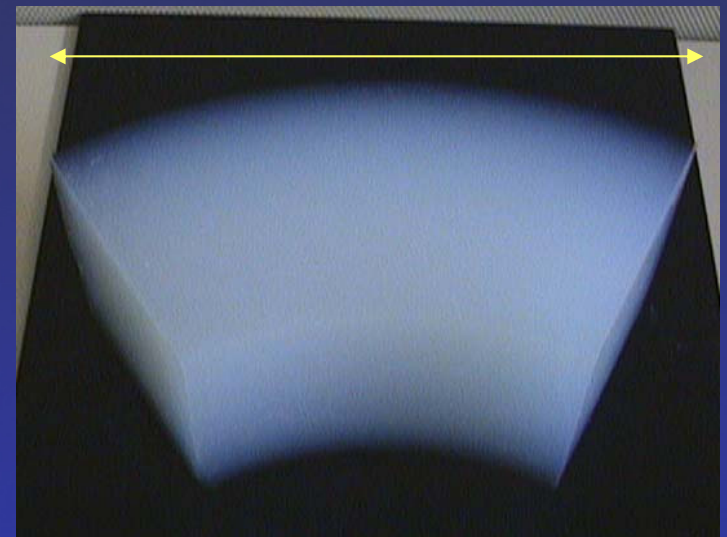


# Helium Tide Control

## Silica Aerogel Constraint



250 mm



- void sizes 100 to 1000 nm
- confines He even in 1g
- passed cryogenic shake test

# STEP: a Landmark in Fundamental Physics



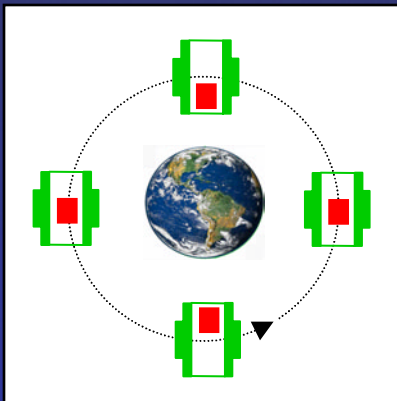
What Physics has done for Space?

“I guess one of the reasons we got here today was because of a gentleman named Galileo....who made a rather significant discovery about falling objects in gravity fields.”

--David Scott



What Space can do for Physics?



STEP (and probably only STEP) has the potential of discovering new forces in Nature that would tell us a lot about why the universe is as it is, and what its ultimate state will be.”

--Thibault Damour

