

# KEY FACTS ABOUT GRAVITY

- THE WEAKEST OF THE FOUR BASIC FORCES.



$$\frac{F_{\text{GRAV}}}{F_{\text{EEM}}} = \frac{GM_p^2/r^2}{e^2/r^2} = \frac{GM_p^2}{e^2} \sim 10^{-40}$$

- UNIVERSAL - COUPLES TO ALL MASS ENERGY.

- LONG-RANGE  $F = \frac{GM_1M_2}{r^2}$

- UNSCREENED - NO NEGATIVE GRAVITATIONAL "CHARGE".

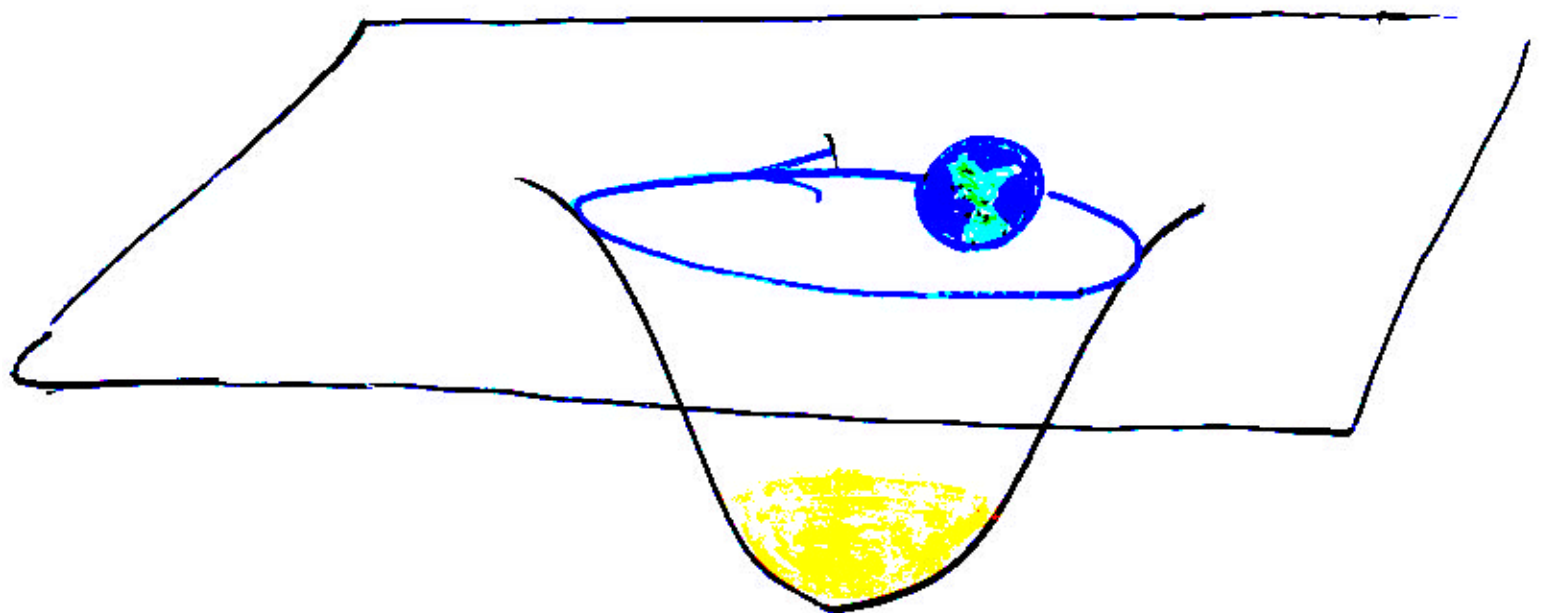


GRAVITY GOVERNS THE STRUCTURE OF MATTER ON THE LARGEST SCALES.

# RELATIVISTIC GRAVITY

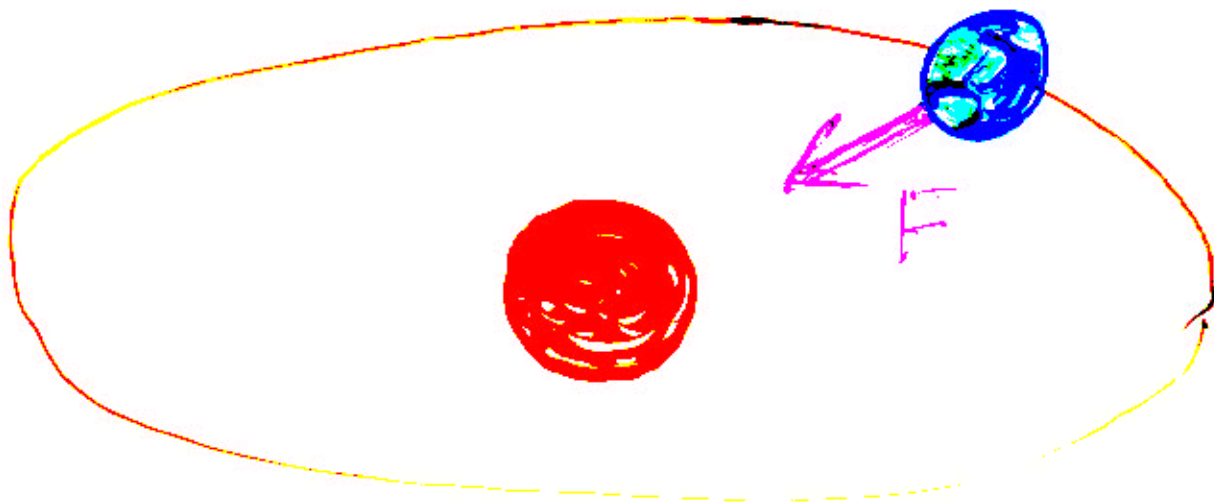
## EINSTEIN'S GENERAL RELATIVITY

- GRAVITY IS GEOMETRY.
- MASS CURVES SPACETIME
- FREE MASS MOVES ON THE STRAIGHTEST PATHS IN CURVED SPACETIME.



## Newtonian Gravity (1687)

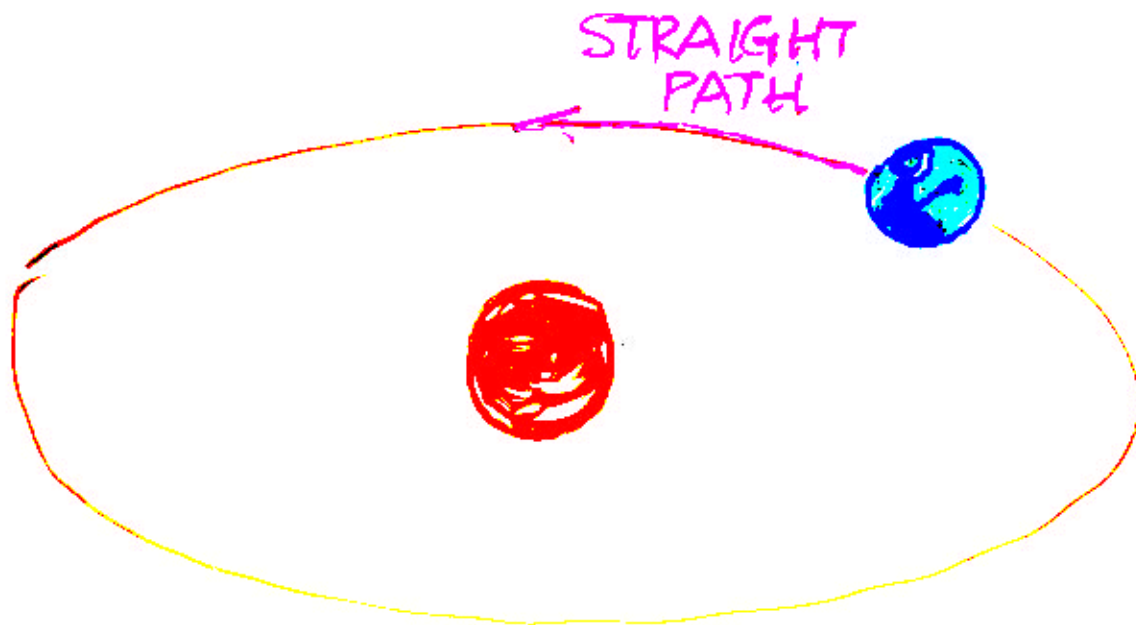
THE EARTH  
TRAVELS  
AROUND  
THE SUN  
BECAUSE  
IT IS PULLED  
BY THE GRAVITATIONAL FORCE  
EXERTED BY THE MASS OF THE SUN.



# Einstein Gravity (General Relativity, 1915)



THE EARTH  
TRAVELS  
AROUND THE  
SUN  
BECAUSE  
ITS FOLLOWING  
THE STRAIGHTEST PATH  
IN THE CURVED SPACETIME  
PRODUCED BY THE SUN'S MASS





# FRONTIER OF THE SMALLEST SCALES

## PLANCK SCALE

$$l = \left( \frac{G\hbar}{c^3} \right)^{1/2} \sim 10^{-33} \text{ cm}$$

## PLANCK ENERGY

$$E_p = \frac{\hbar c}{l} = \left( \frac{\hbar c^5}{G} \right)^{1/2} \sim 10^{19} \text{ GeV}$$

THIS IS THE SCALE CHARACTERIZING:

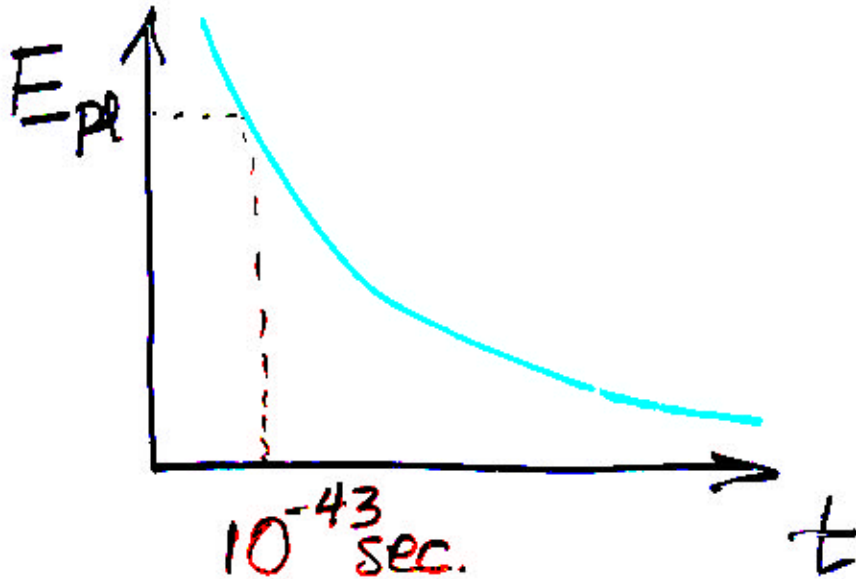
- UNIFIED THEORIES OF THE FOUR BASIC FORCES (E.G. STRING THEORY)
- THE UNION OF GRAVITY AND QUANTUM MECHANICS. (E.G. QUANTUM THEORY OF GEOMETRY)



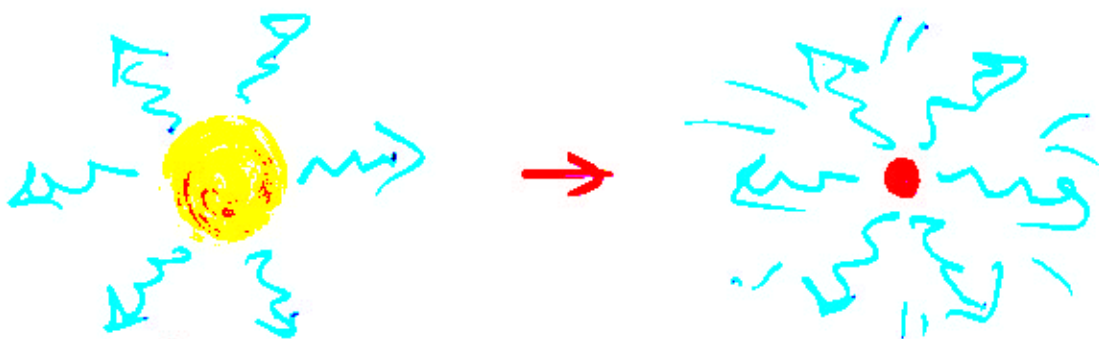
WHERE IN THE UNIVERSE ARE THE HIGHEST ENERGIES REALIZED?

$$E_{Pl} = \left( \frac{\hbar c^5}{G} \right)^{1/2} \sim 10^{19} \text{ GEV.}$$

● THE BIG BANG:



● EXPLODING BLACK HOLES.



HAWKING EFFECT.

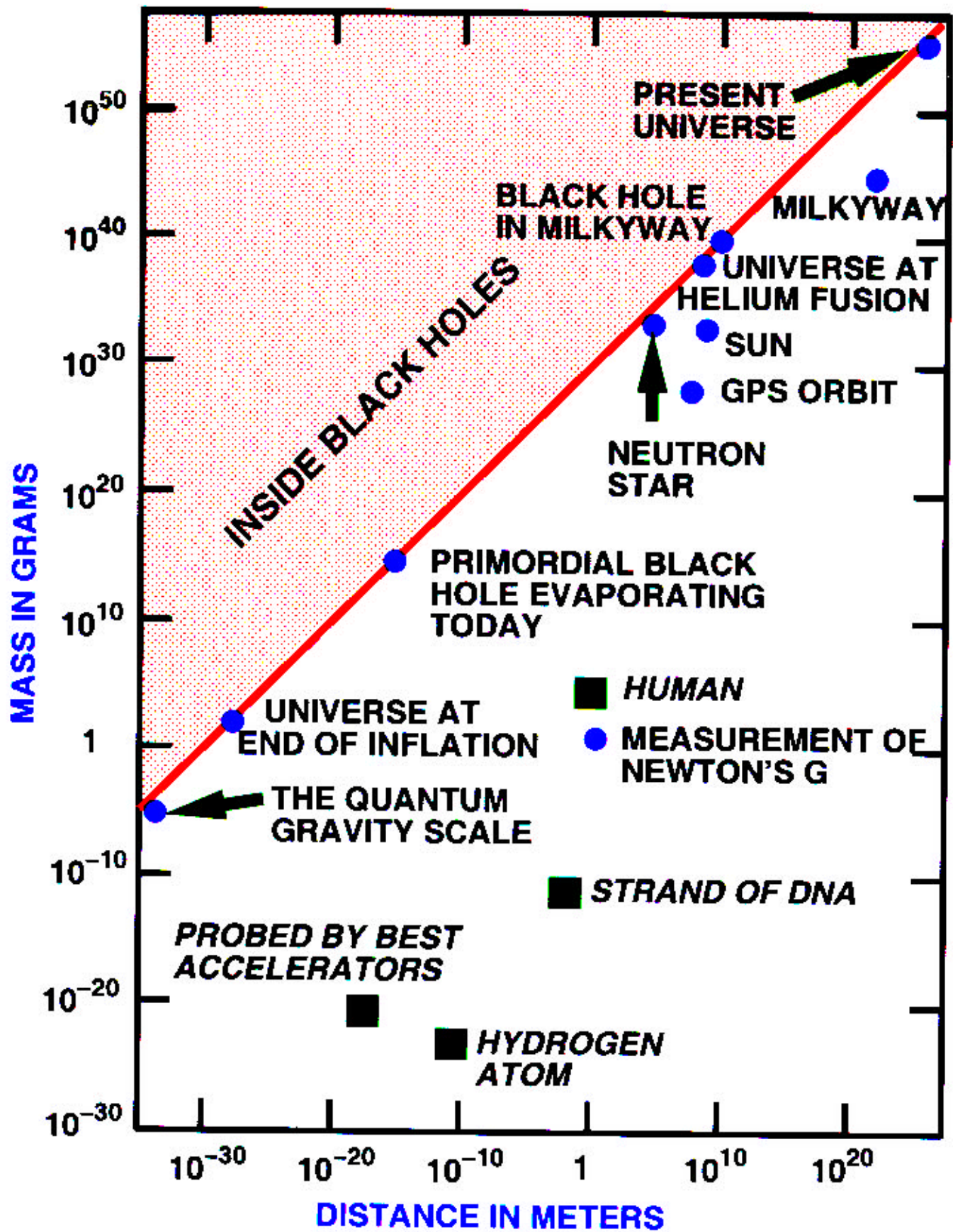
# WHEN IS RELATIVISTIC GRAVITY IMPORTANT?

$$\frac{GM}{c^2 R} \sim 1$$

- SUN  $GM/Rc^2 \sim 10^{-6}$
- NEUTRON STAR  $GM/Rc^2 \sim 0.1$
- BLACK HOLE  $GM/Rc^2 = \frac{1}{2}$  (MAXIMUM VALUE)
- UNIVERSE  $GM/Rc^2 \sim \frac{1}{2}$



# Gravitational Physics: From Quantum to Cosmos



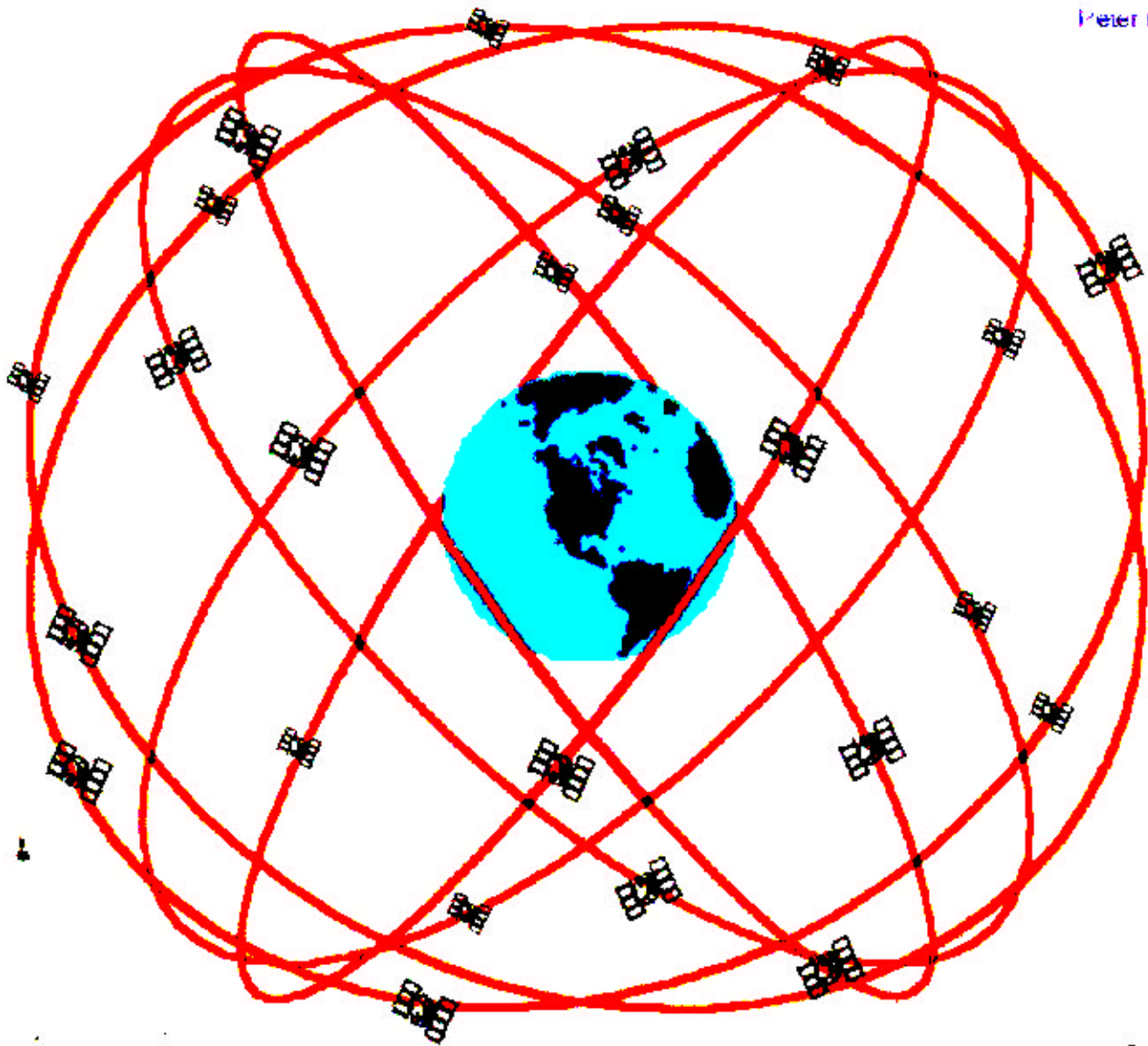
- THE LARGE SCALES OF ASTROPHYSICS AND COSMOLOGY ARE INCREASINGLY ACCESSIBLE TO OBSERVATION.
- THE SMALLEST SCALES ARE INCREASINGLY ACCESSIBLE TO THEORETICAL SPECULATION.

## RESULT:

GENERAL RELATIVITY IS INCREASINGLY INTEGRATED WITH NEIGHBORING AREAS.

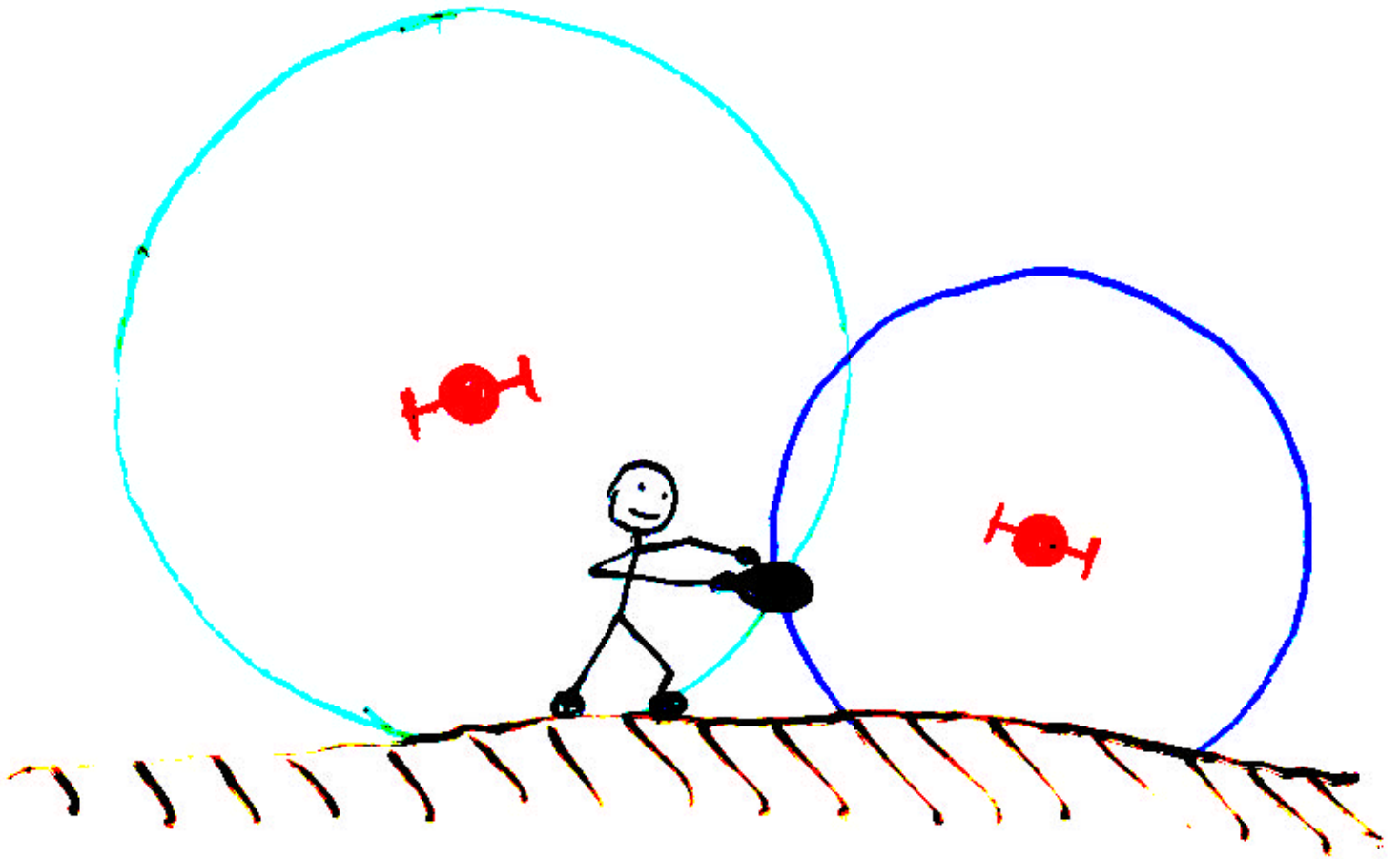
GENERAL RELATIVITY IS INCREASINGLY IN CONTACT WITH EXPERIMENT & OBSERVATION.





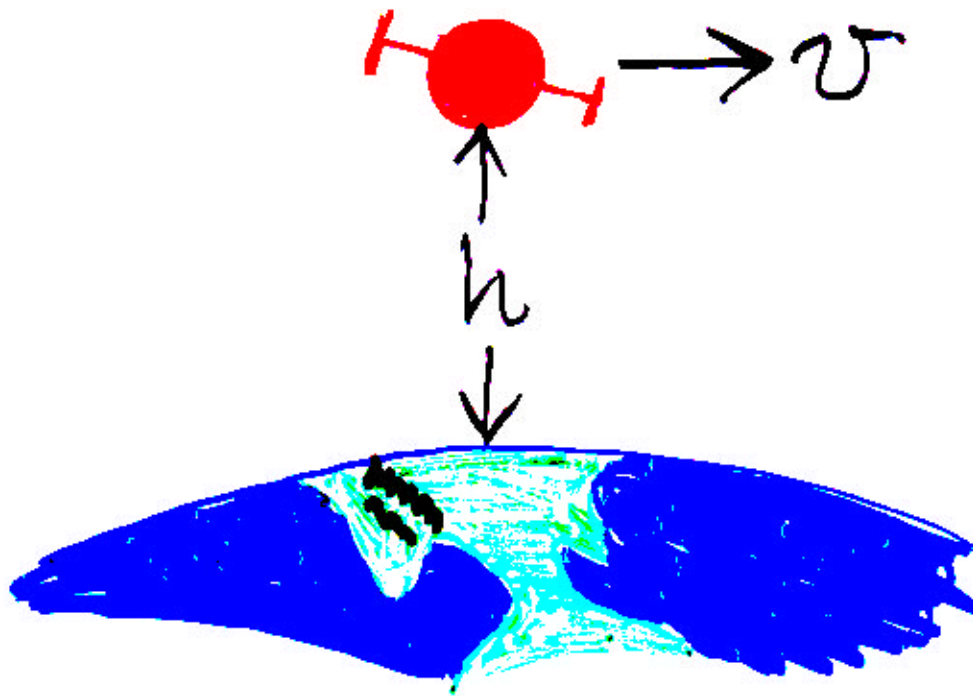
**GPS Nominal Constellation**  
**24 Satellites in 6 Orbital Planes**  
**4 Satellites in each Plane**  
**20,200 km Altitudes, 55 Degree Inclination**

GPS





# RELATIVISTIC EFFECTS IN GPS



- SPECIAL RELATIVITY  
MOVING CLOCKS RUN SLOW

$$\sqrt{1 - v^2/c^2} \rightarrow 1 - 0.8 \times 10^{-10}$$

- GENERAL RELATIVITY  
CLOCKS HIGHER IN A GRAVITATIONAL POTENTIAL RUN FAST.

$$1 + \frac{gh}{c^2} \rightarrow 1 + 5.2 \times 10^{-10}$$

# CLASSICAL SOLAR SYSTEM TESTS

- GRAVITATIONAL REDSHIFT,
- BENDING AND TIME DELAY OF LIGHT,
- PRECESSION OF PERIHELION.

SMALL EFFECT  
DIFFICULT TO DETECT

PRECISION MEASUREMENT  $\lesssim 1\%$

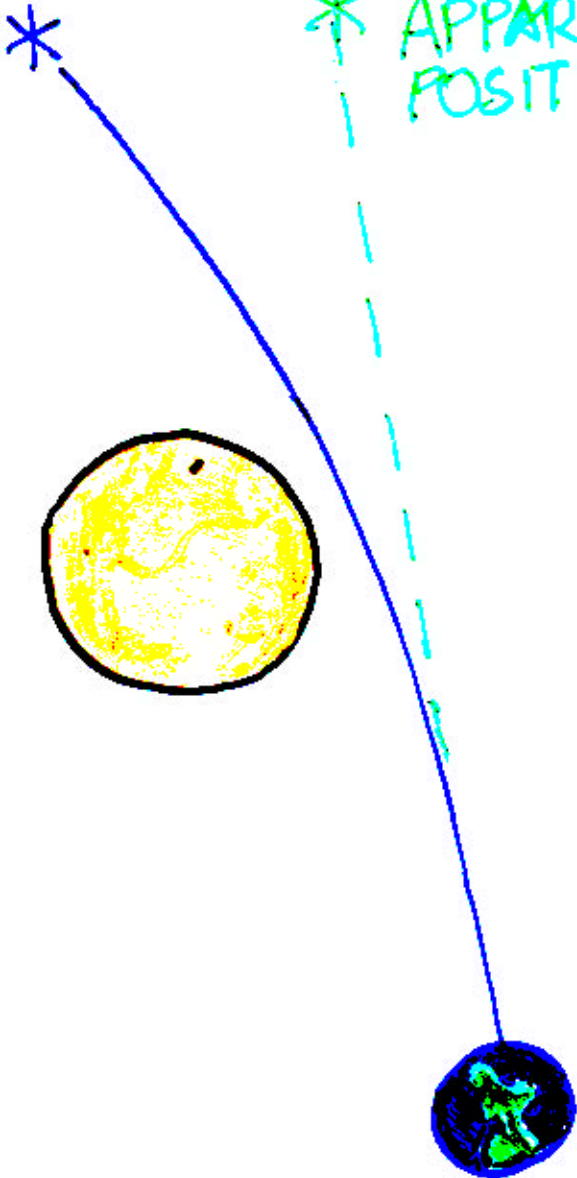
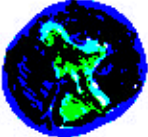
ANNOYING CORRECTION!  
IN OTHER PRECISION MEASUREMENTS

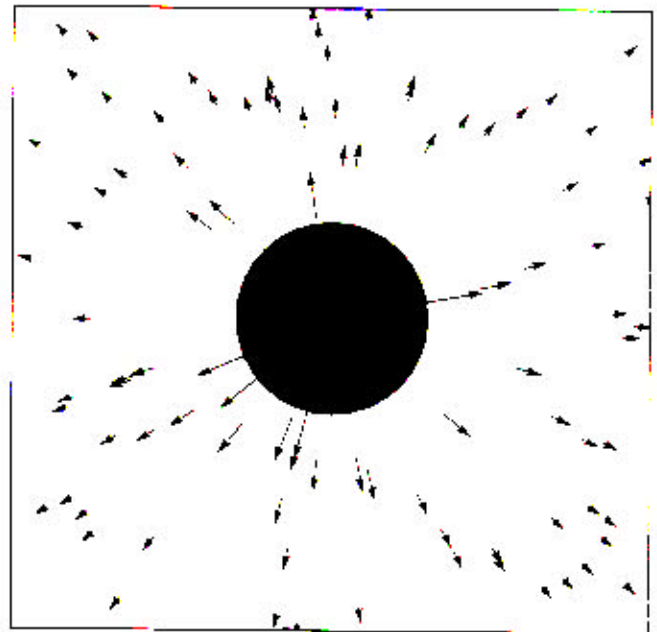
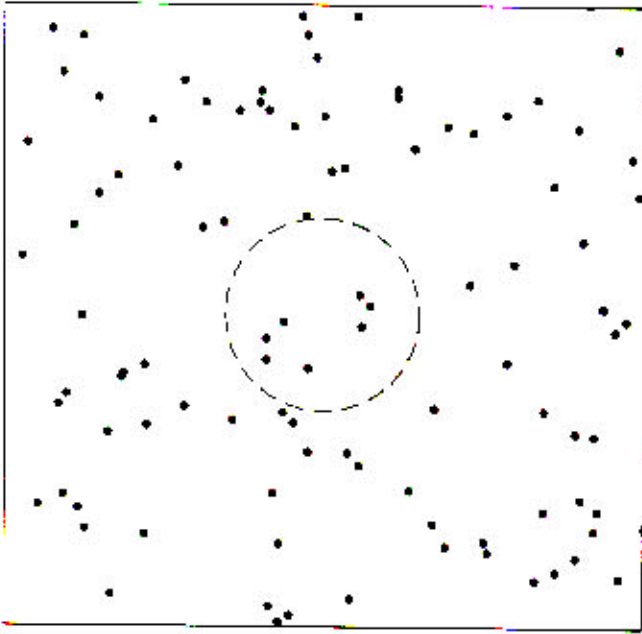
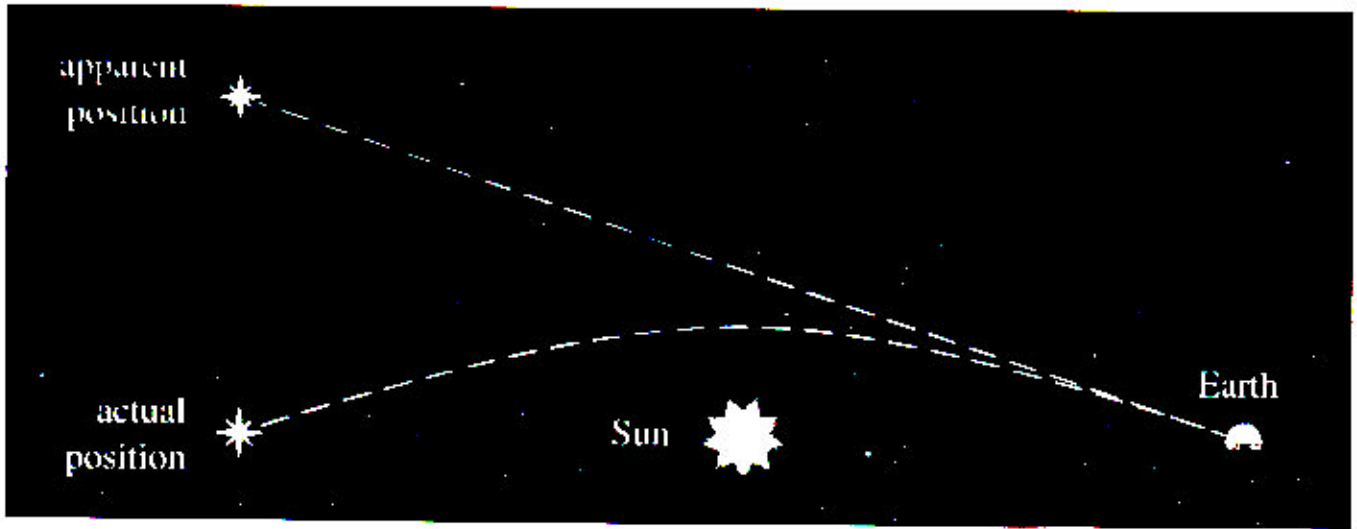
NEW TOOL

# BENDING OF LIGHT BY THE SUN

ACTUAL POSITION \*

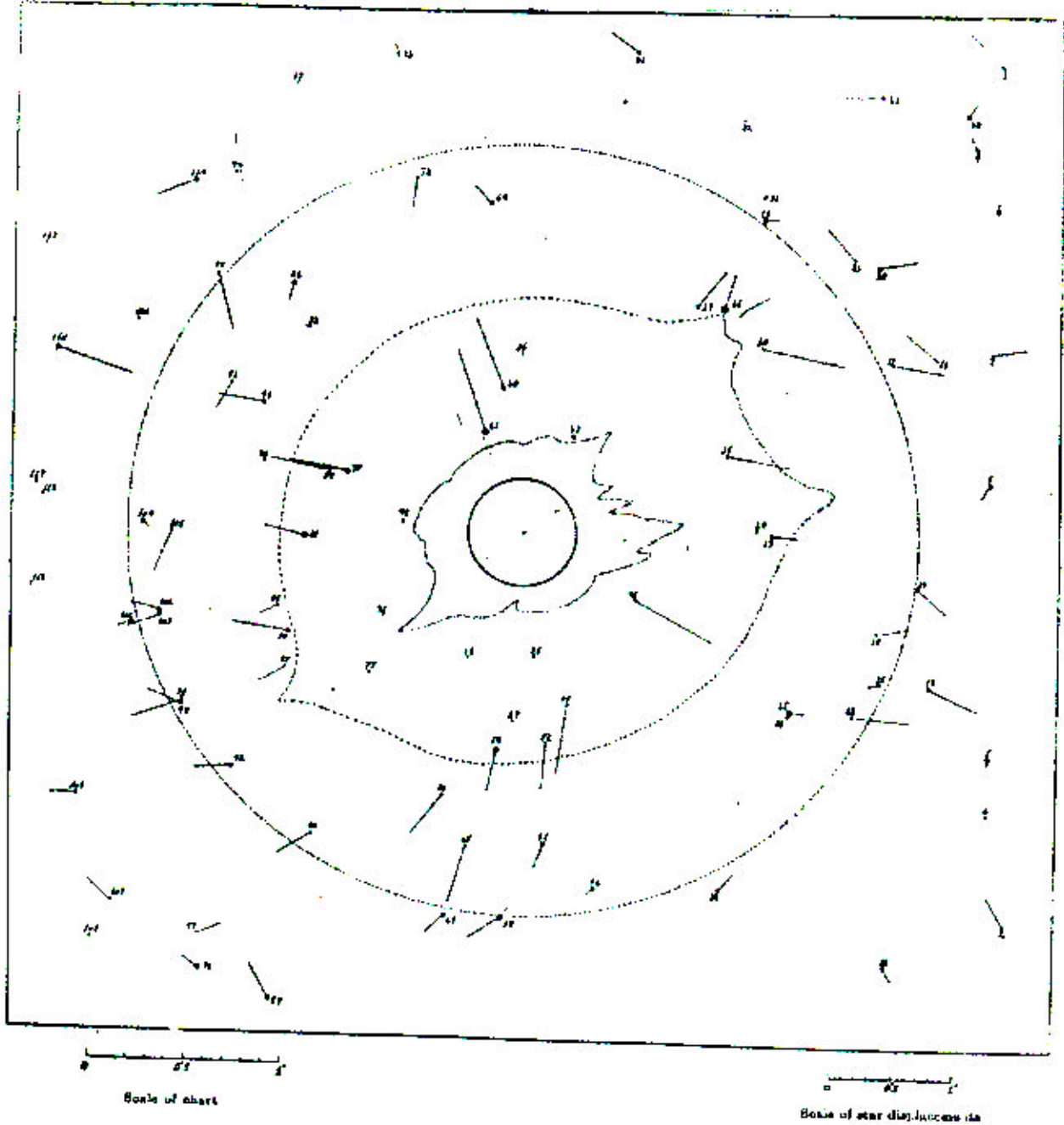
\* APPARENT POSITION







# Solar Eclipse 1922

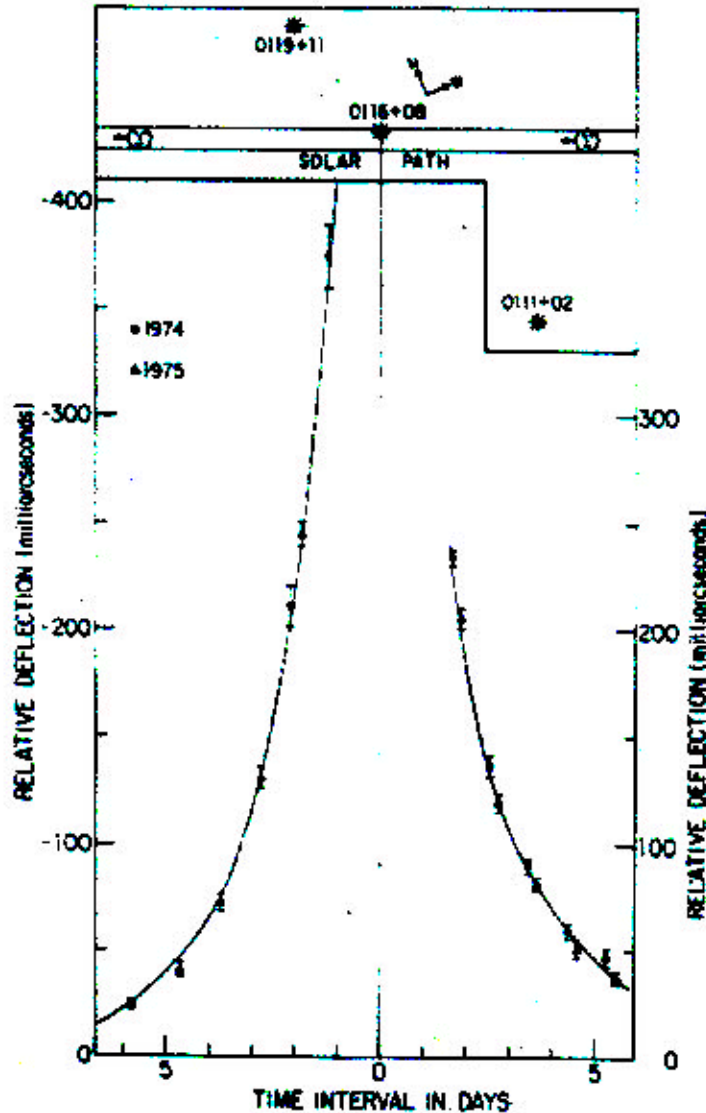
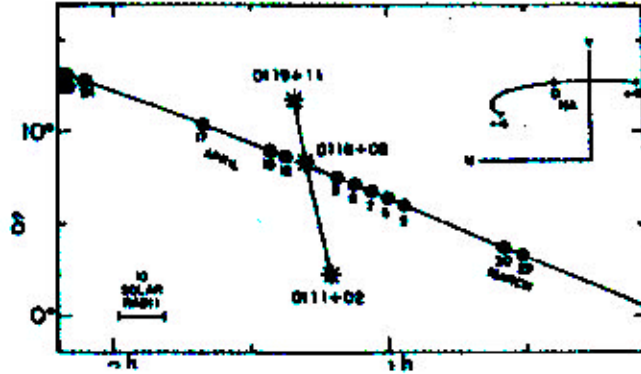


Campbell & Trumpler

$$\delta = 1.007 \pm .009$$

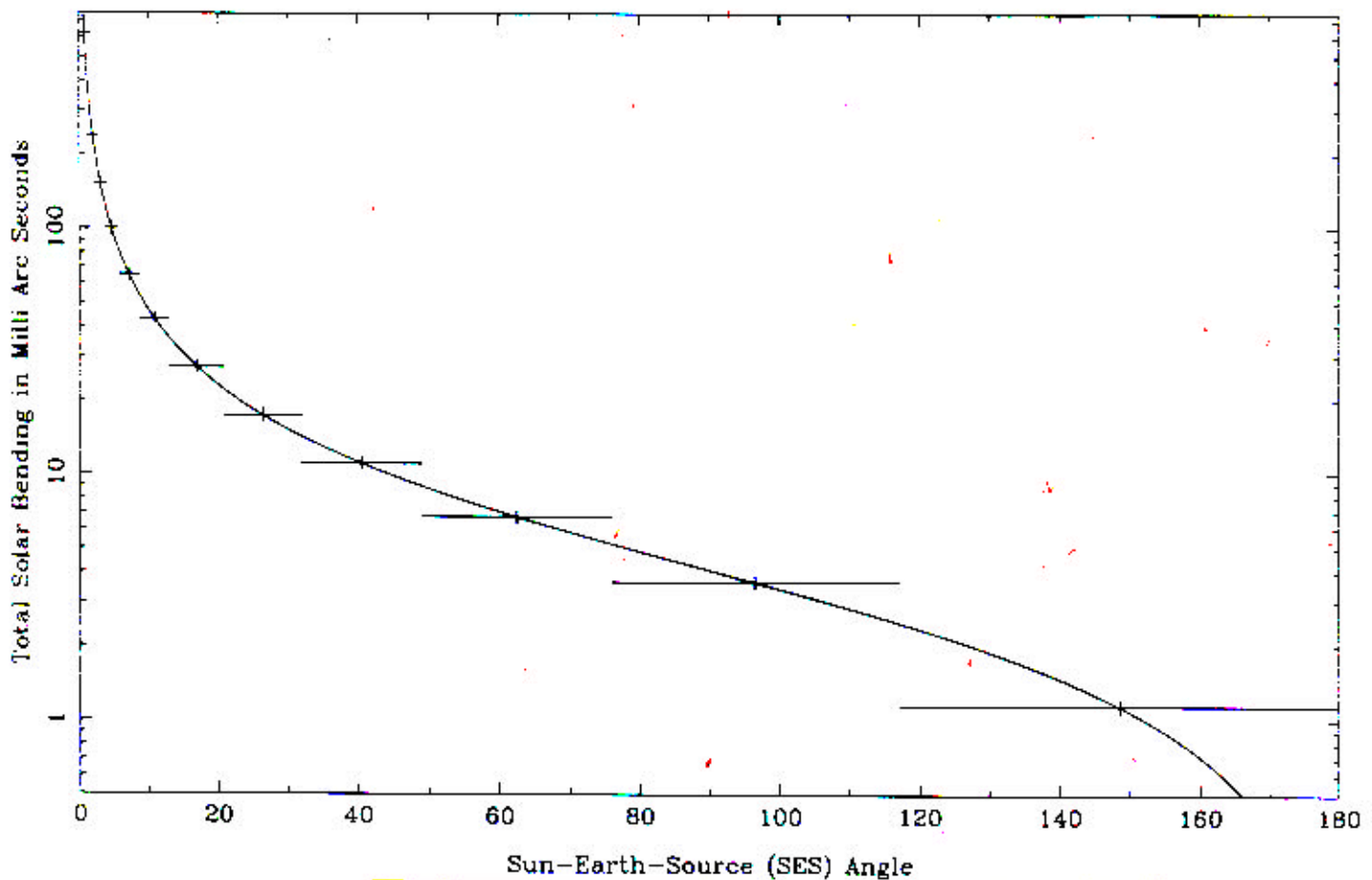
1975

LBI Bending of Light by the Sun



Fomalont & Sramek

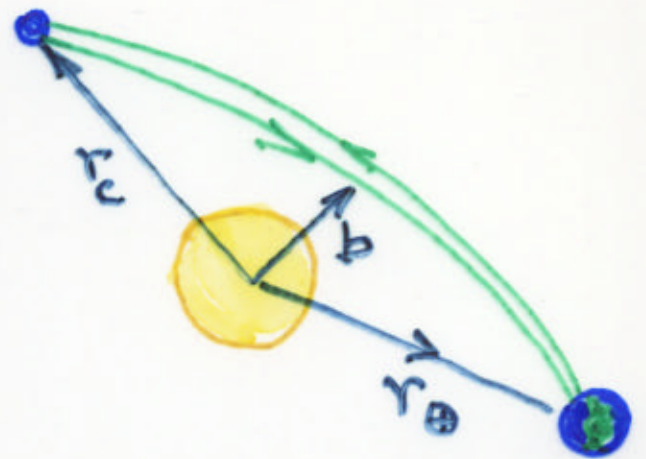
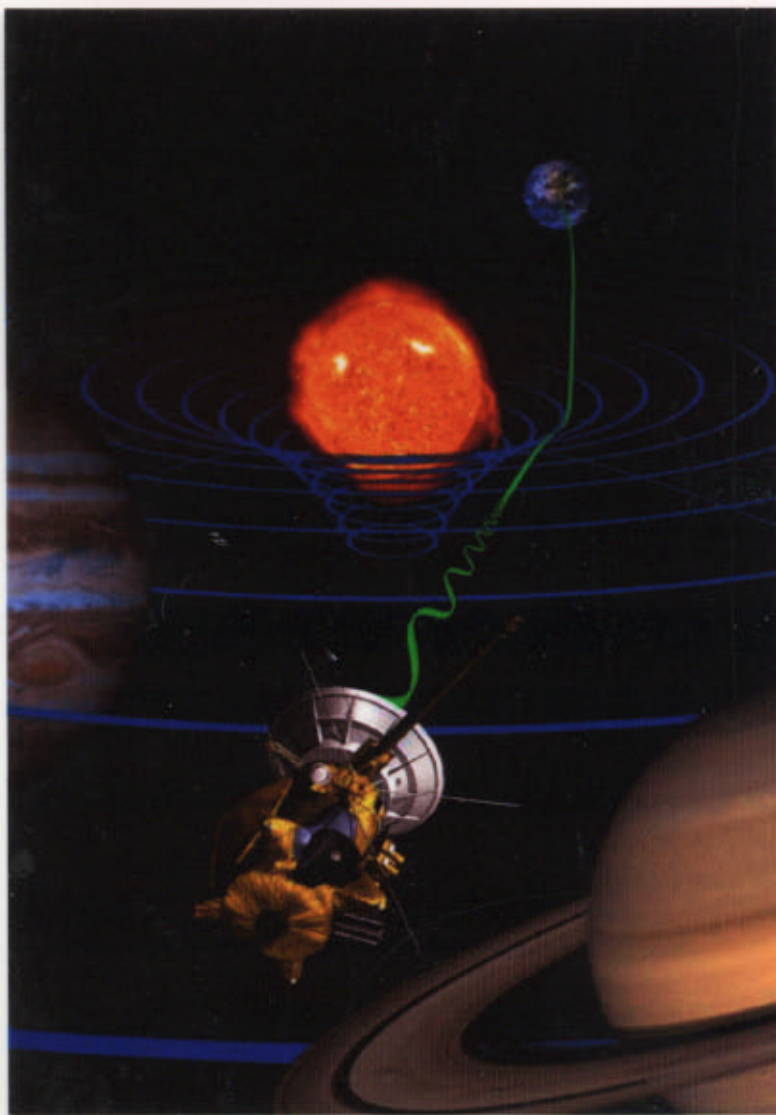
## VLBI Measurements of Bending of Light



$$\gamma = 0.99983 \pm 0.00045$$

S. Shapiro, et al (2004)  
M. Eubanks, et al. (unpub)

# CASSINI DETERMINATION OF $\gamma'$



MEASURE FREQUENCY  
SHIFT OF TWO-WAY  
RADIO SIGNAL

$$y \equiv \frac{\nu_{\text{REC'D}} - \nu_{\text{TRANS}}}{\nu_{\text{TRANS}}}$$

$$y_{\text{gr}} \approx 4(1 + \gamma') \frac{GM_0}{bc^3} \frac{db}{dt}$$

THREE DIFFERENT UP-DOWNLINK PAIRS  
(XX, KK, XK) ALLOW FOR CORRECTION  
BY SOLAR CORONA

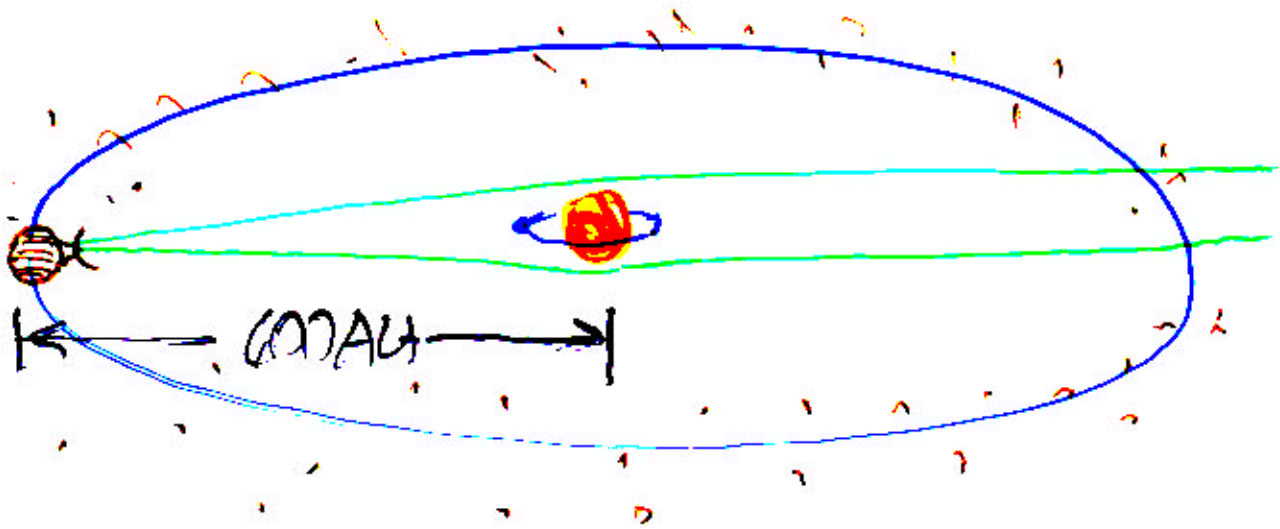
$$\gamma' = 1 \pm (2.1 \pm 2.3) \times 10^{-5}$$





COLEY, TYSON, & TURNER (1996)

# USING THE SUN AS A LENS



OBSERVATORY IN THE KUIPER BELT,

# SOME NEXT STEPS IN PRECISION TESTS

- GRAVITOMAGNETIC  
EFFECTS ( $\sim 1/c^3$ )

GP-B (April 20, 2004)  
Sept 29, 2005)

- NEW TESTS OF THE  
PRINCIPLE OF EQUIVALENCE

STEP ( $\sim ?$ )



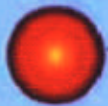




## Geodetic Effect

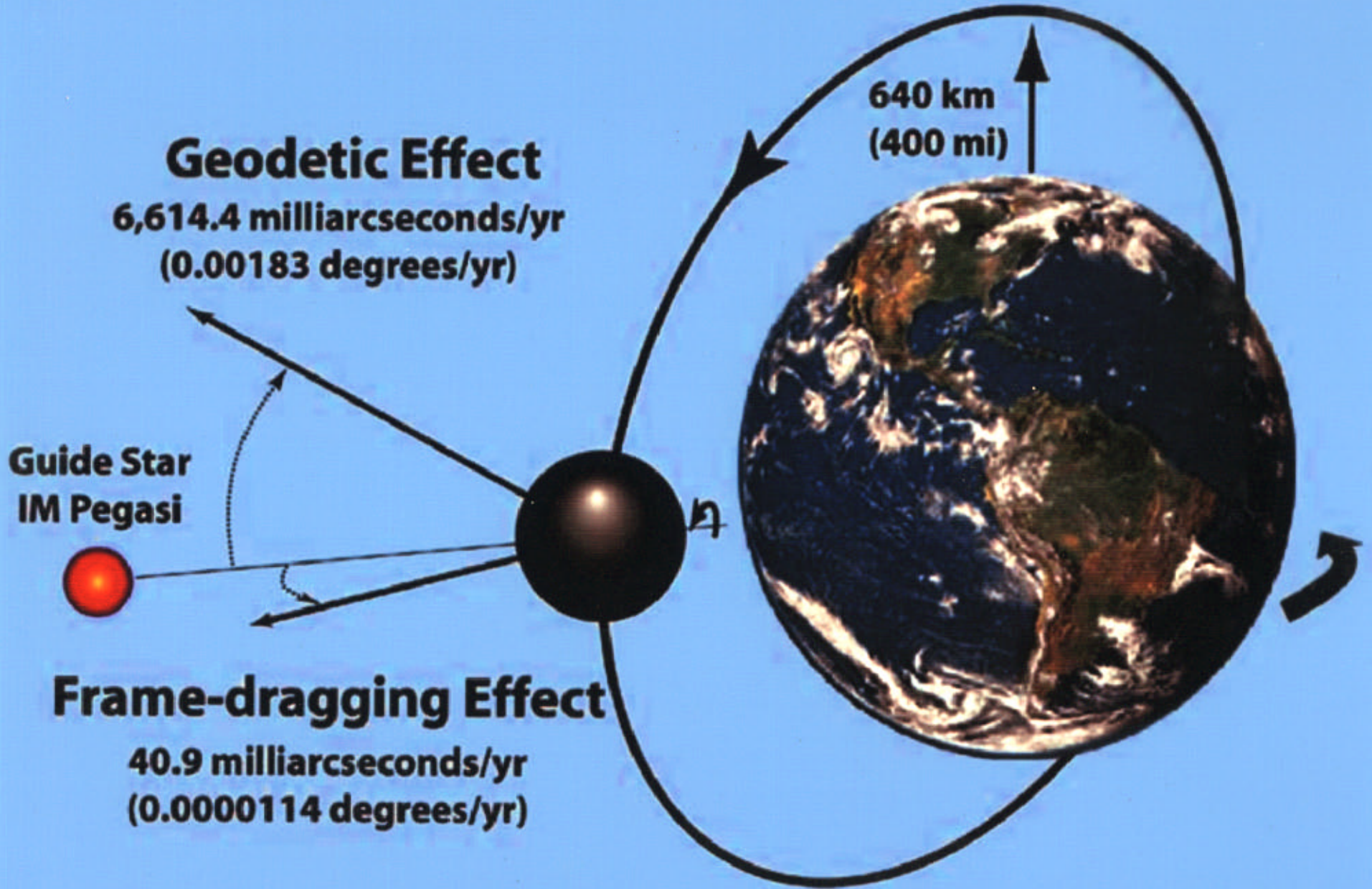
6,614.4 milliarcseconds/yr  
(0.00183 degrees/yr)

Guide Star  
IM Pegasi

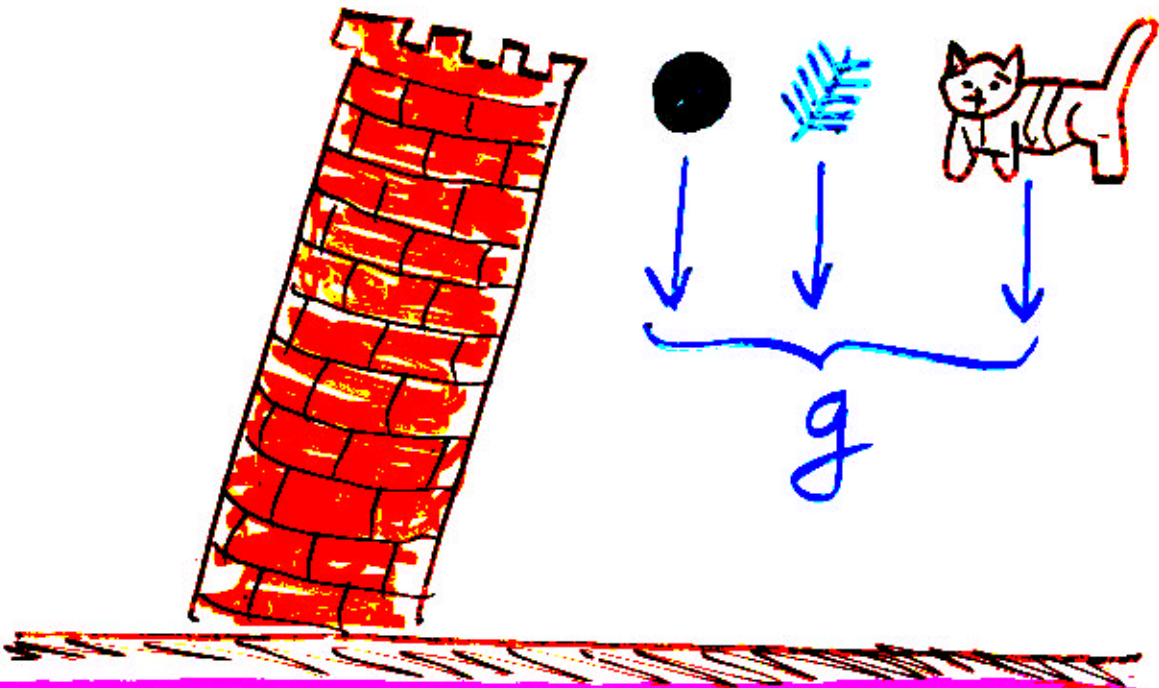


## Frame-dragging Effect

40.9 milliarcseconds/yr  
(0.0000114 degrees/yr)



# PRINCIPLE OF EQUIVALENCE



$g$ 's EQUAL TO  $1.5 \times 10^{-13}$

- CENTRAL TO A GEOMETRIC THEORY OF GRAVITY.
- VIOLATIONS WOULD SIGNAL
- BREAKDOWN OF OUR NOTIONS OF SPACETIME
- NEW FORCES



# Lunar Laser Ranging



- L-1 Lunachok 1
- L-2 Lunachok 2
- A-11 Apollo 11
- A-14 Apollo 14
- A-15 Apollo 15



POSITION OF THE MOON TO ~few CM.

# Los Angeles Times

CO. PAGES

## COLUMN ONE

# Time, Space Obsolete in New View of Universe

■ Many physicists are embracing a revolutionary, still mysterious idea called string theory. The concept rejects several familiar notions and includes the existence of 11 dimensions.

Now, some physicists are taking this revolutionary line of thinking one step further: If their theories are right, in the words of Edward Witten of the Institute for Ad-

## OF SPACE, TIME AND STRINGS

*Rocking the foundations of physics*

■ First in a series

vanced Study in Princeton, space and time may be "doomed."

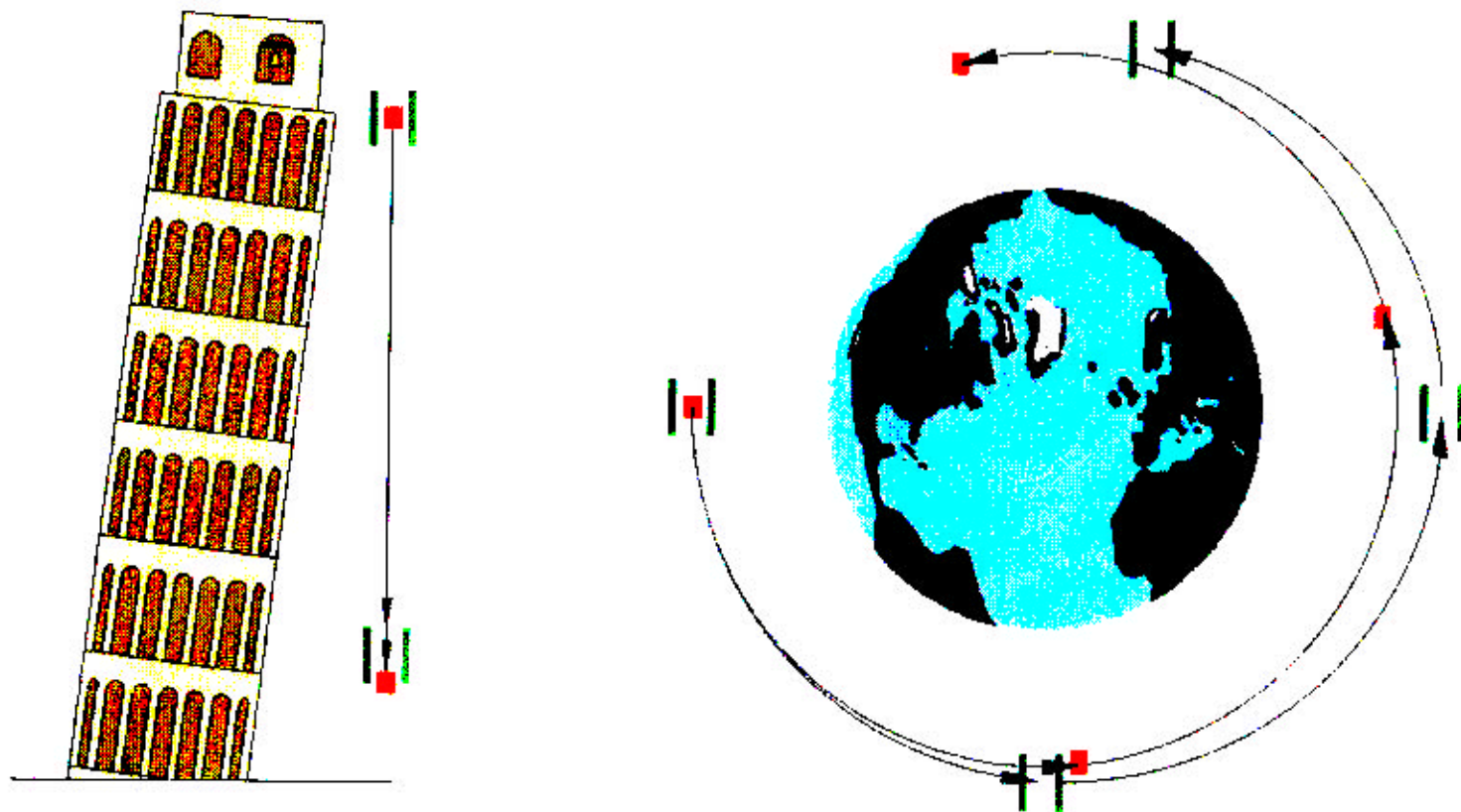
Concurs physicist Nathan Seiberg, also of the institute: "I am almost certain that space and time are illusions. These are primitive notions that will be replaced by something more sophisticated."

SPACE & TIME  
DOOMED!  
WITTEN

SPACE & TIME  
ARE ILLUSIONS!  
SEIBERG

# THEORY: Bizarre Concept Could Explain Universe





$$\Delta g/g \sim 10^{-18}$$

MAJOR THEME  
OF THE THIS DECADE:

THE EXPLORATION OF  
STRONG  
GRAVITATIONAL FIELDS

# STRONG GRAVITY

- GRAVITATIONAL WAVES
- BLACK HOLES
- COSMOLOGY

# GRAVITATIONAL WAVES

EINSTEIN'S THEORY PREDICTS

PROPAGATING RIPPLES

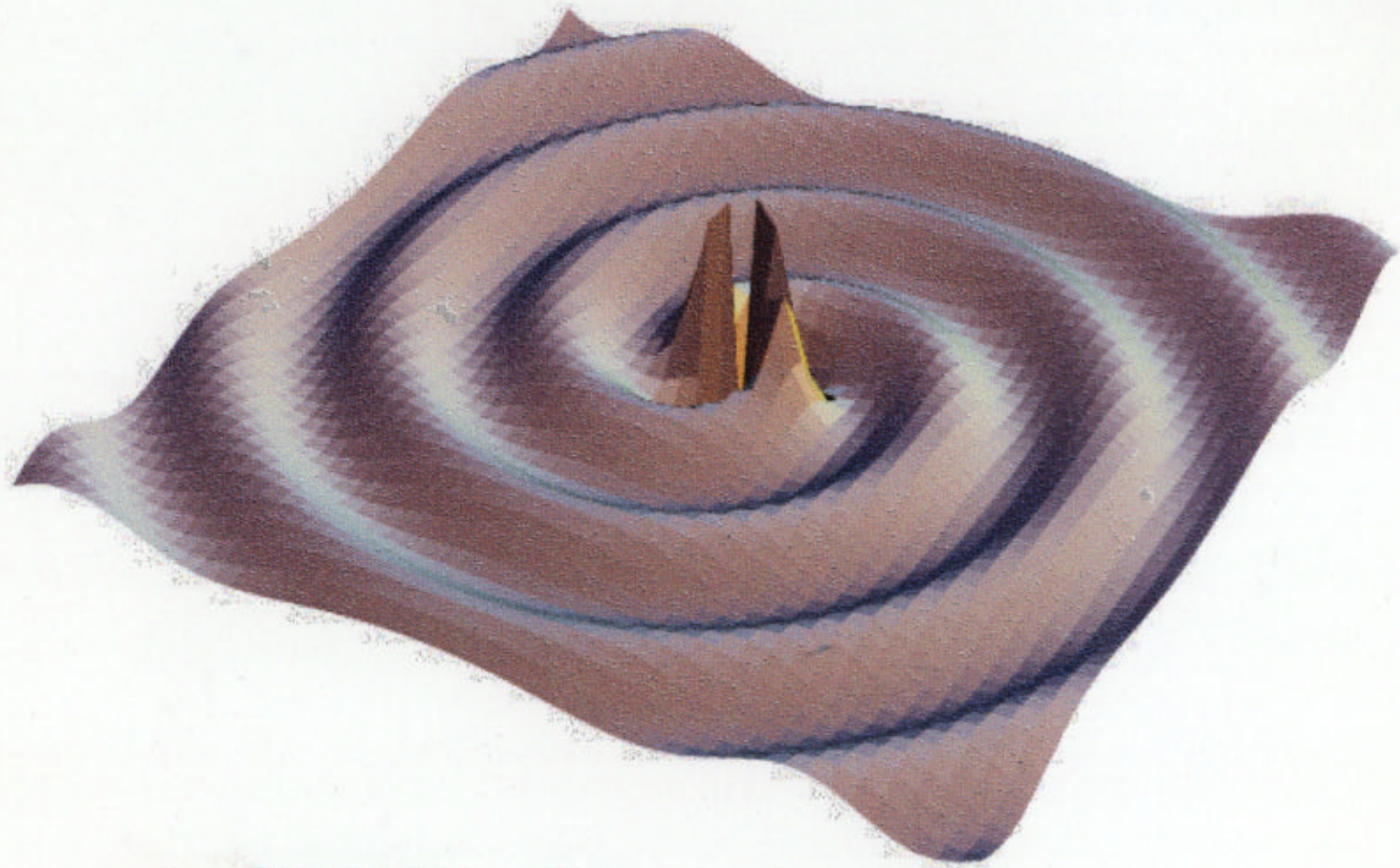
IN

SPACETIME CURVATURE.



# GRAVITATIONAL WAVES

## PROPAGATING RIPPLES IN CURVATURE



● SPEED  $c$

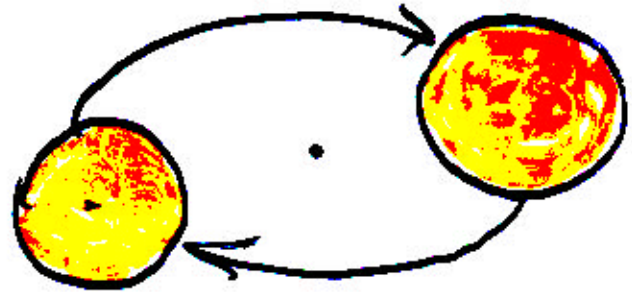
● TWO POLARIZATIONS

# MASS IN MOTION

## THE SOURCE OF GRAVITATIONAL WAVES

### BINARY STARS

$$L_{\text{GW}} \sim L_{\text{EM}} \left( \frac{M}{M_{\odot}} \frac{1 \text{ yr}}{P} \right)^{10/3}$$



EXPLOSIVE COLLAPSE  
IS MUCH BRIGHTER

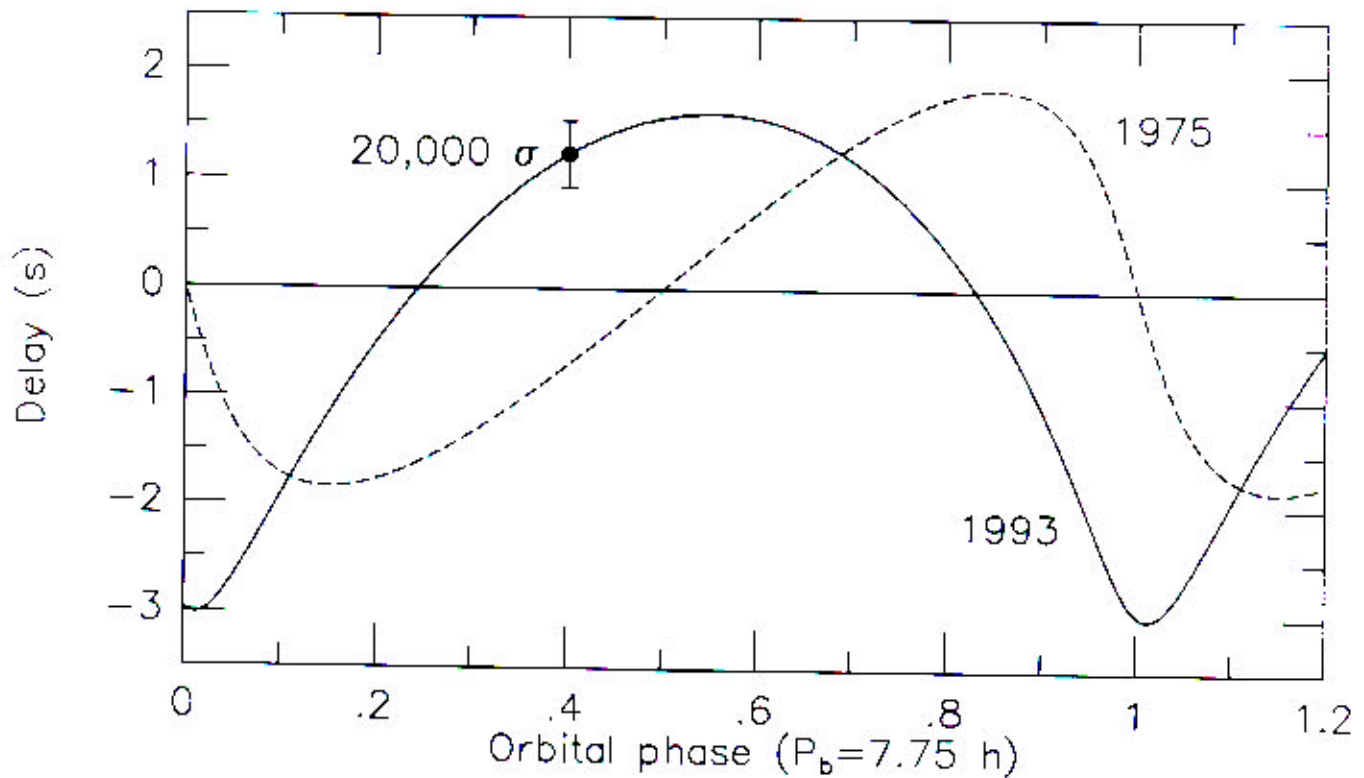
BINARY STARS  
MASS  $M$ , PERIOD  $P$ .

### ONCE PRODUCED, LITTLE IS ABSORBED

WE COULD POTENTIALLY SEE  
TO THE SURFACE OF BLACK HOLES,  
AND TO THE EARLIEST TIMES  
OF THE BIG BANG WITH  
GRAVITATIONAL WAVES



# PSR B1913+16



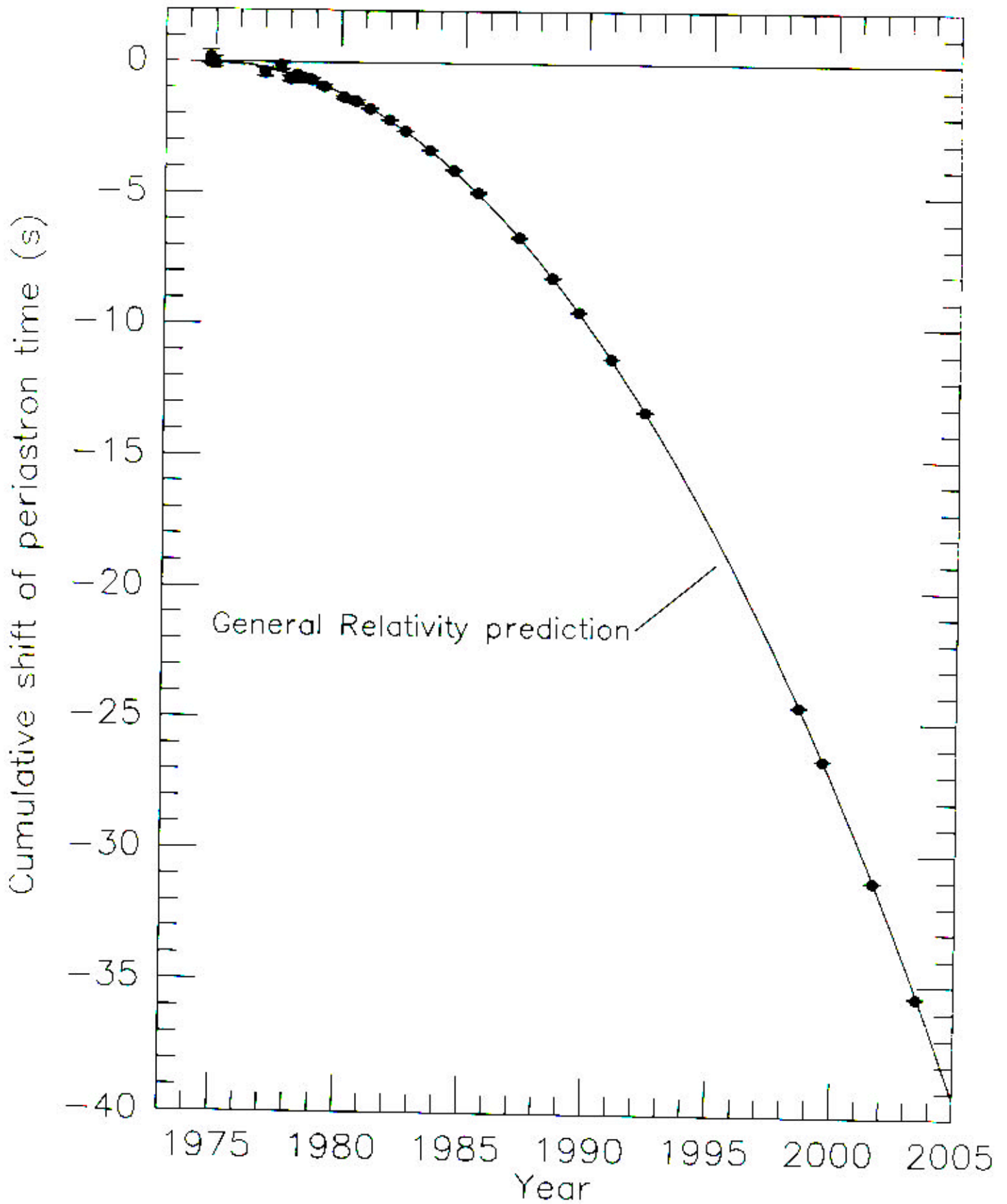
$$P_{\text{rot}} = 0.059029997929613 \pm .00000000000000007 \text{ s}$$

(July 7, 1984 ~ 6h GMT)

↑  
**PLAN AHEAD!**

$$P_{\text{orbit}} = 7.75 \text{ h.}$$

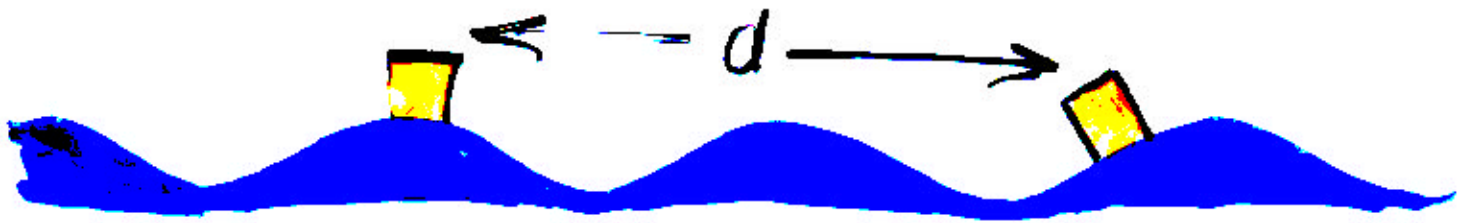
$$\text{(DECREASE DUE TO GRAV. RAD)} \cong 10 \mu\text{s/yr.}$$



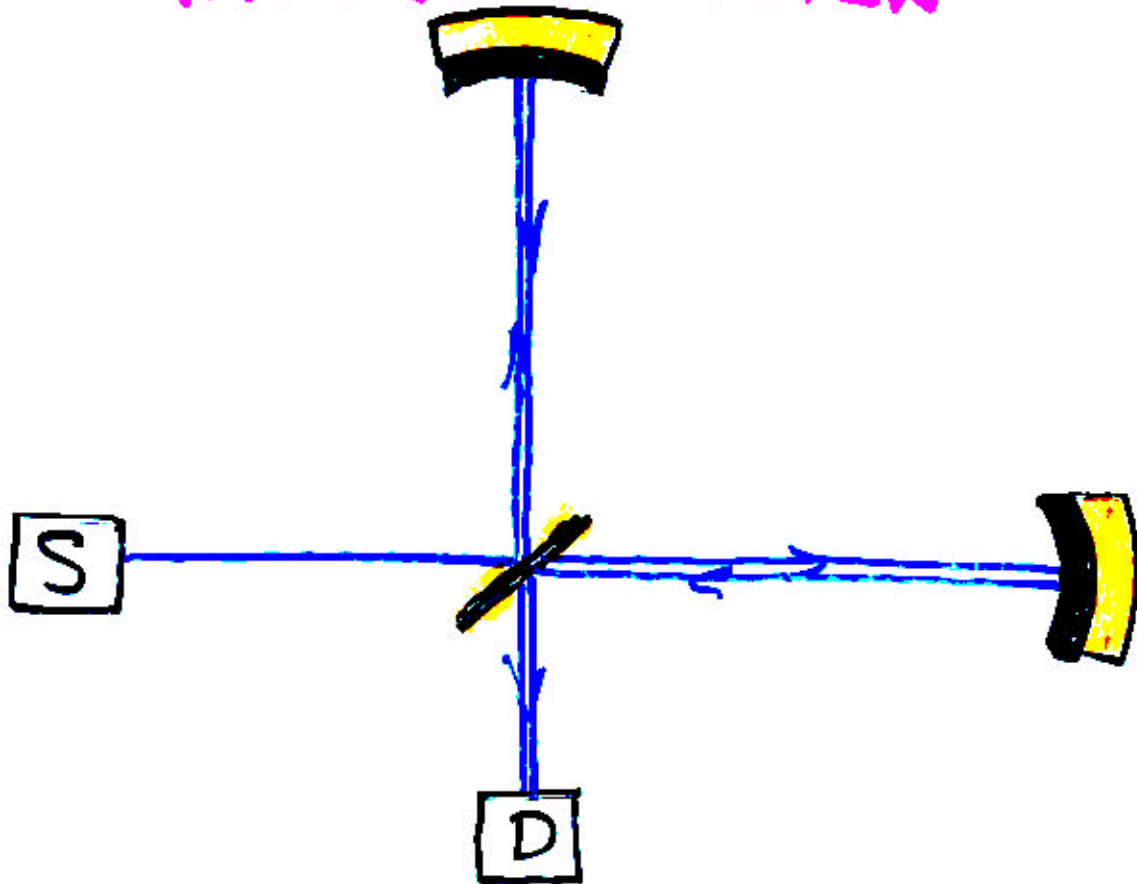
TAYLOR & WEISBERG



# DETECTION OF GRAVITATIONAL WAVES



## LASER INTERFEROMETER



BIG INTERFEROMETERS CAN DETECT DIFFERENCES IN LENGTH OF A FRACTION OF THE SIZE OF THE NUCLEUS OF AN ATOM !

# Gravitational Wave Detectors

- Interferometric
- Resonant-Mass





LIGO



HANFORD, WA.

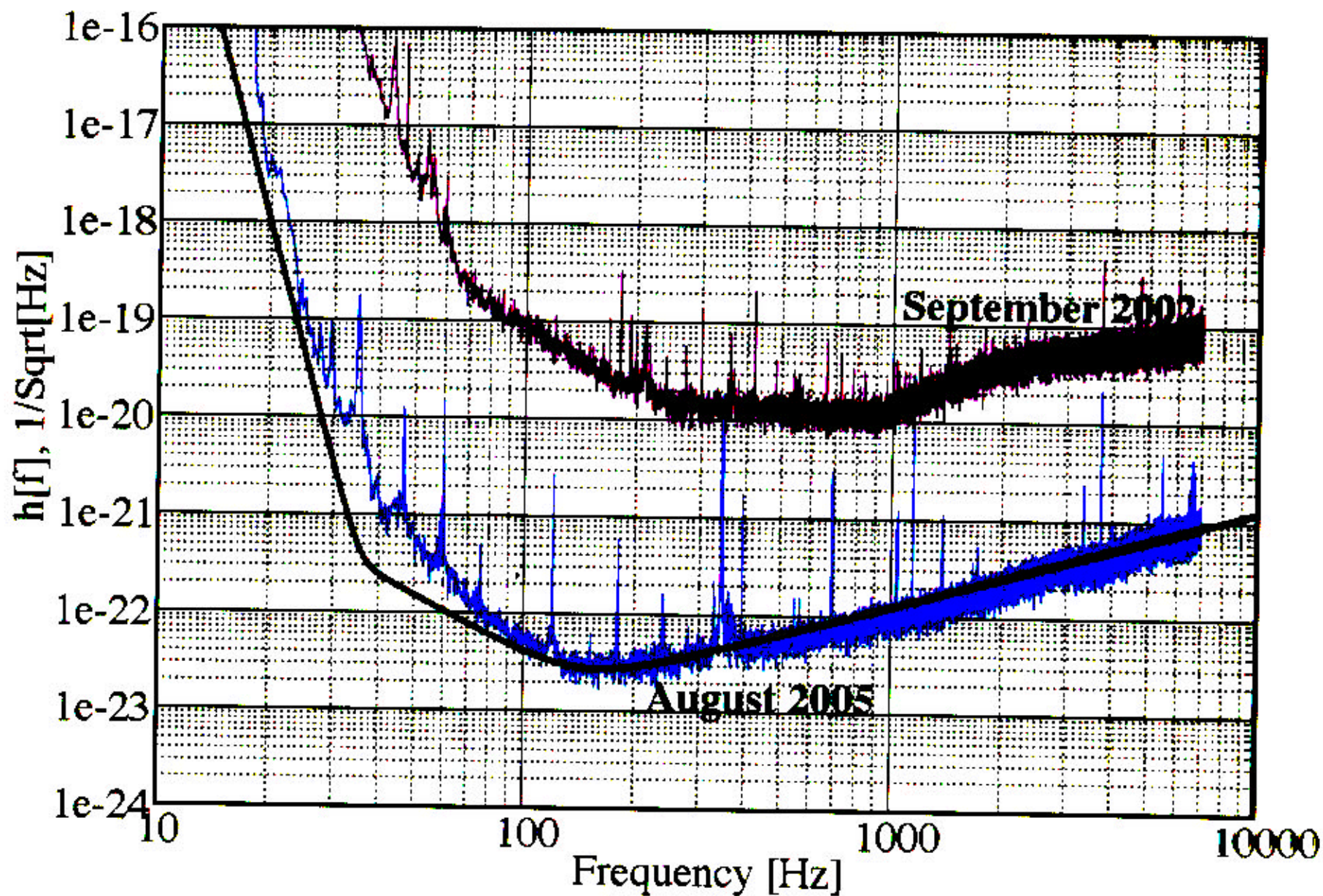


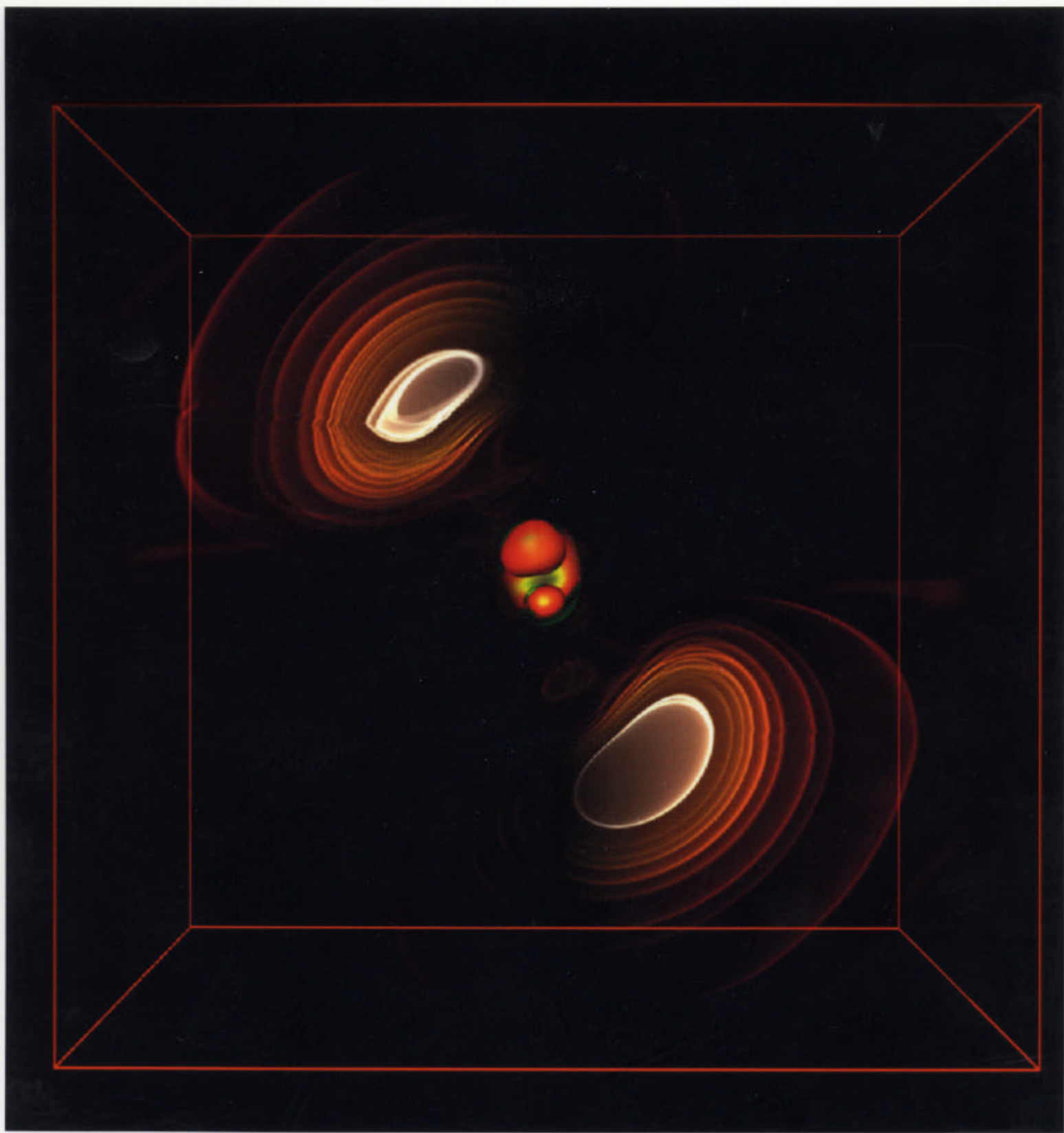




# HANFORD 4km

## Strain Sensivities for the LIGO Interferometers

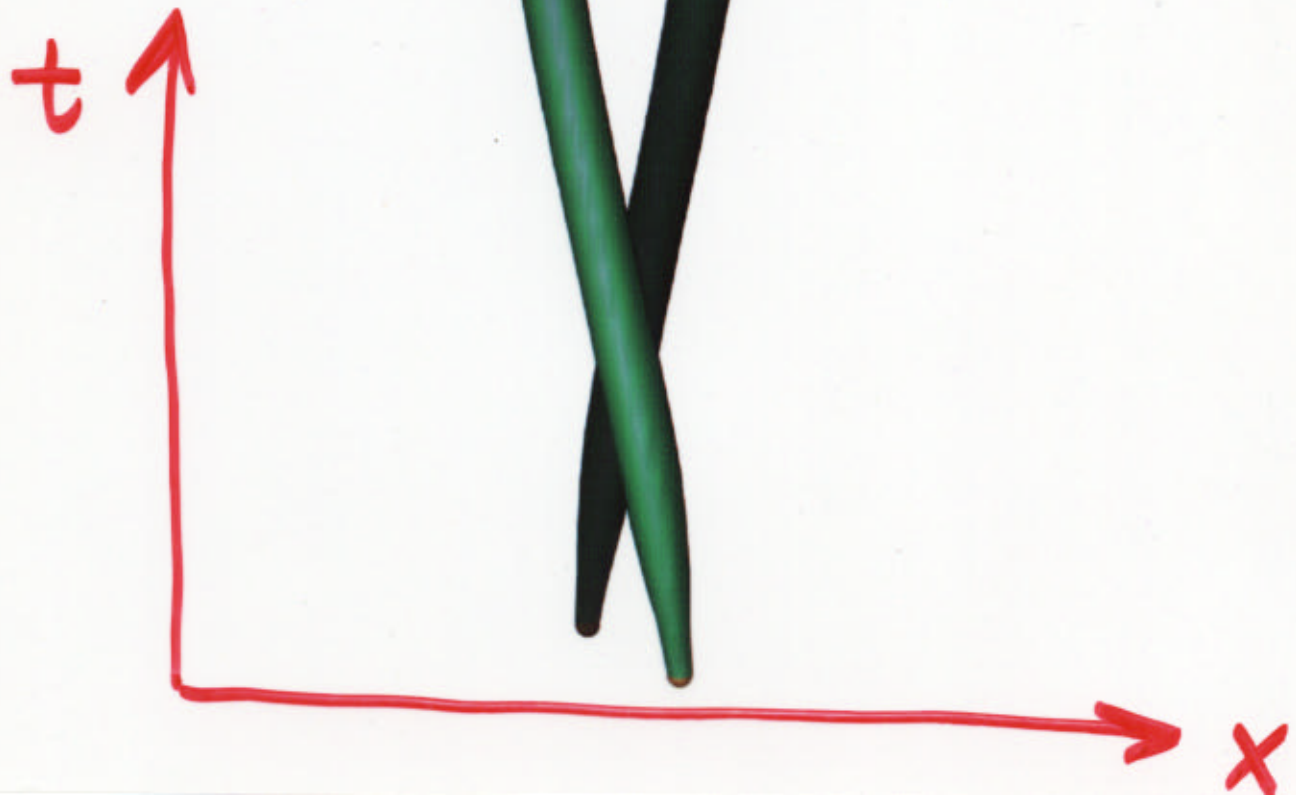




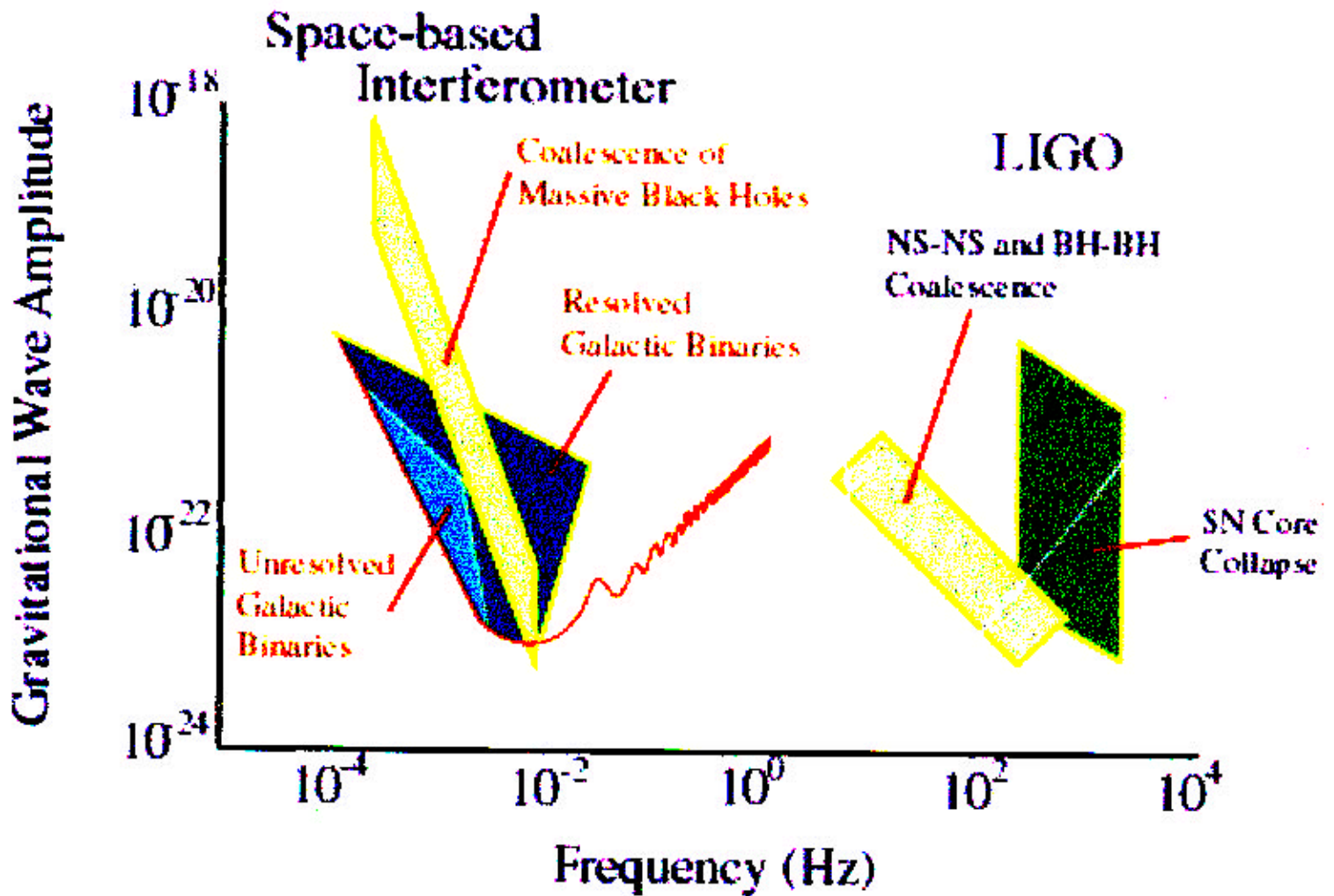


DANCING  
BLACK HOLES

(P. DIENER)

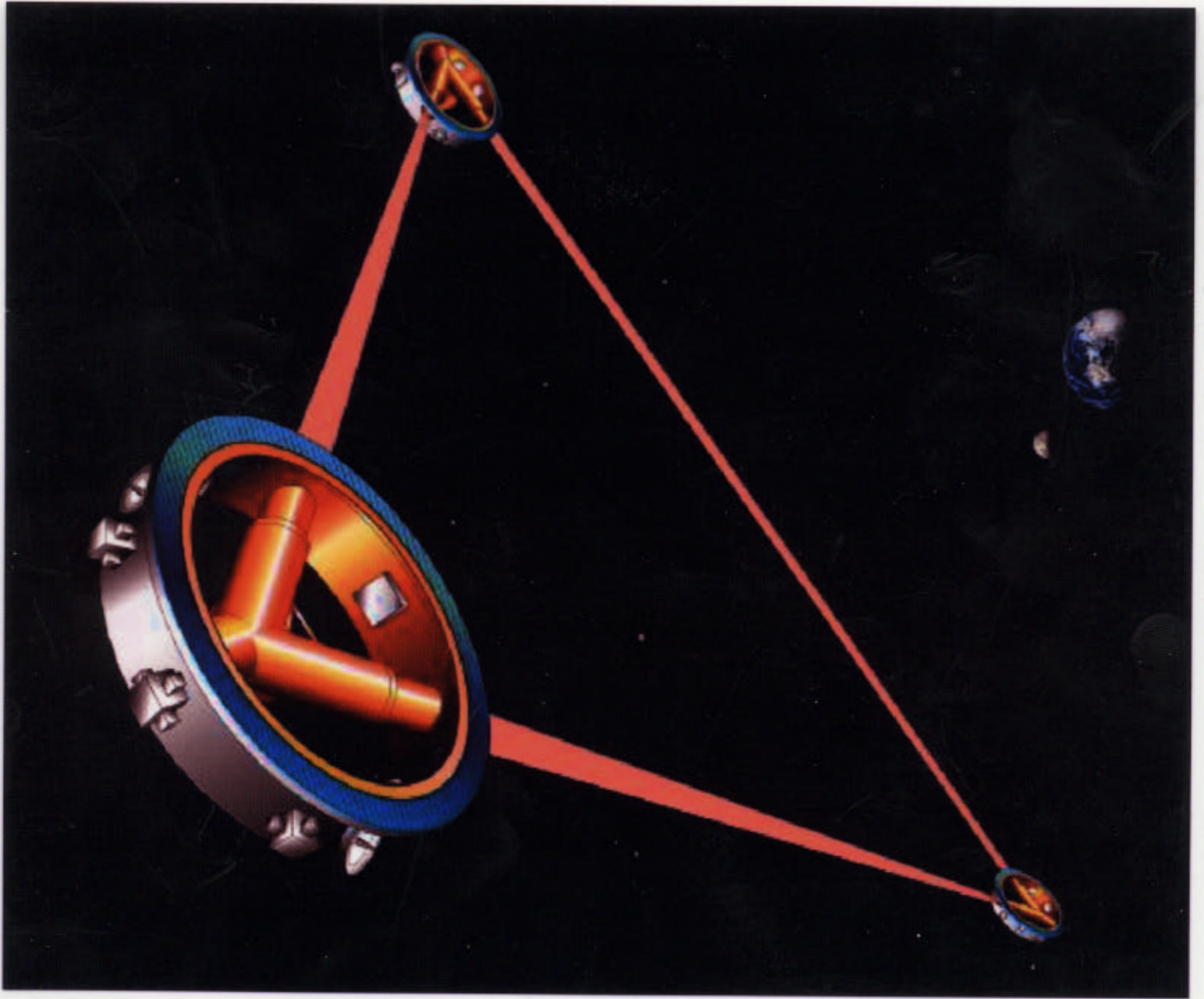


EVENT HORIZON  
OF TWO MERGING  
BLACK HOLES



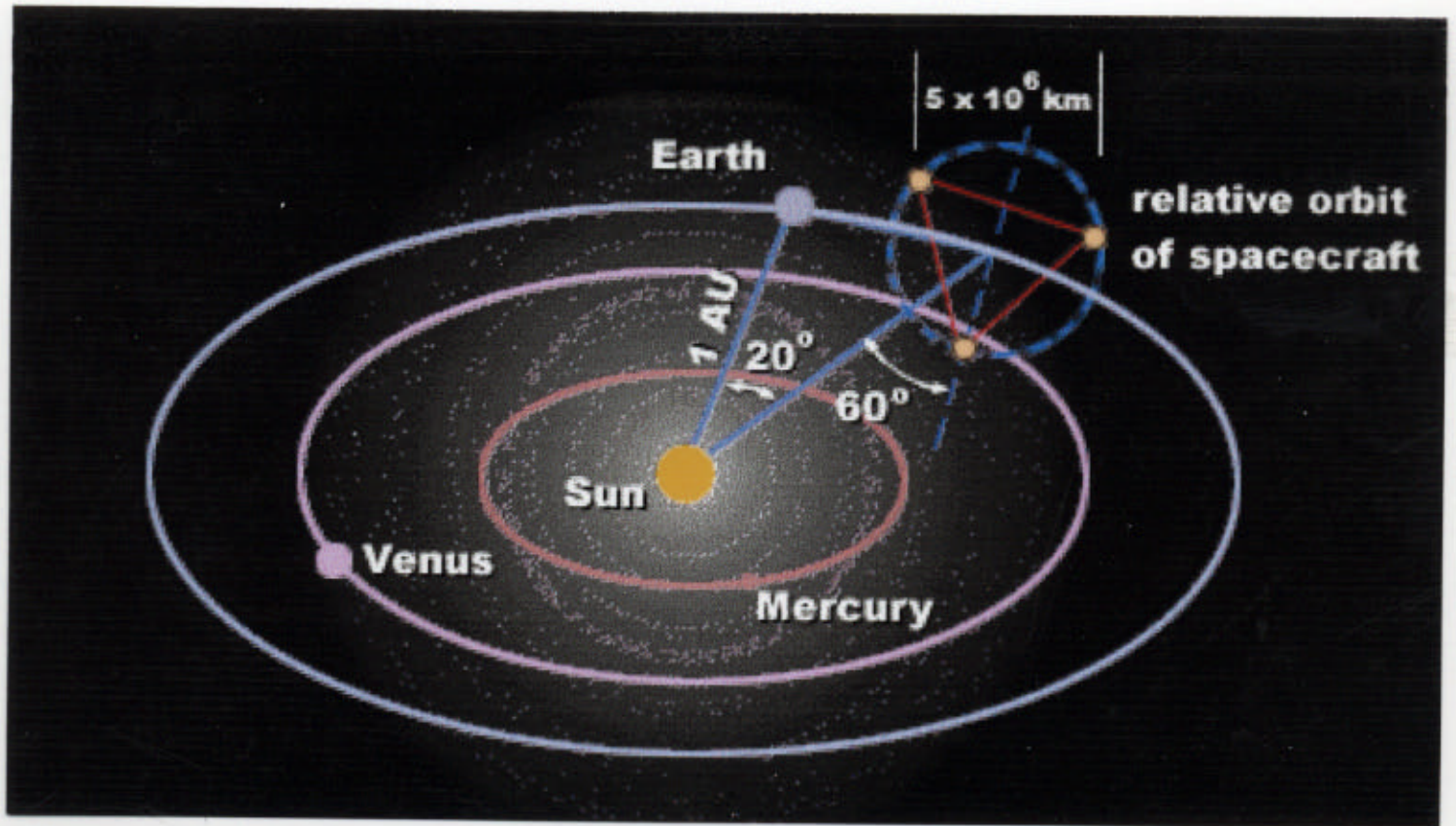
MASS IN MOTION PRODUCES  
GRAVITATIONAL WAVES.

Walker

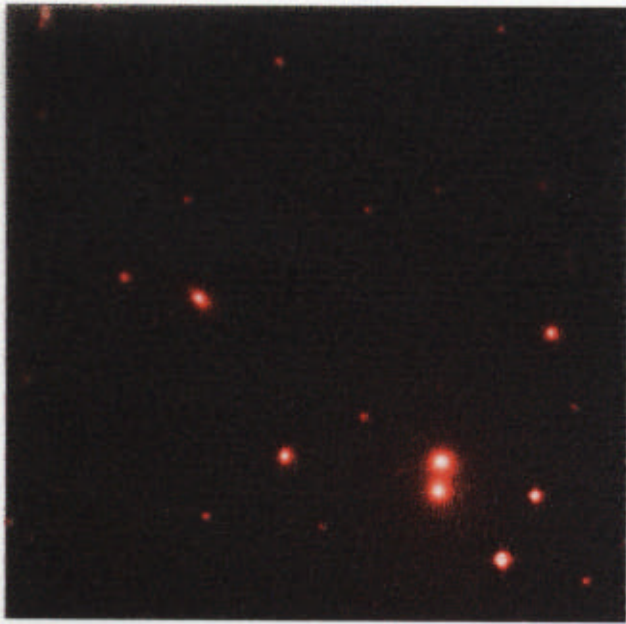


LISA

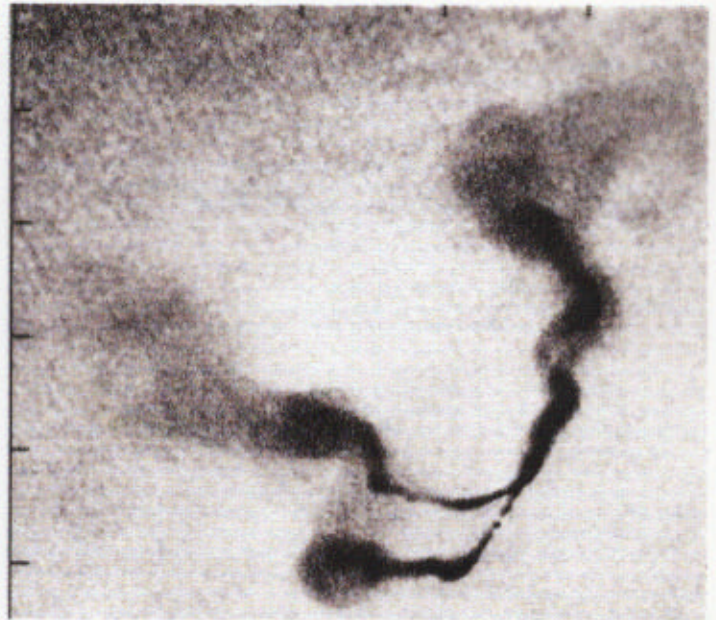




# Abell 400/ 3C75



Optical



Radio, Owen et al.

MERGING BLACK HOLES EMIT

$$E \sim (\text{fraction}) Mc^2$$

IN A TIME

$$t \sim (\text{few}) \times \frac{GM}{c^3}$$

$$L_{\text{GW}} \sim \frac{Mc^2}{GM/c^3} \sim \frac{c^5}{G} \sim 10^{59} \text{ erg/sec}$$

$$L_{\text{GAMMA RAY BURST}} \sim 10^{52} \text{ erg/sec.}$$

MAXIMUM PHYSICAL LUMINOSITY.



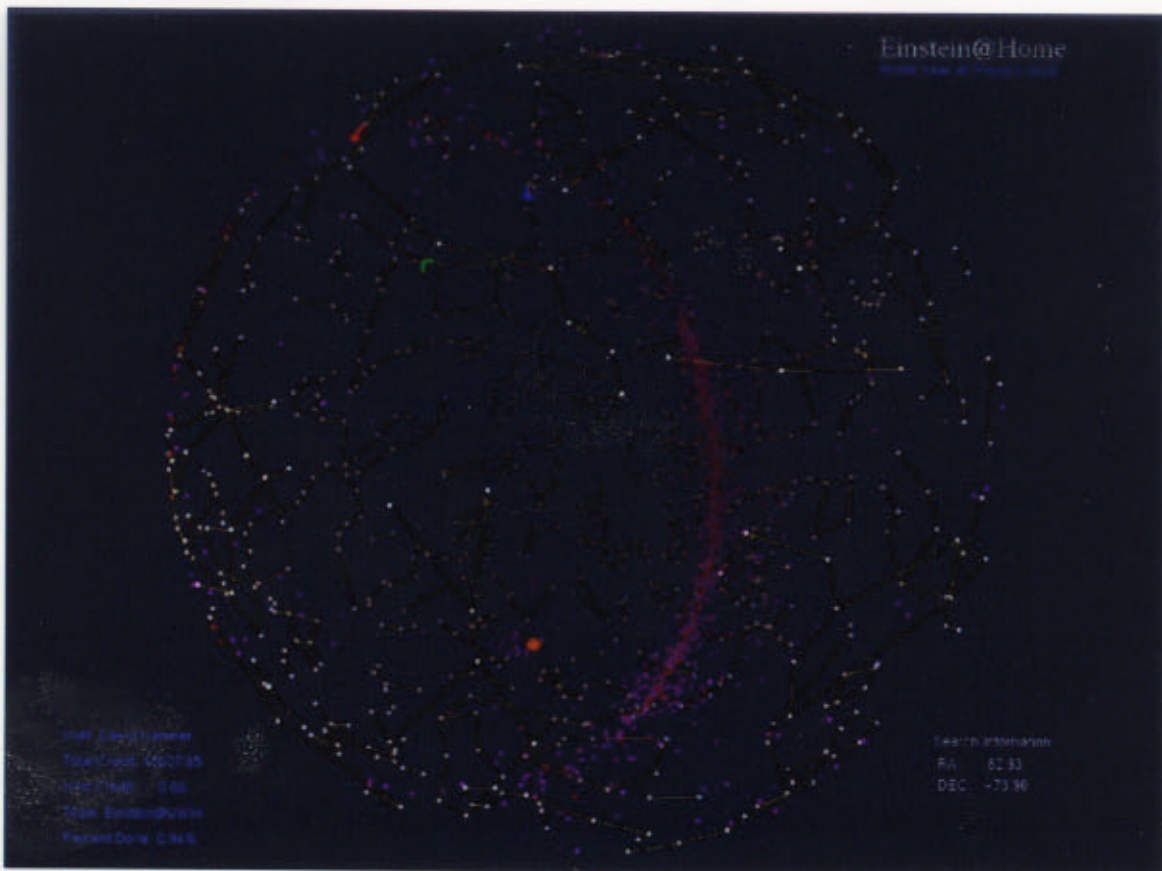
## Analysis of LIGO data for gravitational waves from binary neutron stars

B. Abbott,<sup>17</sup> K. Abbott,<sup>16</sup> R. Adhikari,<sup>14</sup> A. Ageev,<sup>21,28</sup> B. Allen,<sup>40</sup> R. Amin,<sup>35</sup> S. B. Anderson,<sup>15</sup> W. G. Anderson,<sup>41</sup> M. Araya,<sup>7</sup> H. Armandula,<sup>3</sup> F. Asiri,<sup>13a</sup> P. Aufmuth,<sup>32</sup> C. Aulbert,<sup>1</sup> S. Babak,<sup>7</sup> R. Balasubramanian,<sup>1</sup> S. Ballmer,<sup>4</sup> B. C. Barish,<sup>12</sup> D. Barker,<sup>15</sup> C. Barker-Patton,<sup>15</sup> M. Barnes,<sup>13</sup> B. Barr,<sup>36</sup> M. A. Barton,<sup>13</sup> K. Bayer,<sup>14</sup> R. Beausoleil,<sup>27,6</sup> K. Betczynski,<sup>27</sup> R. Bennett,<sup>36a</sup> S. J. Berukoff,<sup>1,0</sup> J. Betzwieser,<sup>14</sup> B. Bhawal,<sup>13</sup> I. A. Bilenko,<sup>21</sup> G. Billingsley,<sup>13</sup> E. Black,<sup>1</sup> K. Blackburn,<sup>7</sup> B. Bland-Weaver,<sup>15</sup> B. Bochner,<sup>14,a</sup> L. Bogue,<sup>13</sup> R. Bork,<sup>13</sup> S. Bose,<sup>41</sup> P. R. Brady,<sup>40</sup> V. B. Braginsky,<sup>2</sup> J. F. Brau,<sup>38</sup> D. A. Brown,<sup>40</sup> S. Brozek,<sup>32,1</sup> A. Bullington,<sup>27</sup> A. Buonanno,<sup>6,8</sup> R. Burgess,<sup>14</sup> D. Busby,<sup>13</sup> W. E. Butler,<sup>36</sup> R. L. Byer,<sup>12</sup> L. Cadonati,<sup>14</sup> G. Cagnoli,<sup>36</sup> J. B. Camp,<sup>22</sup> C. A. Cantley,<sup>36</sup> L. Cardenas,<sup>13</sup> K. Carter,<sup>16</sup> M. M. Casey,<sup>36</sup> J. Castiglione,<sup>25</sup> A. Chandler,<sup>14</sup> J. Chapsky,<sup>13,b</sup> P. Charlton,<sup>13</sup> S. Chatterji,<sup>14</sup> Y. Chen,<sup>6</sup> V. Chikarmane,<sup>7</sup> D. Chin,<sup>7</sup> N. Christensen,<sup>8</sup> D. Churches,<sup>7</sup> C. Colacino,<sup>32,2</sup> R. Coldwell,<sup>35</sup> M. Coles,<sup>16,1</sup> D. Cook,<sup>15</sup> T. Corbitt,<sup>14</sup> D. Coyne,<sup>12</sup> J. D. E. Creighton,<sup>40</sup> T. D. Creighton,<sup>3</sup> D. R. M. Crooks,<sup>36</sup> P. Csatorday,<sup>14</sup> B. J. Cusack,<sup>3</sup> C. Cutler,<sup>1</sup> E. D'Ambrosio,<sup>13</sup> K. Danzmann,<sup>32,20</sup> R. Davies,<sup>7</sup> E. Daw,<sup>17,d</sup> D. DeBra,<sup>27</sup> T. Delker,<sup>35,k</sup> R. DeSalvo,<sup>3</sup> S. Dhurandhar,<sup>12</sup> M. Diaz,<sup>40</sup> H. Ding,<sup>12</sup> R. W. P. Drever,<sup>4</sup> R. J. Dupuis,<sup>36</sup> C. Ebeling,<sup>8</sup> J. Edlund,<sup>13</sup> P. Ehrens,<sup>13</sup> E. J. Elliffe,<sup>36</sup> T. Etzel,<sup>13</sup> M. Evans,<sup>15</sup> T. Evans,<sup>16</sup> C. Fallnich,<sup>32</sup> D. Farnham,<sup>13</sup> M. M. Fejer,<sup>27</sup> M. Fine,<sup>13</sup> L. S. Finn,<sup>29</sup> E. Flanagan,<sup>9</sup> A. Freise,<sup>24</sup> R. Frey,<sup>38</sup> P. Fritschel,<sup>14</sup> V. Frolov,<sup>16</sup> M. Fyffe,<sup>16</sup> K. S. Ganezer,<sup>5</sup> J. A. Giaime,<sup>17</sup> A. Gillespie,<sup>13,m</sup> K. Goda,<sup>14</sup> G. Gonzalez,<sup>17</sup> S. Goßler,<sup>32</sup> P. Grandclément,<sup>24</sup> A. Grant,<sup>36</sup> C. Gray,<sup>15</sup> A. M. Gretarsson,<sup>16</sup> D. Grmnetz,<sup>13</sup> H. Grote,<sup>2</sup> S. Grunewald,<sup>1</sup> M. Guenther,<sup>15</sup> E. Gustafson,<sup>27,n</sup> R. Gustafson,<sup>17</sup> W. O. Hamilton,<sup>17</sup> M. Hammond,<sup>16</sup> J. Hanson,<sup>16</sup> C. Hardham,<sup>27</sup> G. Harry,<sup>14</sup> A. Hartunian,<sup>13</sup> J. Heefner,<sup>13</sup> Y. Hefetz,<sup>14</sup> G. Heinzl,<sup>2</sup> I. S. Heng,<sup>32</sup> M. Hennessy,<sup>27</sup> N. Hepler,<sup>29</sup> A. Heptonstall,<sup>36</sup> M. Heurs,<sup>32</sup> M. Hewitson,<sup>36</sup> N. Hindman,<sup>15</sup> P. Hoang,<sup>13</sup> J. Hough,<sup>36</sup> M. Hyrnyevych,<sup>13,0</sup> W. Hua,<sup>27</sup> R. Ingley,<sup>34</sup> M. Ito,<sup>38</sup> Y. Itoh,<sup>1</sup> A. Ivanov,<sup>14</sup> O. Jennrich,<sup>36,o</sup> W. W. Johnson,<sup>17</sup> W. Johnston,<sup>30</sup> L. Jones,<sup>13</sup> D. Jungwirth,<sup>13,q</sup> V. Kalogera,<sup>24</sup> E. Katsavounidis,<sup>4</sup> K. Kawabe,<sup>20,2</sup> S. Kawamura,<sup>23</sup> W. Kells,<sup>13</sup> J. Kern,<sup>6</sup> A. Khan,<sup>16</sup> S. Killbourn,<sup>36</sup> C. J. Killow,<sup>36</sup> C. Kim,<sup>24</sup> C. King,<sup>13</sup> P. King,<sup>13</sup> S. Klimenko,<sup>15</sup> P. Kloeveborn,<sup>2</sup> S. Koranda,<sup>40</sup> K. Kötter,<sup>32</sup> J. Kovalik,<sup>16</sup> D. Kozak,<sup>13</sup> B. Krishnan,<sup>1</sup> M. Landry,<sup>15</sup> J. Langdale,<sup>16</sup> B. Lantz,<sup>27</sup> R. Lawrence,<sup>14</sup> A. Lazzarini,<sup>13</sup> M. Lei,<sup>13</sup> V. Leonhardt,<sup>32</sup> I. Leonor,<sup>38</sup> K. Libbrecht,<sup>14</sup> P. Lindquist,<sup>13</sup> S. Liu,<sup>15</sup> J. Logan,<sup>13,r</sup> M. Lormand,<sup>16</sup> M. Lubinski,<sup>15</sup> H. Lück,<sup>32,2</sup> T. T. Lyons,<sup>13,t</sup> B. Machenschalk,<sup>1</sup> M. MacInnis,<sup>14</sup> M. Mageswaran,<sup>13</sup> K. Mailand,<sup>13</sup> W. Majid,<sup>13,b</sup> M. Malec,<sup>32</sup> F. Mann,<sup>13</sup> A. Marin,<sup>14,s</sup> S. Märka,<sup>15</sup> E. Maros,<sup>15</sup> J. Mason,<sup>15</sup> K. Mason,<sup>14</sup> O. Matherny,<sup>15</sup> L. Matone,<sup>15</sup> N. Mavalvala,<sup>14</sup> R. McCarthy,<sup>15</sup> D. E. McClelland,<sup>7</sup> M. McHugh,<sup>14</sup> P. McNamara,<sup>36,0</sup> G. Mendell,<sup>15</sup> S. Meshkov,<sup>15</sup> C. Messenger,<sup>34</sup> V. P. Mitrofanov,<sup>21</sup> G. Mitselmakher,<sup>35</sup> R. Mittleman,<sup>4</sup> O. Miyakawa,<sup>13</sup> S. Miyoki,<sup>35,s</sup> S. Mohanty,<sup>36</sup> G. Moreno,<sup>15</sup> K. Mossavi,<sup>2</sup> B. Mours,<sup>13,x</sup> G. Mueller,<sup>35</sup> S. Mukherjee,<sup>16</sup> J. Myers,<sup>15</sup> S. Nagano,<sup>2</sup> T. Nash,<sup>10,y</sup> H. Naundorf,<sup>1</sup> R. Nayak,<sup>12</sup> G. Newton,<sup>36</sup> F. Nocera,<sup>13</sup> P. Nutzman,<sup>24</sup> T. Olson,<sup>25</sup> B. O'Reilly,<sup>16</sup> D. J. Ottaway,<sup>14</sup> A. Ottewill,<sup>40,z</sup> D. Oulmette,<sup>13,9</sup> H. Overmier,<sup>26</sup> B. J. Owen,<sup>29</sup> M. A. Papa,<sup>1</sup> C. Parameswaran,<sup>16</sup> V. Parameswaran,<sup>15</sup> M. Pedraza,<sup>13</sup> S. Penn,<sup>11</sup> M. Pitkin,<sup>36</sup> M. Plišt,<sup>36</sup> M. Pratt,<sup>14</sup> V. Quetschke,<sup>32</sup> F. Raab,<sup>15</sup> H. Radkins,<sup>15</sup> R. Rahkola,<sup>38</sup> M. Rakhmanov,<sup>35</sup> S. R. Rao,<sup>13</sup> D. Redding,<sup>13,b</sup> M. W. Regehr,<sup>14,1</sup> T. Regimbau,<sup>14</sup> K. I. Reilly,<sup>13</sup> K. Reithmaier,<sup>13</sup> D. H. Reitze,<sup>15</sup> S. Richman,<sup>14,27</sup> R. Riesen,<sup>16</sup> K. Riles,<sup>5</sup> A. Rizzi,<sup>6,28</sup> D. I. Robertson,<sup>36</sup> N. A. Robertson,<sup>36,27</sup> L. Robison,<sup>15</sup> S. Roddy,<sup>16</sup> J. Rollins,<sup>14</sup> J. D. Romano,<sup>30,29</sup> J. Romie,<sup>11</sup> H. Rong,<sup>35,0</sup> D. Rose,<sup>13</sup> E. Rothhoff,<sup>29</sup> S. Rowan,<sup>36</sup> A. Rüdiger,<sup>20,2</sup> P. Russell,<sup>13</sup> K. Ryan,<sup>15</sup> I. Salzman,<sup>14</sup> G. H. Sanders,<sup>13</sup> V. Sannibale,<sup>13</sup> B. Sathyaprakash,<sup>7</sup> P. R. Saulson,<sup>28</sup> R. Savage,<sup>15</sup> A. Sazonov,<sup>35</sup> R. Schilling,<sup>20,2</sup> K. Schlautman,<sup>29</sup> V. Schmidt,<sup>14,30</sup> R. Scholfield,<sup>38</sup> M. Schrempel,<sup>32,ee</sup> B. F. Schutz,<sup>17</sup> P. Schwinberg,<sup>12</sup> S. M. Scott,<sup>1</sup> A. C. Searle,<sup>7</sup> B. Sears,<sup>13</sup> S. Seel,<sup>15</sup> A. S. Sengupta,<sup>12</sup> C. A. Shapiro,<sup>29,1</sup> P. Shawhan,<sup>13</sup> D. H. Shoemaker,<sup>14</sup> Q. Z. Shu,<sup>35,8f</sup> A. Sibley,<sup>16</sup> X. Siemens,<sup>40</sup> L. Sievers,<sup>13,h</sup> D. Sigg,<sup>15</sup> A. M. Sintes,<sup>13,h</sup> K. Skeldon,<sup>36</sup> J. R. Smith,<sup>3</sup> M. Smith,<sup>14</sup> M. R. Smith,<sup>13</sup> P. Sneddon,<sup>36</sup> R. Spero,<sup>13,h</sup> G. Stappert,<sup>16</sup> K. A. Strain,<sup>36</sup> D. Strom,<sup>38</sup> A. Stuver,<sup>29</sup> T. Summerscales,<sup>29</sup> M. C. Sumner,<sup>13</sup> P. J. Sutton,<sup>26,3</sup> J. Sylvestre,<sup>14</sup> A. Takamori,<sup>13</sup> D. B. Tanner,<sup>35</sup> H. Tariq,<sup>13</sup> I. Taylor,<sup>7</sup> R. Taylor,<sup>13</sup> K. S. Thorne,<sup>6</sup> M. Tibbitts,<sup>29</sup> S. Tilay,<sup>13,hh</sup> M. Tinto,<sup>34</sup> K. V. Tokmakov,<sup>21</sup> C. Torres,<sup>30</sup> C. Torrie,<sup>13,36</sup> S. Traeger,<sup>32,1</sup> G. Traylor,<sup>16</sup> W. Tyler,<sup>13</sup> D. Ugolini,<sup>41</sup> M. Vallisneri,<sup>6,1</sup> M. van Putten,<sup>14</sup> S. Vass,<sup>13</sup> A. Vecchio,<sup>34</sup> C. Vorvick,<sup>15</sup> S. P. Vyachanin,<sup>21</sup> L. Wallace,<sup>14</sup> H. Walther,<sup>20</sup> H. Ward,<sup>36</sup> B. Ware,<sup>13,4</sup> K. Watts,<sup>16</sup> D. Webber,<sup>13</sup> A. Weidner,<sup>20,2</sup> U. Weiland,<sup>32</sup> A. Weinstein,<sup>14</sup> R. Weiss,<sup>14</sup> H. Welling,<sup>42</sup> L. Wen,<sup>15</sup> S. Wen,<sup>17</sup> J. T. Whelan,<sup>14</sup> S. E. Whitcomb,<sup>13</sup> B. F. Whiting,<sup>35</sup> P. A. Willems,<sup>14</sup> P. R. Williams,<sup>13,k</sup> R. Williams,<sup>4</sup> B. Willke,<sup>12,2</sup> A. Wilson,<sup>13</sup> B. J. Winjum,<sup>20,d</sup> W. Winkler,<sup>20,2</sup> S. Wise,<sup>35</sup> A. G. Wiseman,<sup>40</sup> G. Woan,<sup>36</sup> R. Woolley,<sup>16</sup> J. Worden,<sup>15</sup> I. Yakushin,<sup>16</sup> H. Yamamoto,<sup>3</sup> S. Yoshida,<sup>26</sup> I. Zawischa,<sup>32,1</sup> L. Zhang,<sup>15</sup> N. Zotov,<sup>18</sup> M. Zucker,<sup>16</sup> and J. Zweizig<sup>13</sup>

(LIGO Scientific Collaboration)<sup>000h</sup><sup>1</sup>Albert-Einstein Institut, Max-Planck-Institut für Gravitationsphysik, D-14476 Golm, Germany<sup>2</sup>Albert-Einstein Institut, Max-Planck-Institut für Gravitationsphysik, D-30167 Hannover, Germany<sup>3</sup>Australian National University, Canberra, 0200, Australia<sup>4</sup>California Institute of Technology, Pasadena, California 91125, USA<sup>5</sup>California State University Dominguez Hills, Carson, California 90747, USA<sup>6</sup>Caltech-CaRT, Pasadena, California 91125, USA<sup>7</sup>Cardiff University, Cardiff, CF2 3YB, United Kingdom<sup>8</sup>Carleton College, Northfield, Minnesota 55057, USA



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# BLACK HOLES

EINSTEIN'S THEORY PREDICTS  
THAT WHEN

MASS IS COMPRESSED

TO A SMALL ENOUGH VOLUME

THE GRAVITATIONAL FORCE

IS SO STRONG

NOTHING CAN ESCAPE,

# BLACK HOLES

- GENERAL RELATIVITY PREDICTS THAT THE GEOMETRY OF ASTROPHYSICAL BLACK HOLES IS CHARACTERIZED BY JUST TWO PARAMETERS:

$M$  — MASS

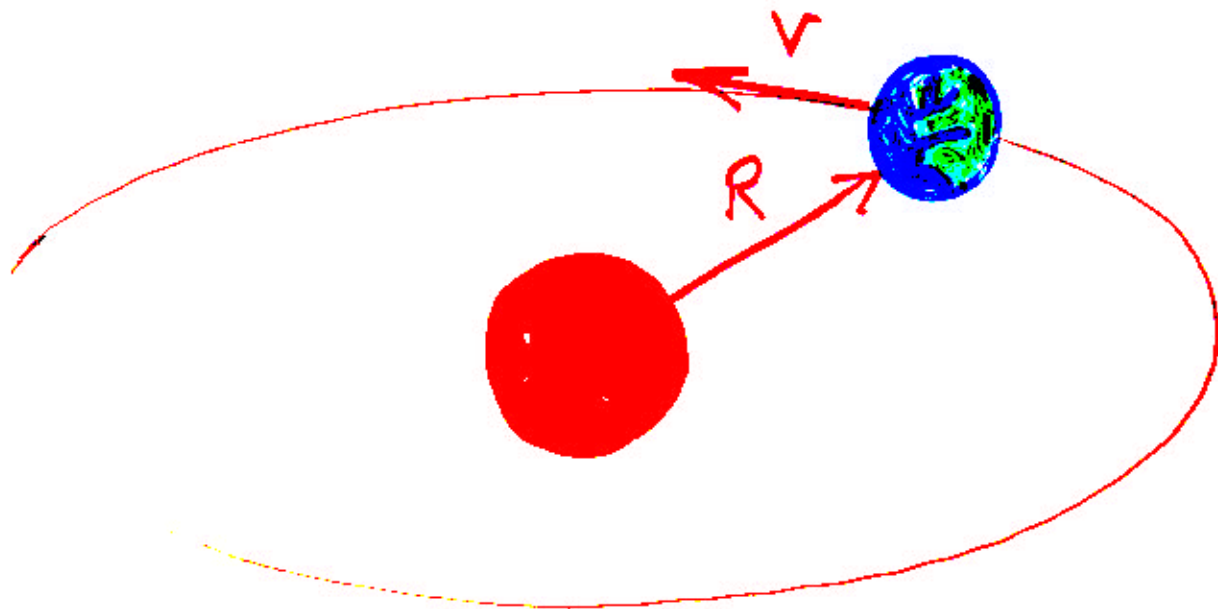
$J$  — ANGULAR MOMENTUM

- BLACK HOLES PROVIDE THE CLEANEST CONNECTION BETWEEN ASTROPHYSICS AND FUNDAMENTAL PHYSICS.



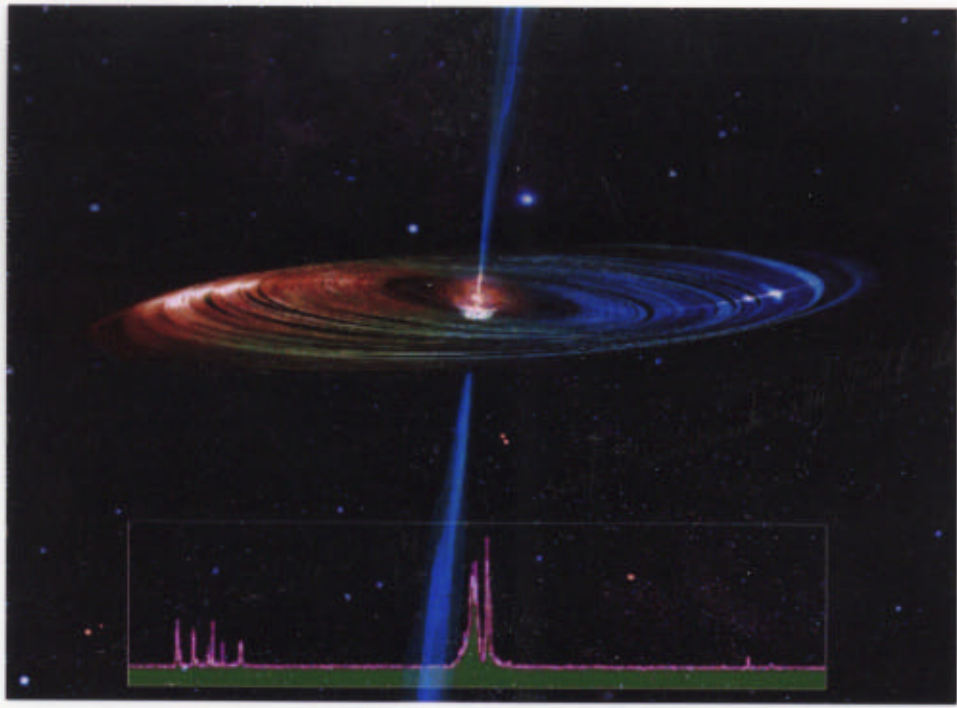
# WEIGHING THE SUN

THE MASS OF THE SUN CAN BE DETERMINED FROM THE SPEED OF THE EARTH AND THE SIZE OF ITS ORBIT.

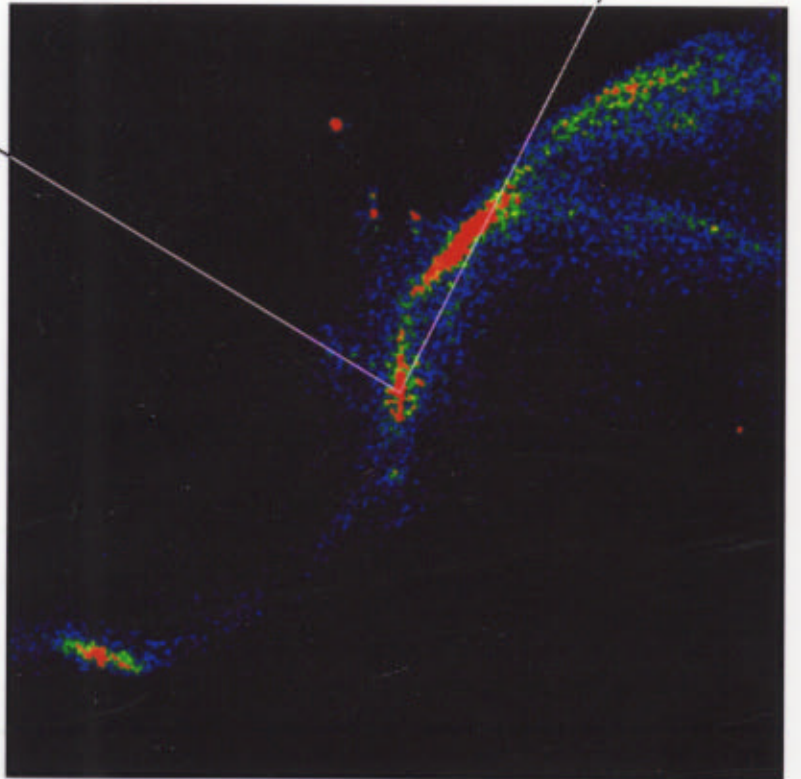
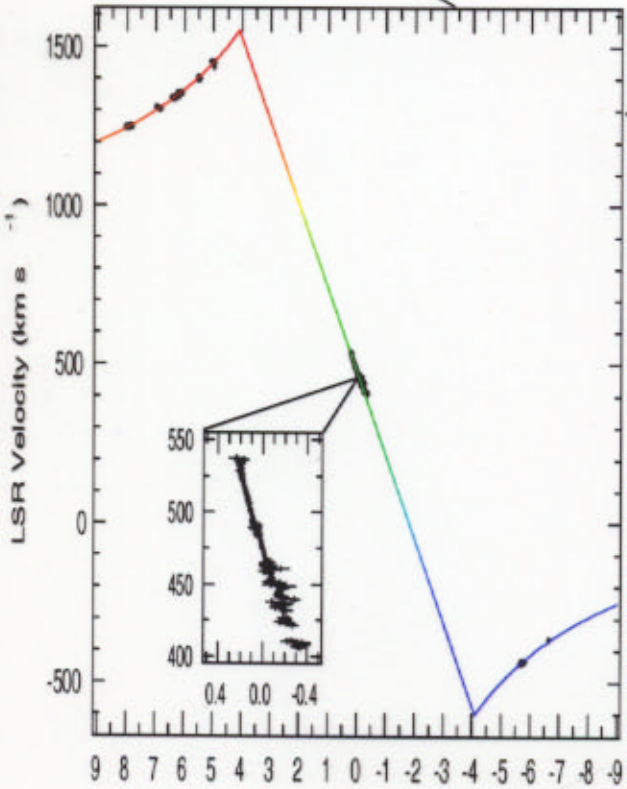
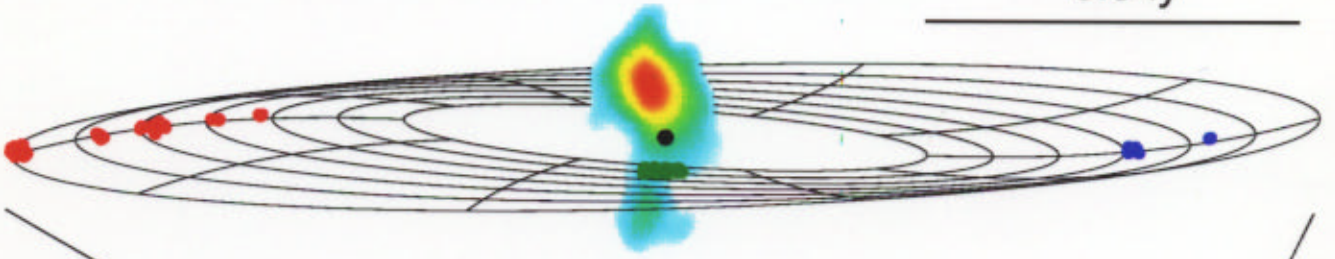


THE EARTH WOULD GO AROUND TWICE AS FAST IF THE SUN WERE FOUR TIMES MORE MASSIVE.

# NGC 4258

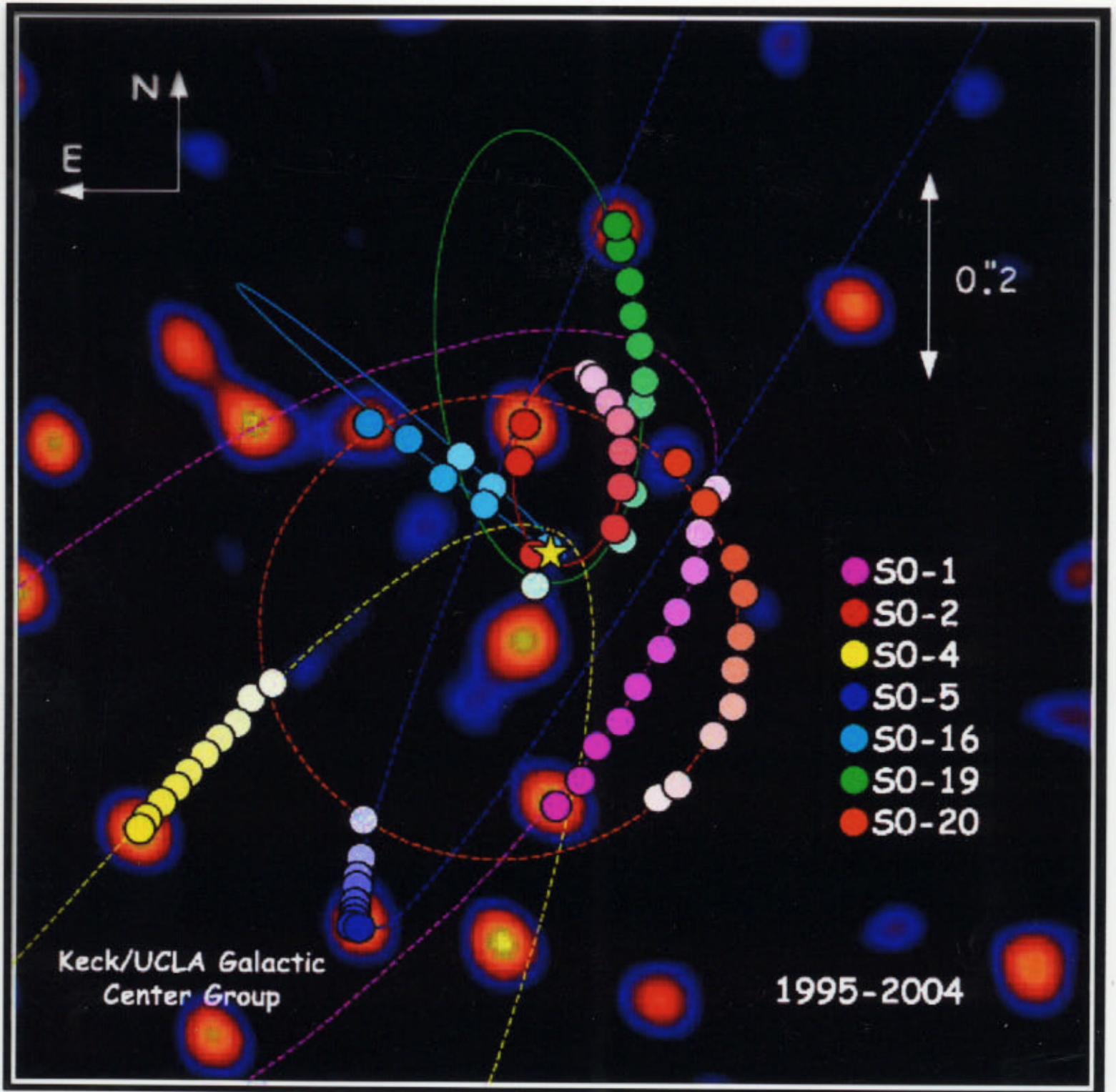


0.5 ly





# THE BLACK HOLE IN OUR GALAXY



(A. Ghez, et al. 2004)



BLACK HOLES POWER SOME OF THE MOST ENERGETIC PHENOMENA IN THE UNIVERSE.

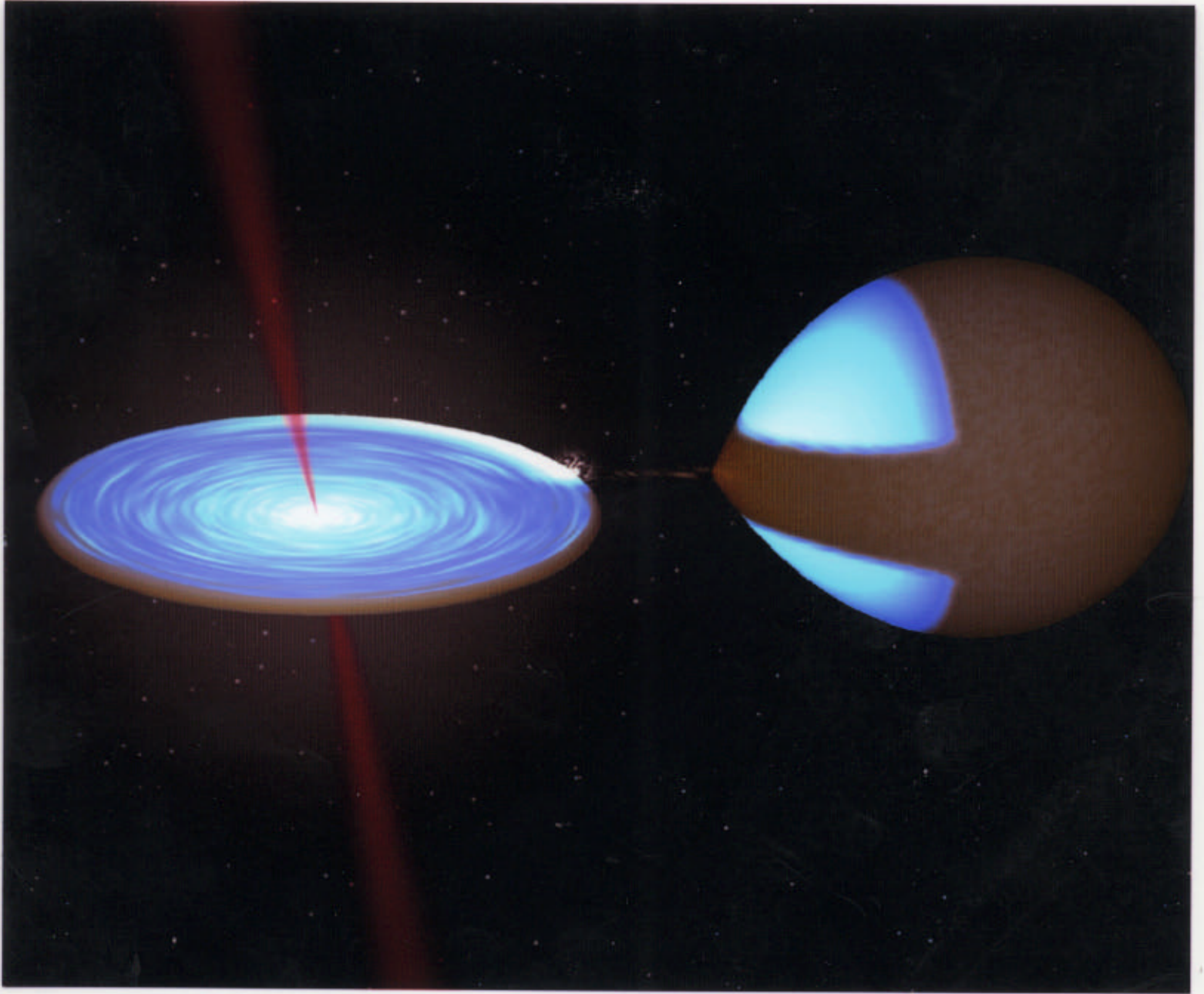
- X-RAY SOURCES

$$L_x \sim 10^{38} \text{ ergs/sec} \quad L_{\odot} \sim 10^{33} \text{ ergs/sec}$$

- ACTIVE GALACTIC NUCLEI

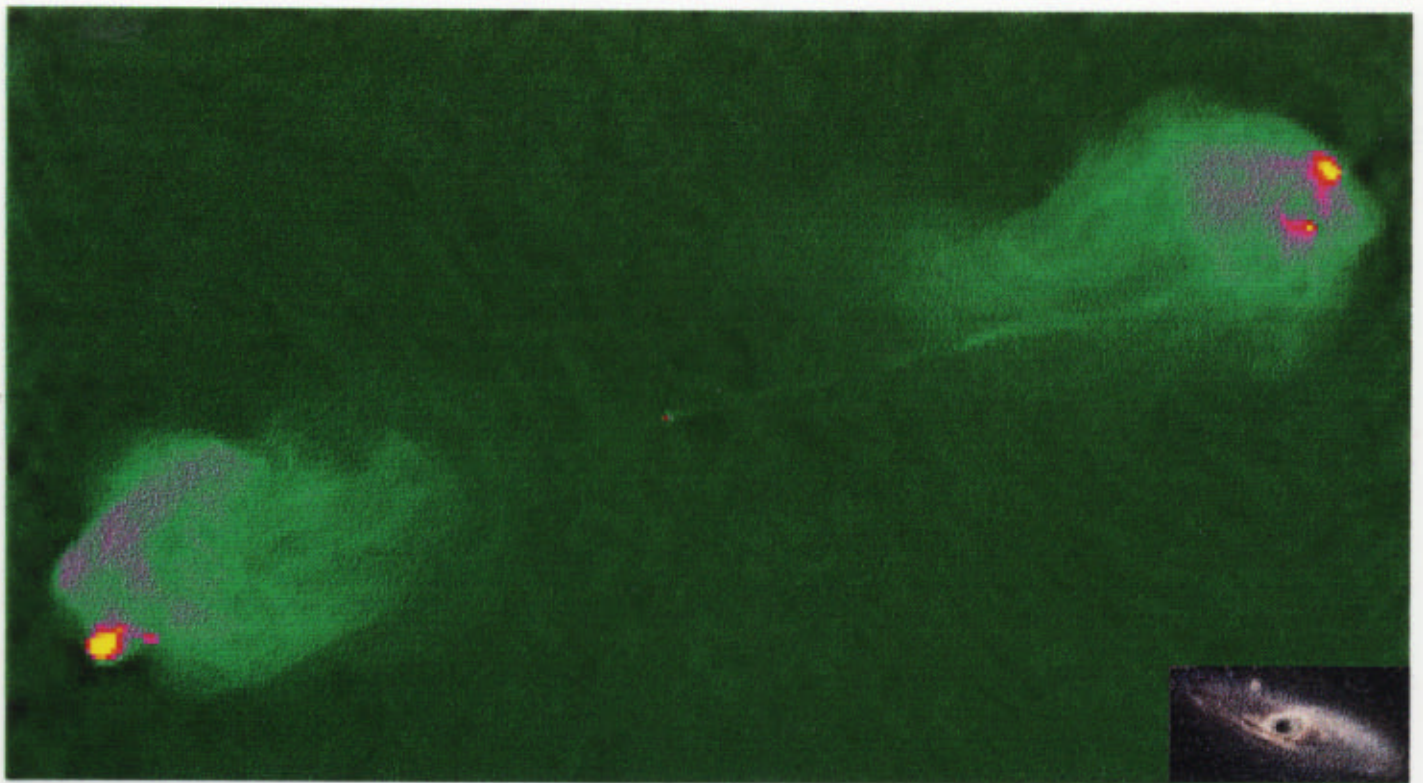
$$L \sim 10^{42} - 10^{48} \text{ ergs/sec}$$

$$L_{\text{galaxy}} \sim 10^{44} \text{ ergs/sec.}$$



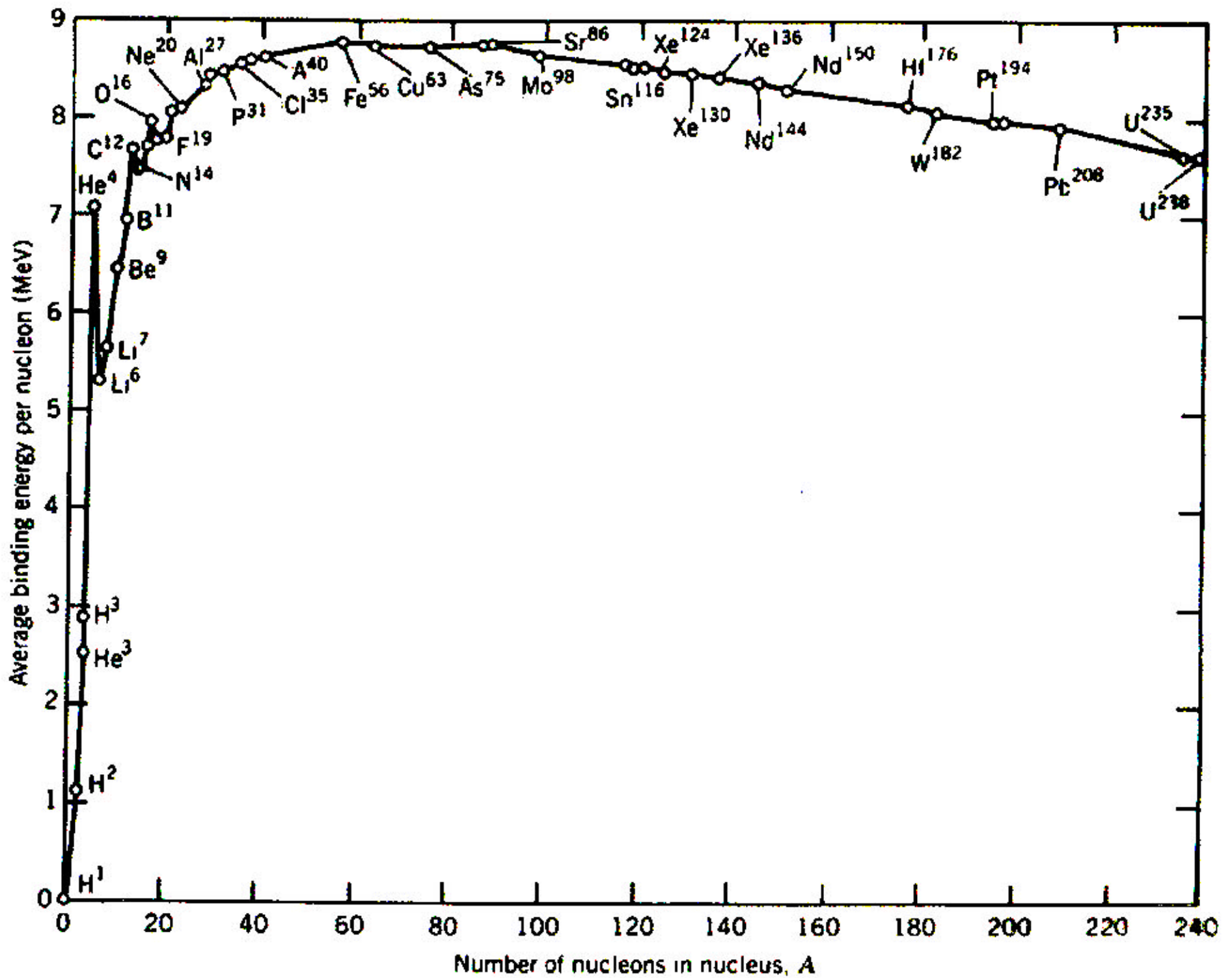
# ACTIVE GALACTIC NUCLEI

$\frac{1}{100}$  - 10,000 TIMES BRIGHTER  
THAN ALL THE STARS IN THE  
GALAXY.



Cygnus A



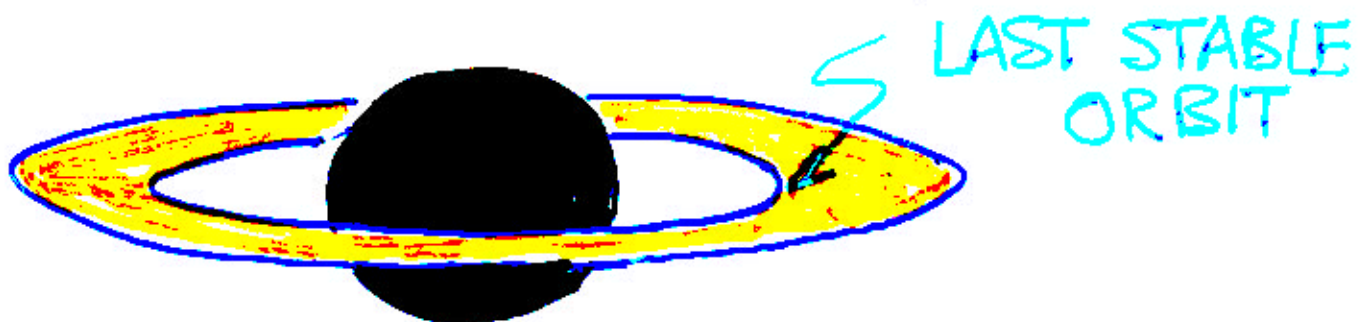


# GRAVITATIONAL BINDING VS THERMONUCLEAR FUSION

● FUSION  $4H^1 \rightarrow He^4 + (\text{ENERGY OUT})$

$$\frac{(\text{ENERGY OUT})}{(\text{REST ENERGY IN})} \approx 1\%$$

● GRAVITATIONAL BINDING



NEUTRON STAR:

$$\frac{(\text{ENERGY OUT})}{(\text{REST ENERGY IN})} \sim 6\%$$

EXTREME ROTATING BLACK HOLE:

$$\frac{(\text{ENERGY OUT})}{(\text{REST ENERGY IN})} \sim 30\%$$

THE BLACK HOLES OF  
GENERAL RELATIVITY (1915)  
ARE THE MOST EFFICIENT  
WAY OF REALIZING

$$E = mc^2$$

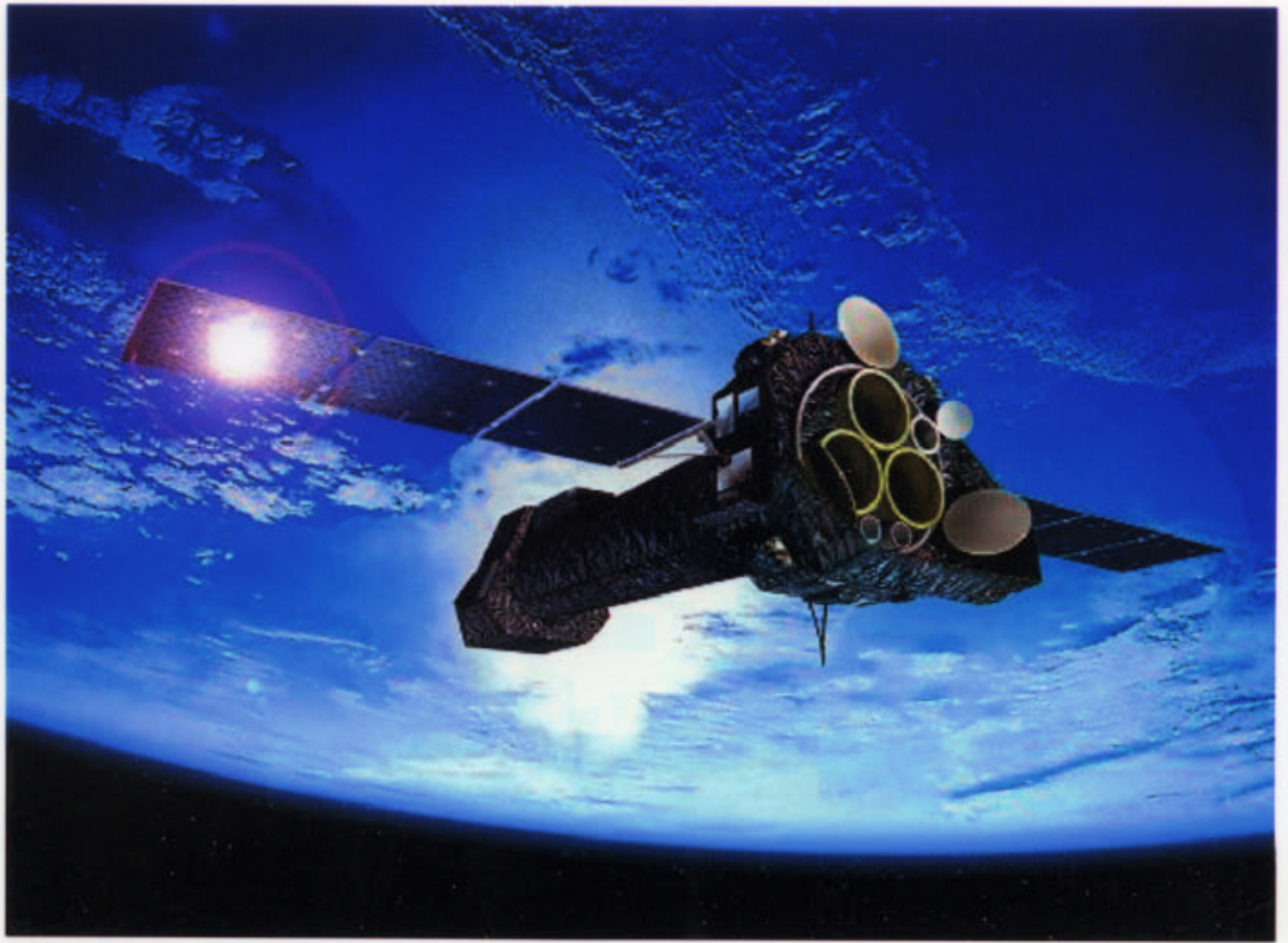
OF SPECIAL RELATIVITY (1905)



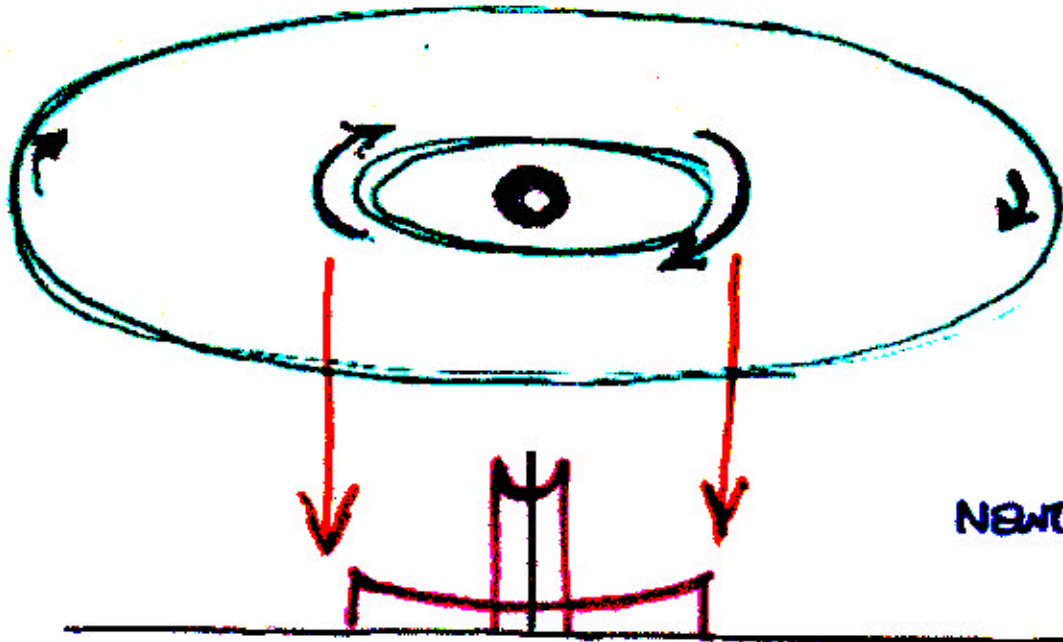
# FUTURE OF BLACK HOLE PHYSICS

- TODAY: IDENTIFIED PLACES WHERE THERE IS MUCH MASS IN A SMALL VOLUME,
- FUTURE: CHECK THE DETAILED PREDICTION OF EINSTEIN'S THEORY FOR THE GEOMETRY AROUND BLACK HOLES.

X-Ray  
GRAVITATIONAL WAVES



XMM



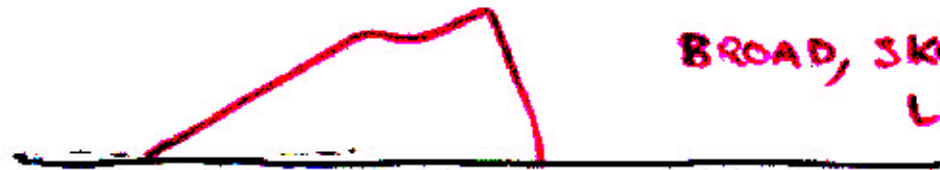
NEWTONIAN



SPECIAL  
RELATIVITY



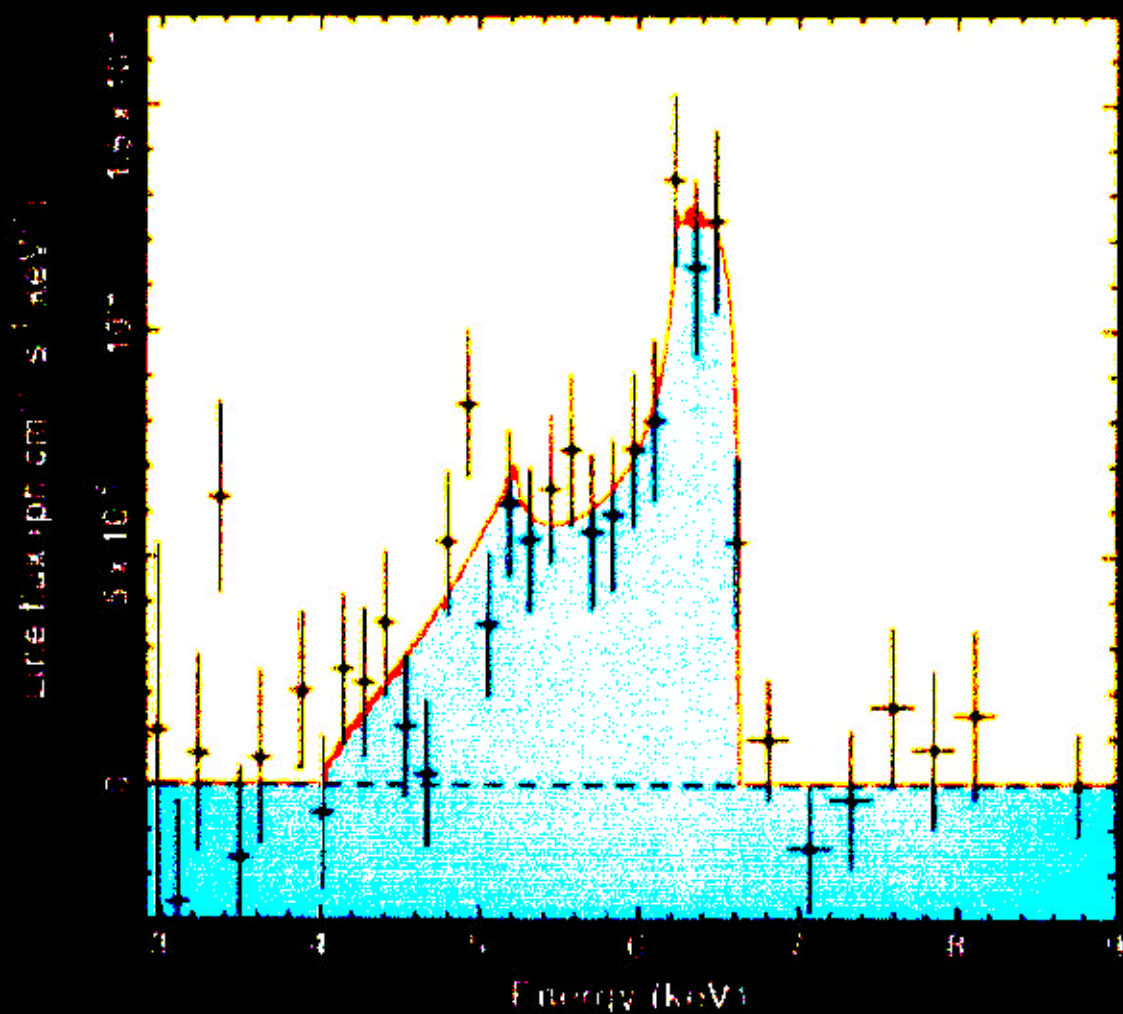
GENERAL  
RELATIVITY



BROAD, SKEW  
LINE

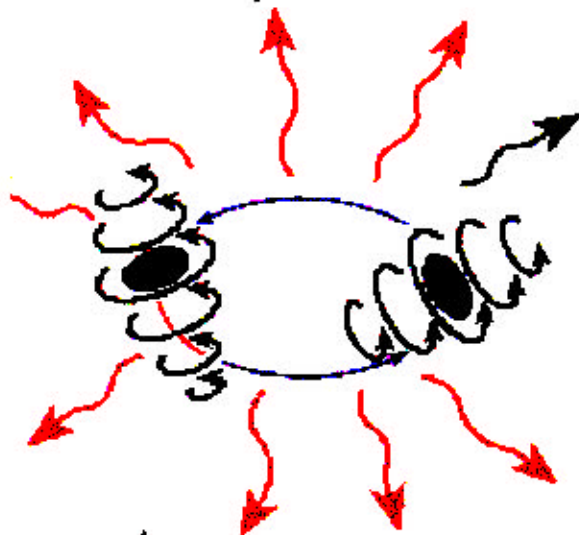


# Fe Line in MCG-6-30-15

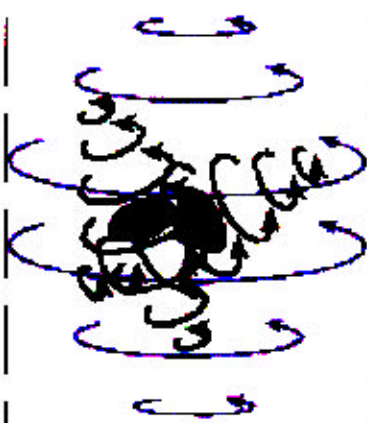


Tanaka, et al. (1995)

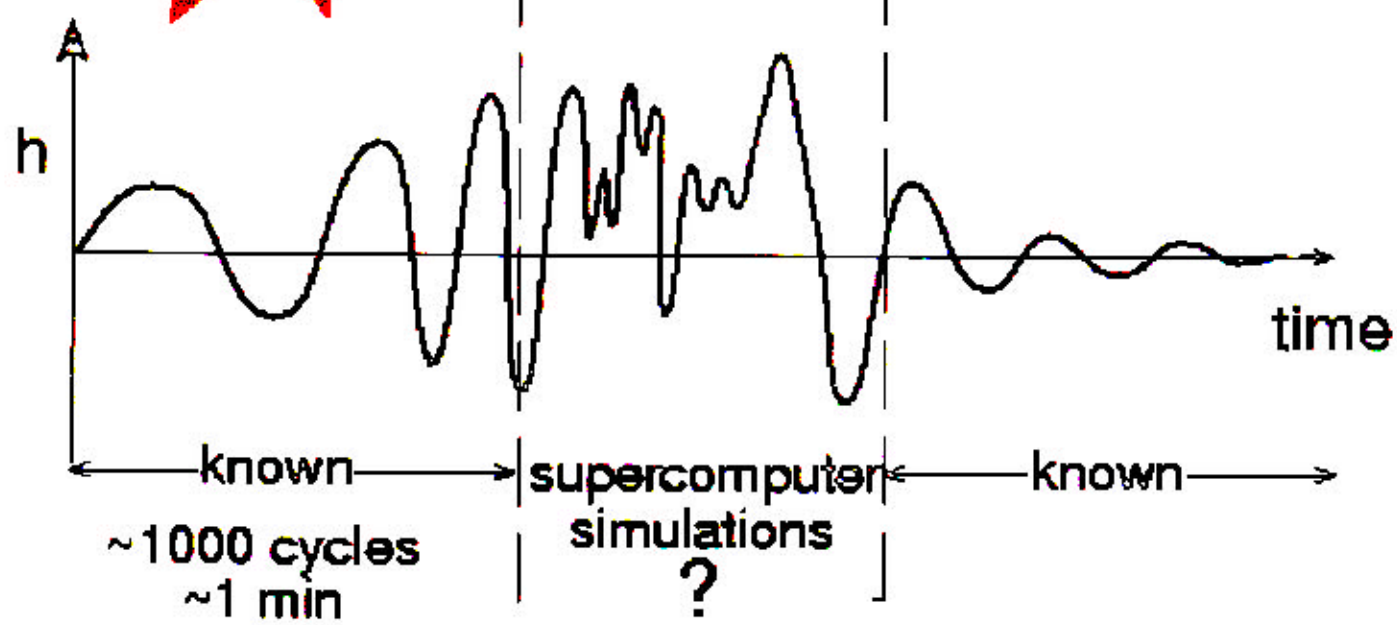
Inspiral



Merger



Ringdown



# COSMOLOGY

EINSTEIN'S THEORY PREDICTS

THE EXPANSION OF

THE UNIVERSE

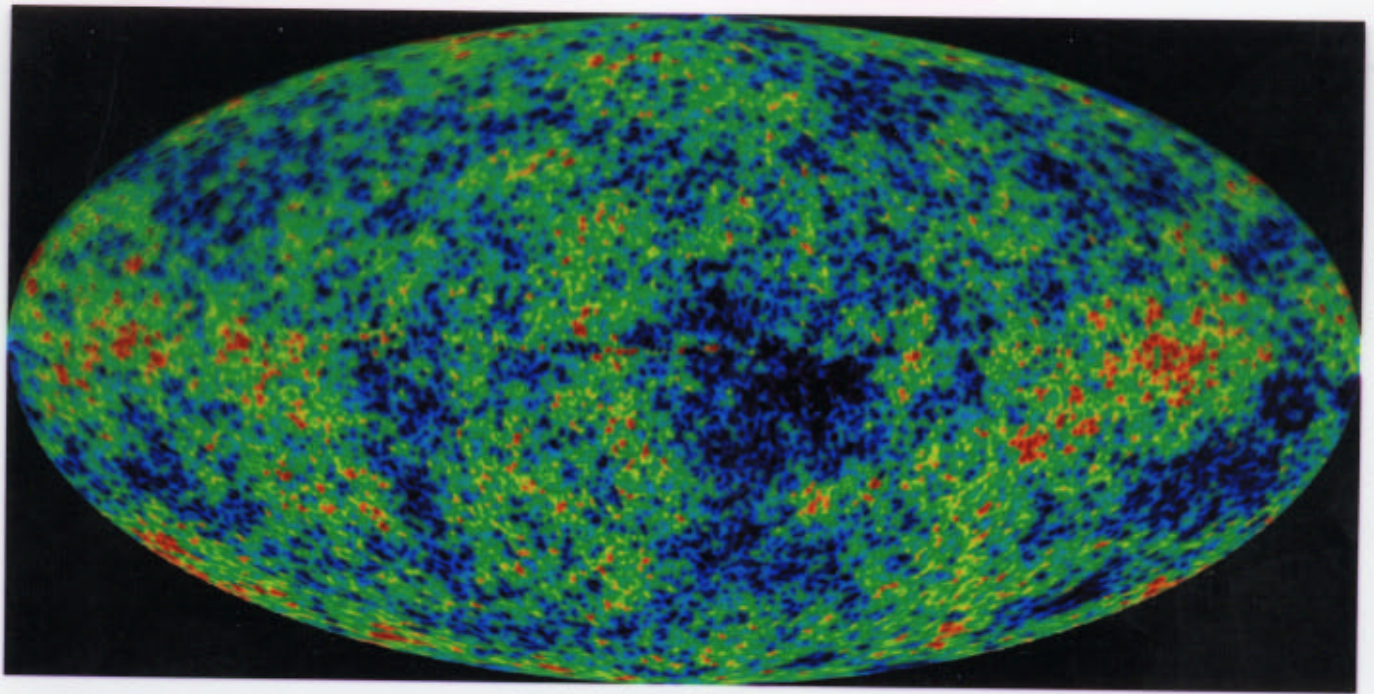
FROM A BIG BANG.



# GRAVITATIONAL ISSUES IN COSMOLOGY

- WHAT IS THE GEOMETRY OF THE UNIVERSE?
- WHAT IS THE SOURCE OF CURVATURE?
  - DARK MATTER
  - VACUUM ENERGY  
(COSMOLOGICAL CONSTANT.)
- DOES THE UNIVERSE OBEY THE EINSTEIN EQUATION?

# COSMOLOGY TODAY



## WMAP

- HOMOGENEOUS, ISOTROPIC ( $d \gtrsim 100 \text{ Mpc}$ )
- INITIAL MATTER THERMAL EQUILIBRIUM

## FRW MODELS

$$H_0 = 71 \text{ (km/s)/Mpc}$$

$$\Omega_{\text{VACUUM}} = .73$$

$$\Omega_{\text{RADIATION}} = 8 \times 10^{-5}$$

$$\Omega_{\text{BARYON}} = .04$$

$$\Omega_{\text{DARK MATTER}} = .23$$

COSMOLOGICAL  
PARAMETERS.



# COSMOLOGY TOMORROW

$$R_{\mu\nu} - \frac{1}{2}g_{\mu\nu}R = 8\pi G T_{\mu\nu}$$

IS EINSTEIN'S  
GRAVITATIONAL  
THEORY CORRECT  
ON THE  
SCALES OF  
COSMOLOGY?

$$\Omega_V = .73$$

$$\Omega_r = 8 \times 10^{-5}$$

$$\Omega_b = .04$$

$$\Omega_{DM} = .23$$

96% OF UNCERTAIN  
CHARACTER



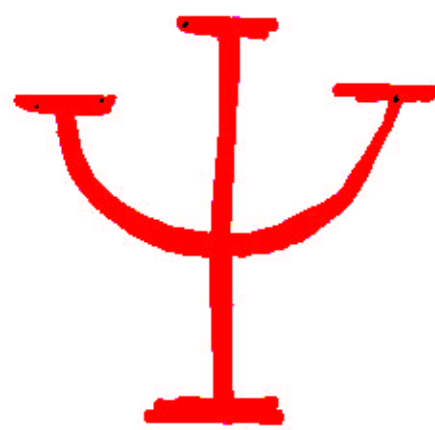


# COSMOLOGY AND QUANTUM GRAVITY

- AT THE BIG BANG  
PLANCK ENERGIES  
ARE REACHED  
CHARACTERISTIC OF QUANTUM  
FLUCTUATIONS IN GEOMETRY  
AND THE CONJECTURED SCALE  
OF THE UNIFICATION OF ALL FORCES.

$$E_p = \left( \frac{c^5}{\hbar G} \right)^{1/2}$$

- AT THE BIG BANG  
WE ARE CLOSEST  
TO THE QUANTUM  
INITIAL CONDITION  
OF THE UNIVERSE.



- AT THE BIG BANG, LARGE  
AND SMALL ARE ONE.

# THIS DECADE AND THE NEXT

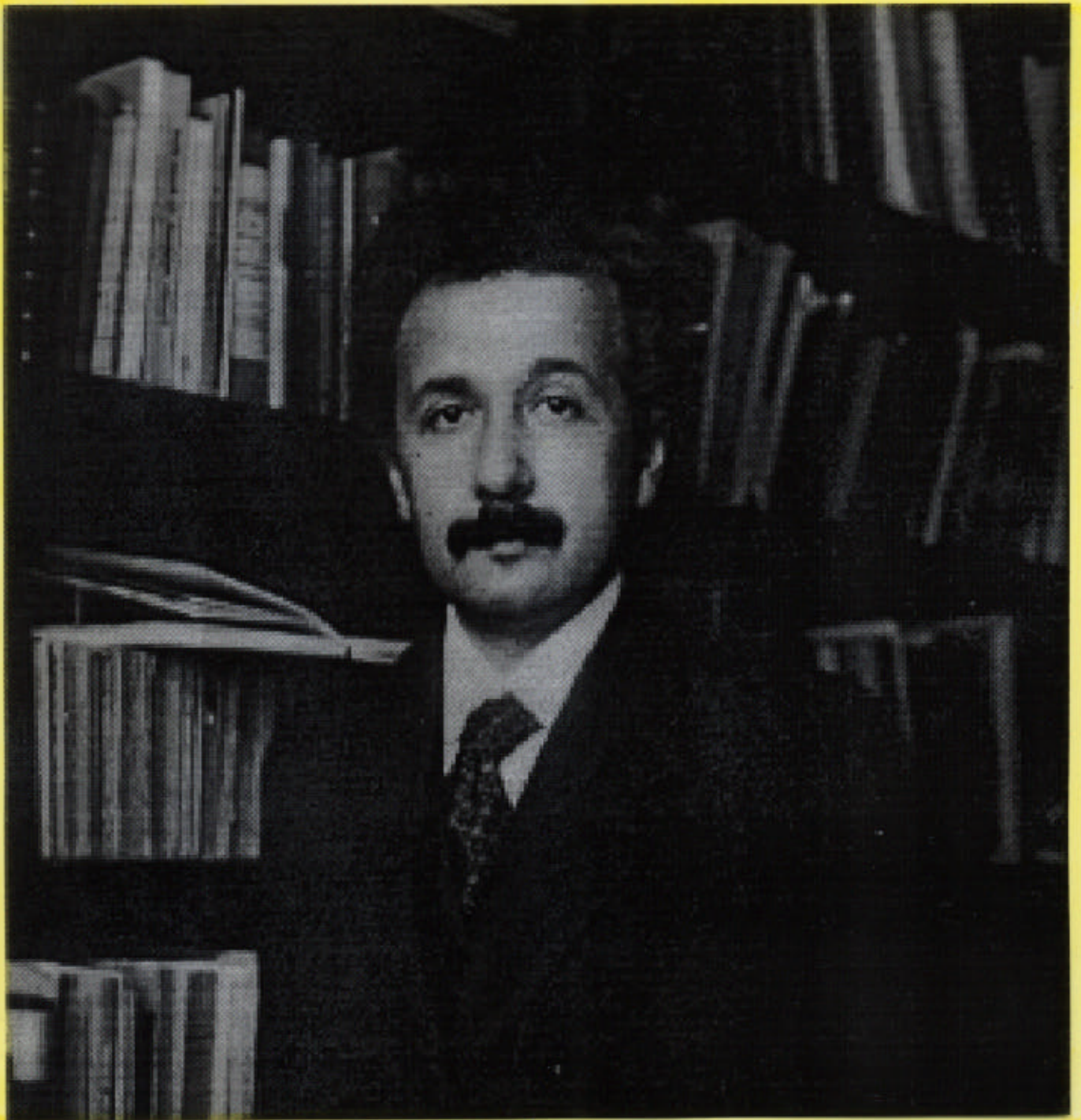
- THE DIRECT DETECTION OF GRAVITATIONAL WAVES AND THEIR APPLICATION TO PROBE REGIONS OF STRONG GRAVITY
- VERIFICATION OF THE DETAILED PREDICTIONS OF GR FOR THE GEOMETRY OUTSIDE BLACK HOLES.
- TESTS OF GENERAL RELATIVITY ON COSMOLOGICAL SCALES
- EXPLORATION OF THE LIMITS OF GENERAL RELATIVITY AND NEW PRECISION TESTS.
- OBSERVATIONAL SIGNATURES OF QUANTUM GRAVITY.



A THEORY IS MORE IMPRESSIVE  
THE GREATER THE SIMPLICITY  
OF ITS PREMISES,  
THE MORE DIFFERENT  
KINDS OF THINGS IT RELATES,  
AND THE MORE EXTENDED  
ITS AREA OF APPLICABILITY.

A. EINSTEIN

ALBERT EINSTEIN



1916