

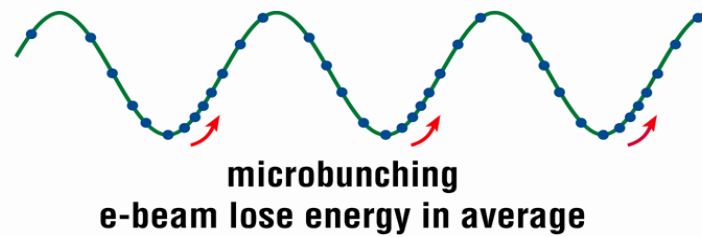
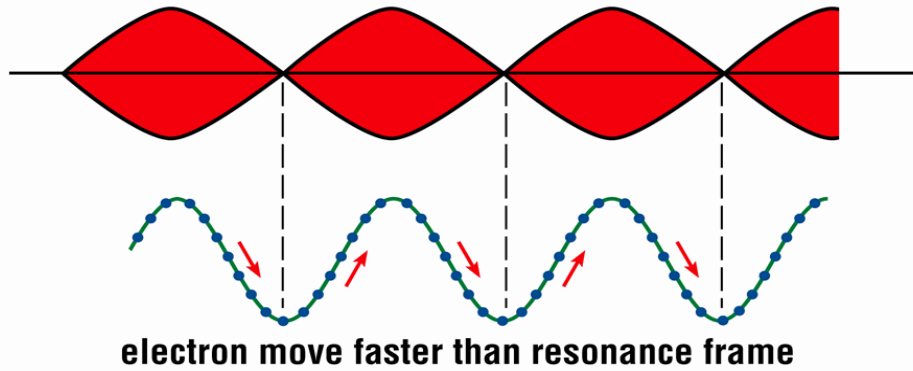
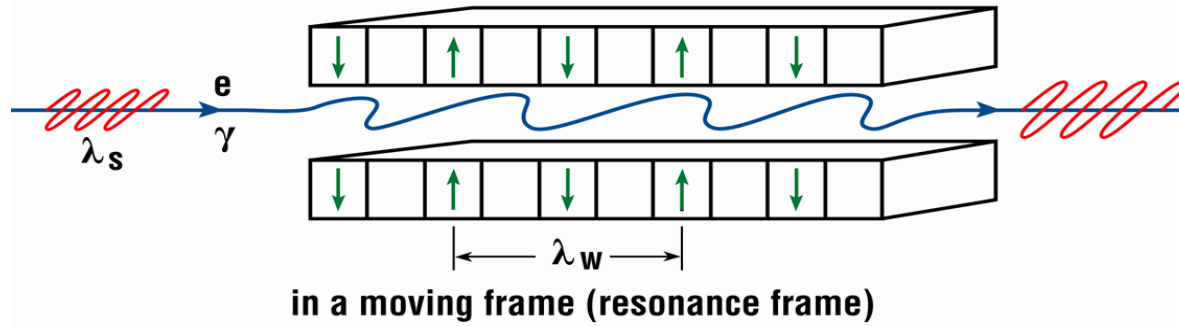
# R&D Towards X-ray Free Electron Laser

Li Hua Yu  
Brookhaven National Laboratory

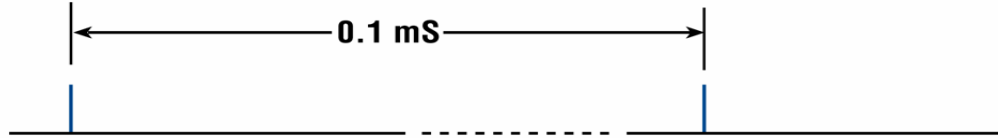
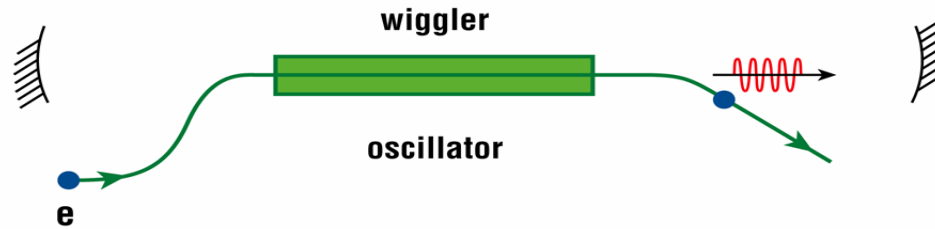
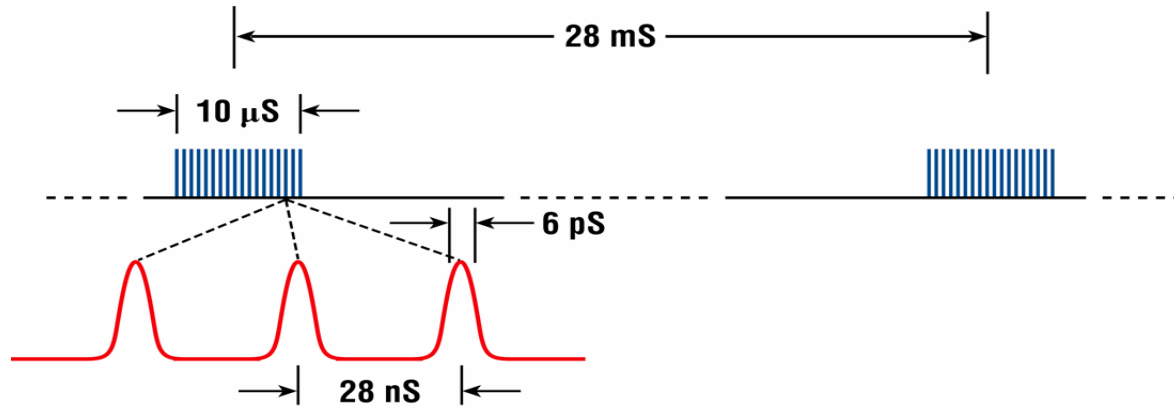
1/23/2004

# FEL Principle

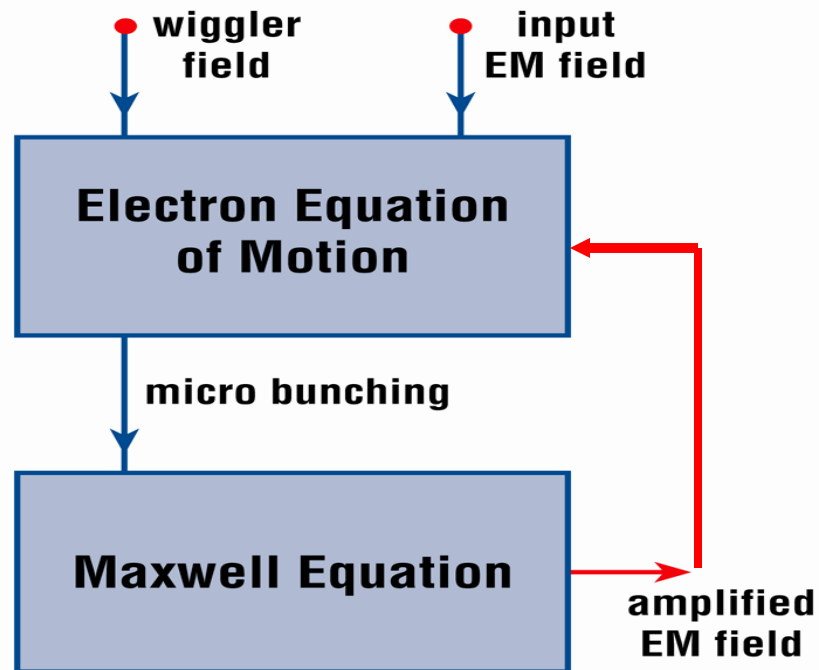
J. Maday *et. al.* (1976)



# Time Structure and Different FEL Configurations



## High Gain FEL – Exponential Growth



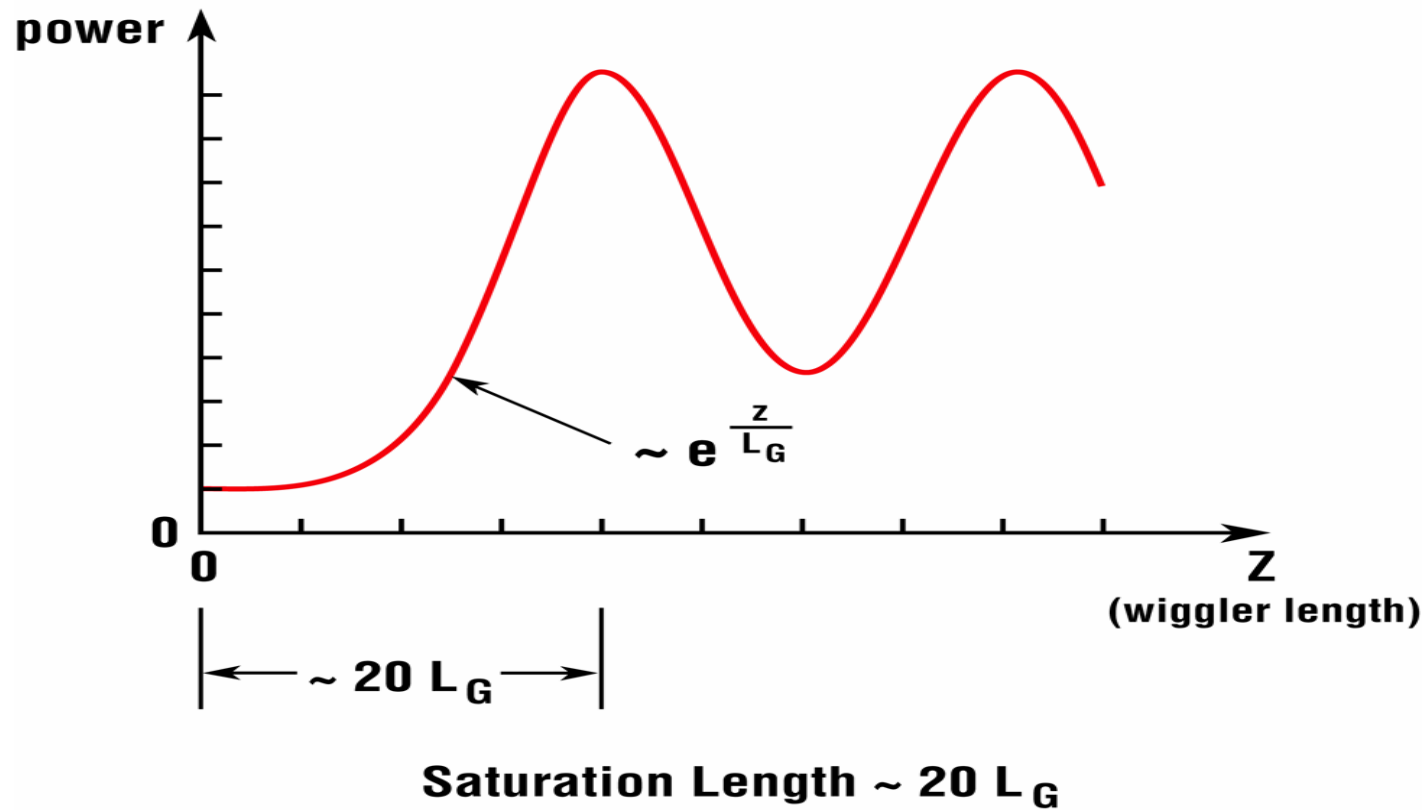
**1D Theory: exponential growth, proposes **SASE****  
**Self Amplified Spontaneous Emission**

Saldin et al. (1982)

Bonifacio, Pellegrini et al. (1984)

# Gain Length

(Length power increases by a factor of e)



# Beam quality requirement: Scaling Function of Gain

L.H.Yu, S.Krinsky, R. Gluckstern PRL. 64, 3011 (1990)

$$\frac{\lambda_w}{L_G D} = G \left( \frac{\varepsilon_n}{\gamma \lambda_s}, \frac{\sigma_\gamma}{\gamma D}, \frac{\lambda_w}{\lambda_\beta D} \right)$$

Electron beam energy  $\gamma$

Current  $I_0$

Wavelength:  $\lambda_s = \frac{\lambda_w}{2\gamma^2} (1 + a_w^2) \quad (a_w \sim 1)$

Emittance  $\varepsilon$  (transverse phase space size = beam size\*angular spread)

$$\gamma \propto 1 / \sqrt{\lambda_s} \quad , \quad I_0 \propto 1 / \sqrt{\lambda_s} \quad ,$$

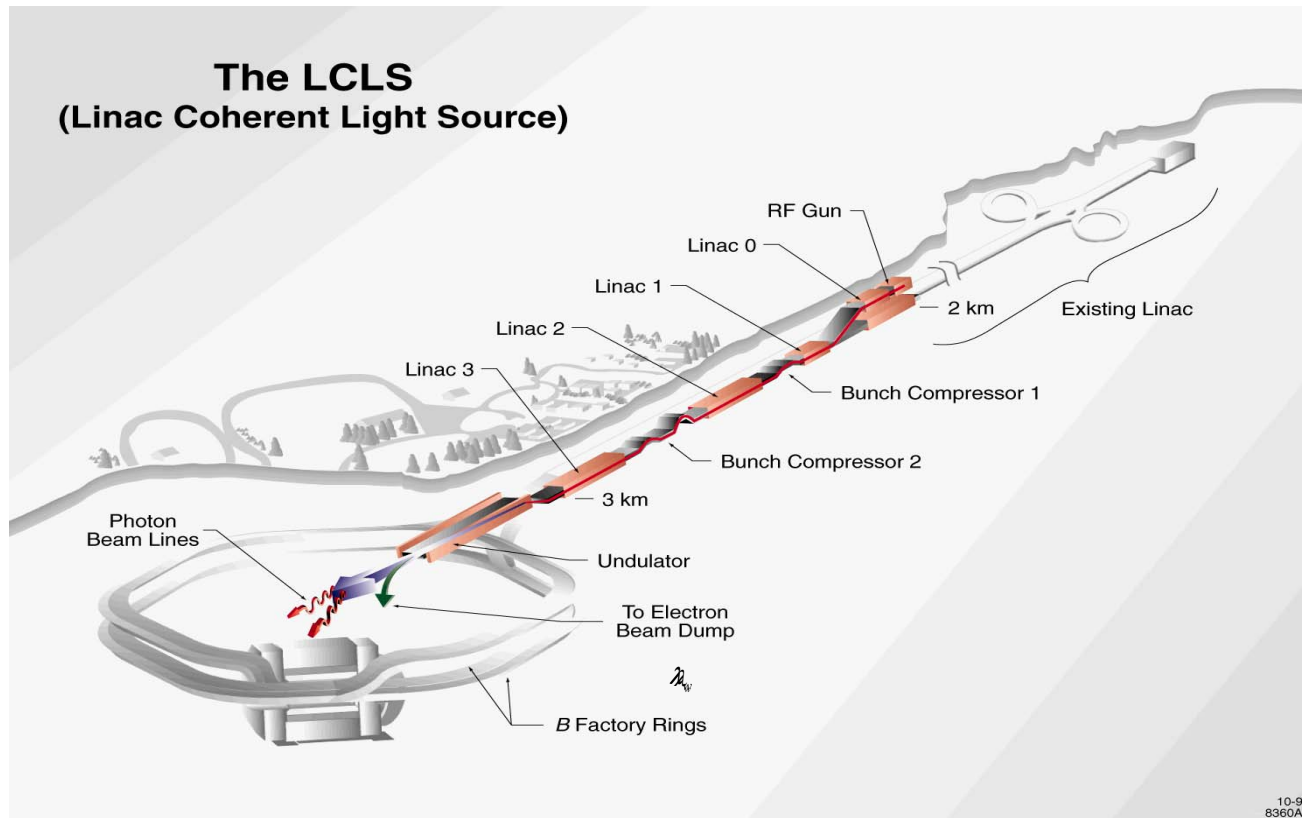
Scaling:

$$\varepsilon_n \propto \sqrt{\lambda_s}$$

**The weak square root scaling is favorable for  
going to short wavelength**

**The development of Photocathode RF Gun (R. Schefield, 1985) with Emittance Compensation Solenoid (B. Carleston) makes it possible to consider high gain X-ray FEL because it provides high brightness beam**





## LCLS (SLAC) Proposal

Wavelength	$\lambda$	1.5 Å
Electron energy	$\gamma$	14.35 GeV
Norm. emittance (rms)	$\varepsilon_n$	1.5 $\pi$ mm mrad
Peak current	$I_0$	3,400 A
Wiggler Period	$\lambda_w$	3 cm
Gain Length	$L_G$	5.8 m

Courtesy: Max  
Cornaccia



Linac Coherent Light Source

Stanford Linear Accelerator Center

Stanford Synchrotron Radiation Laboratory

*& Schedule*

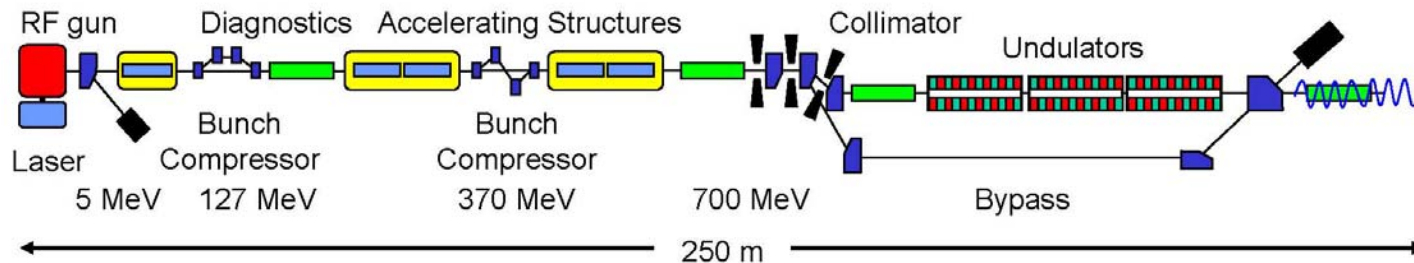
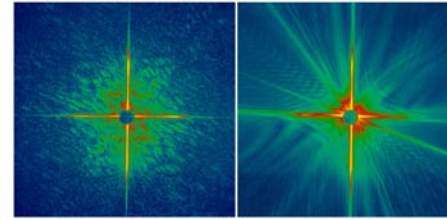
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**Schedule:**

1999 - 2002	Research and Development
2003 - 2006	Project Engineering Design
2005	Long-Lead Procurement
2006 - 2008	Facilities Construction, Startup
2009	Start Operations

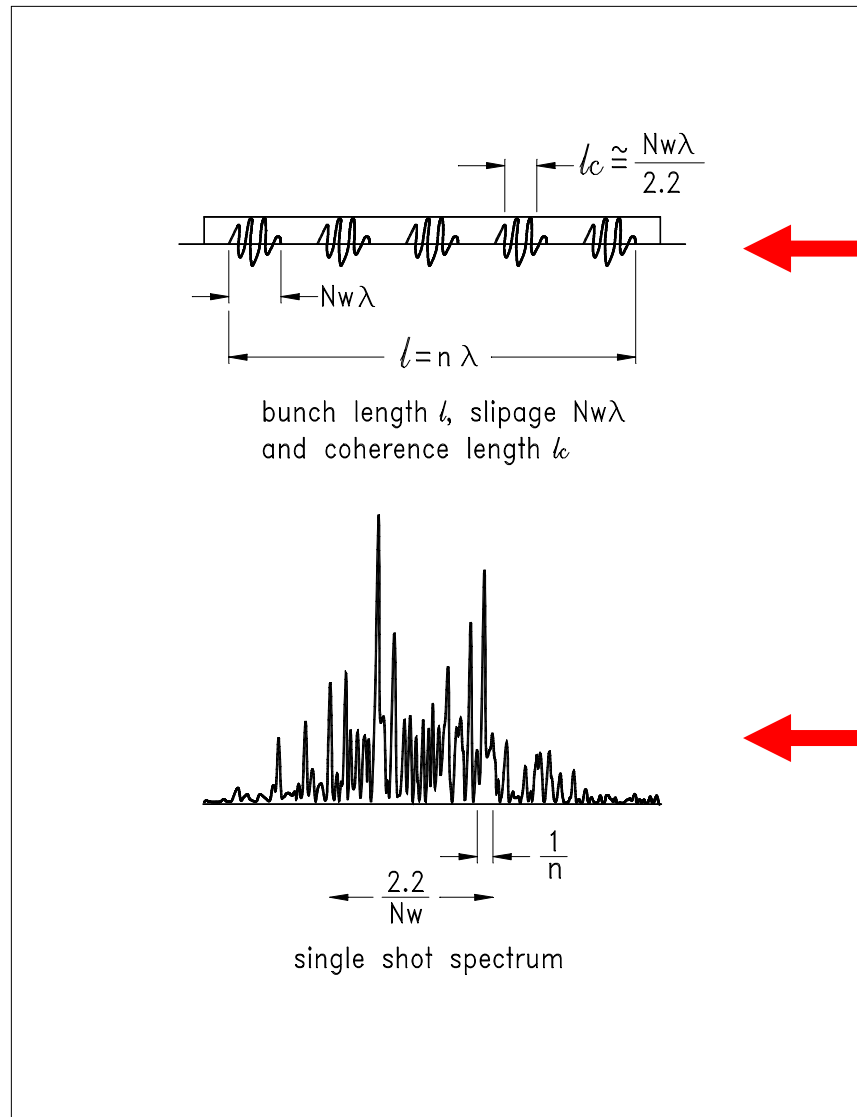
**Cost:**

The Linac Coherent Light Source makes use of the last 1/3 of the Two-Mile Linac at the Stanford Linear Accelerator Center, reducing the cost of construction by more than \$300,000,000.



**Free-electron laser at DESY delivers highest power at at wavelengths between 13.5 and 13.8 nanometers with an average power of 10 milliwatts and record energies of up to 170 microjoules per pulse.**

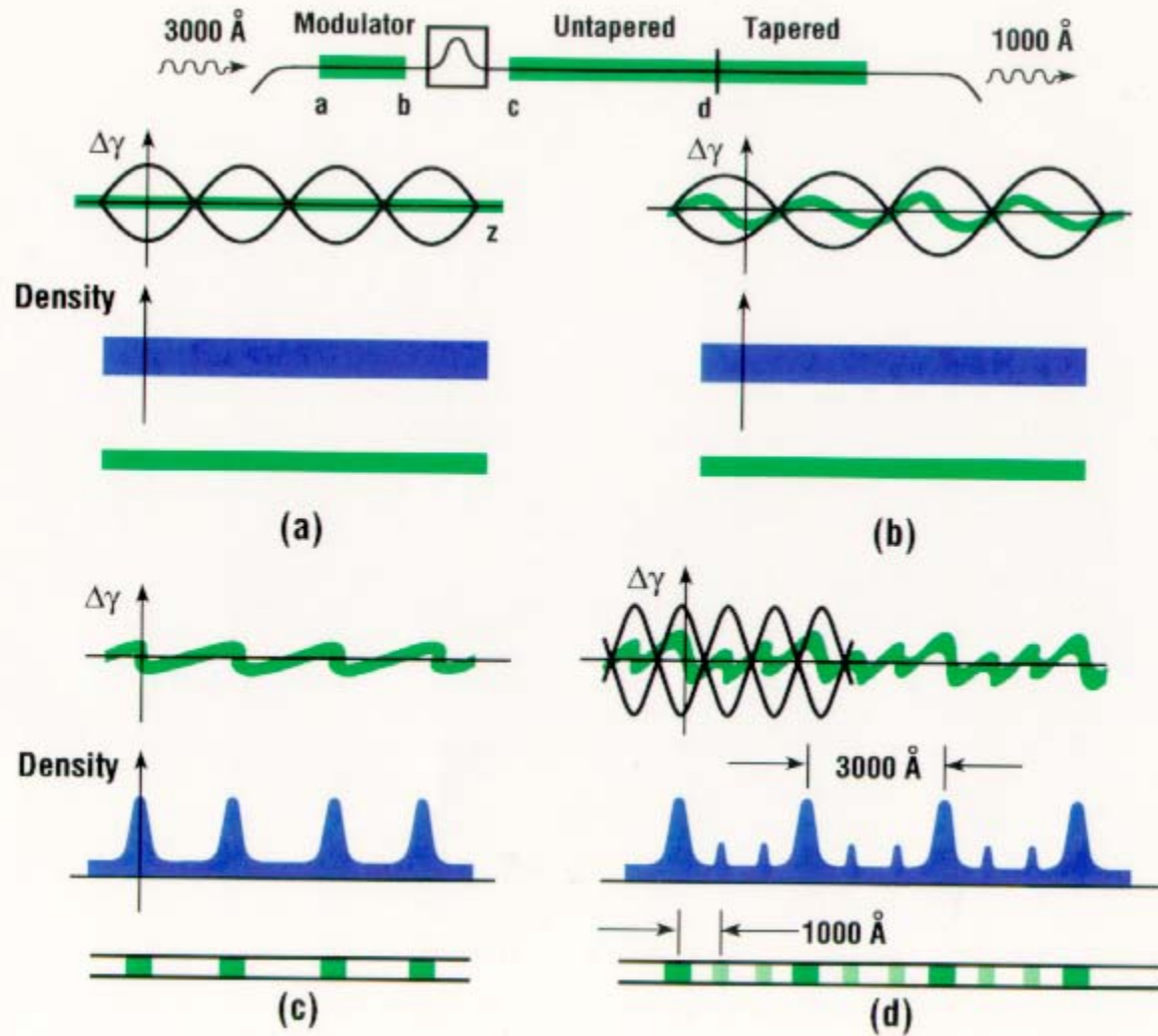
# SASE spectrum is not coherent



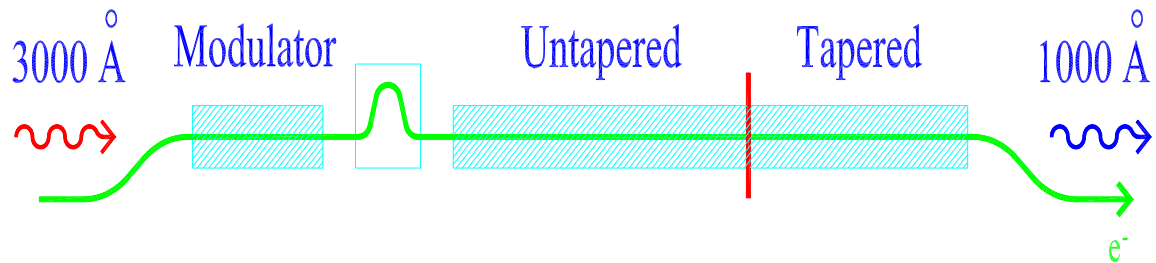
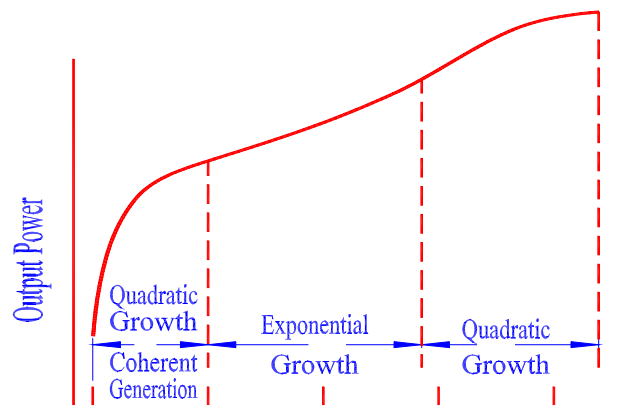
Train of pulses with independent phase

Spectrum consists spikes with 100% fluctuation

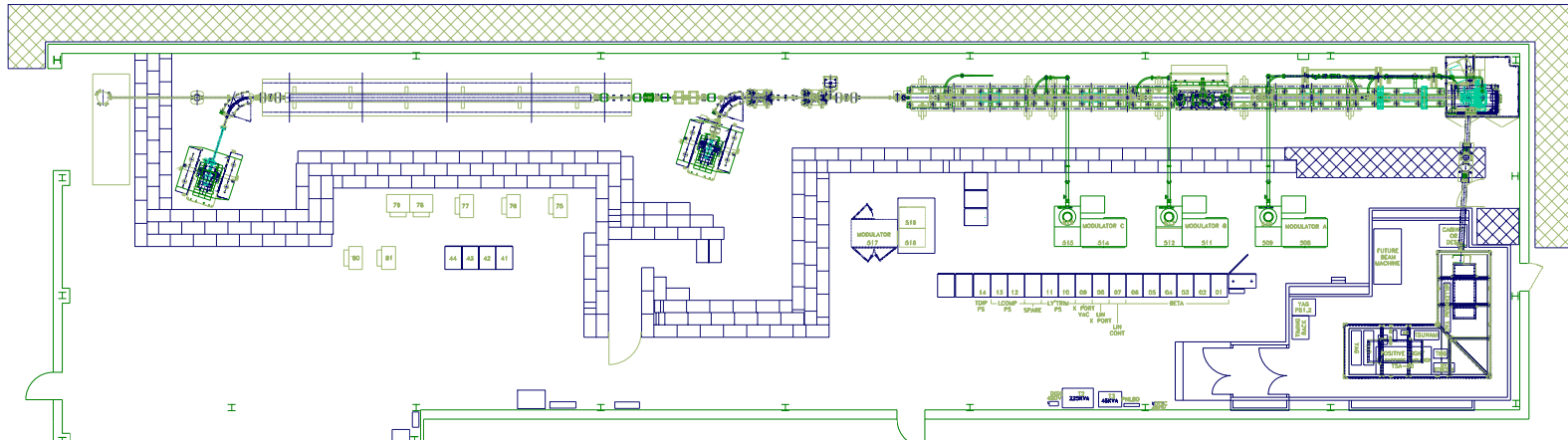
# High Gain Harmonic Generation (HGHG) for coherence and stability



# HIGH GAIN HARMONIC GENERATION (HGHG)

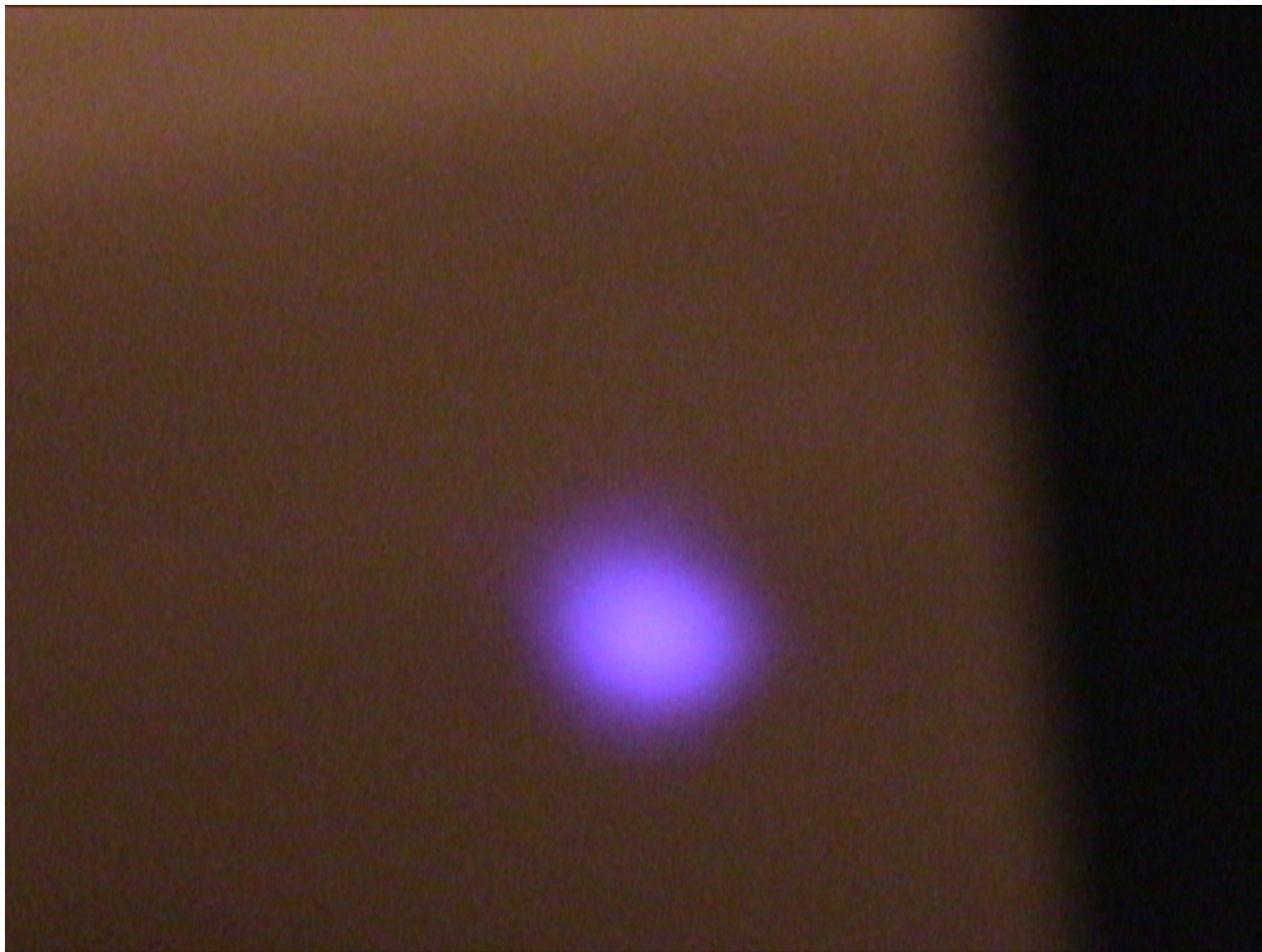


# DUVFEL Configuration

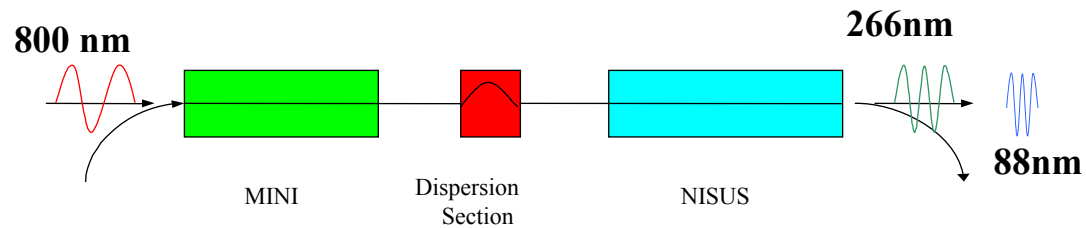


DUVFEL using NISUS wiggler  
Step 1. SASE at 400 nm  
Step 2. direct seeding at 266nm  
Step 3. HGHG 800nm  $\rightarrow$  266 nm

Camera image after exit window



# HGHG Experiment

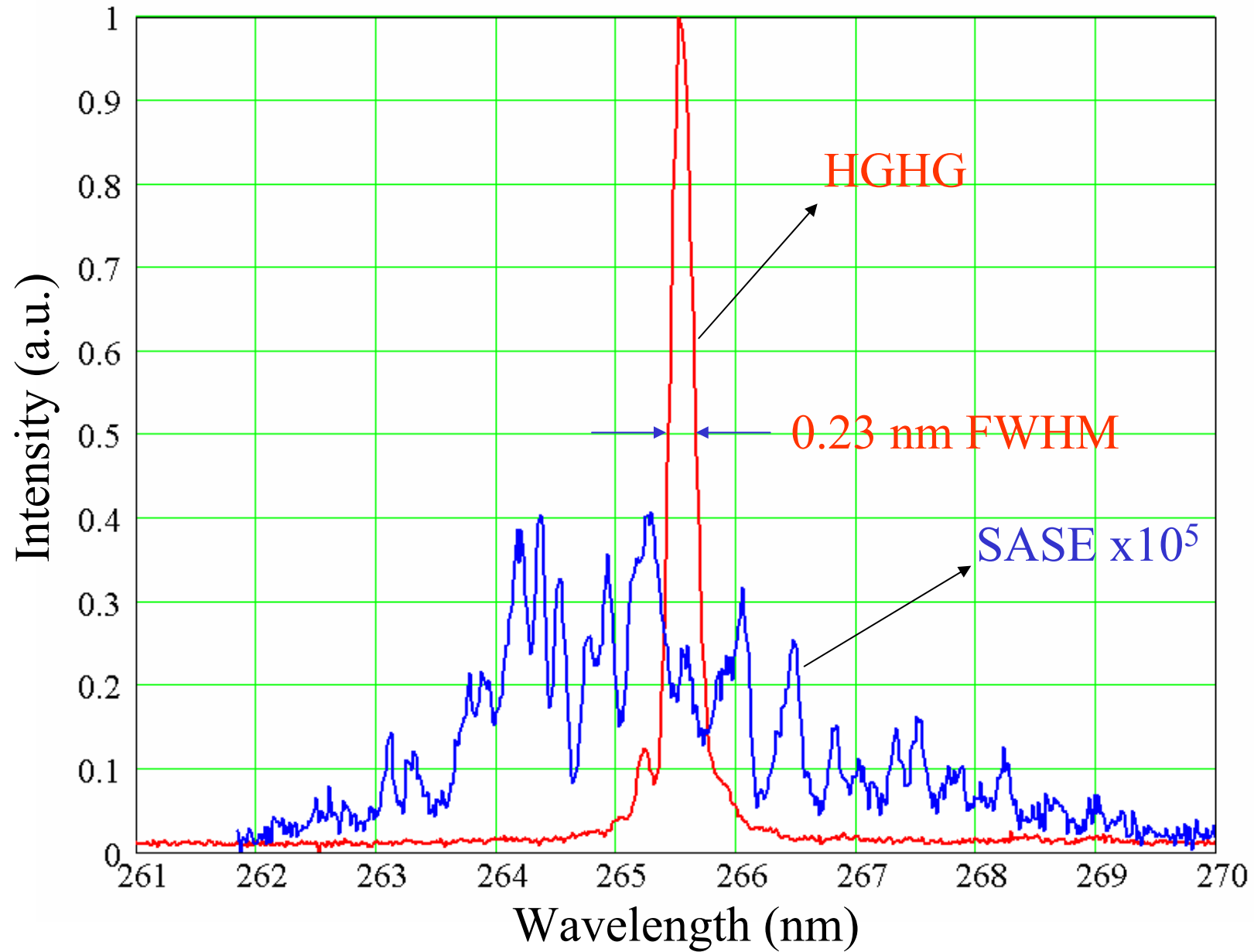


**HGHG from 800 nm → 266 nm,  
Output at 266 nm: ~ 120 μJ,  
e-beam: 300 Amp, 3 mm-mrad,  
Energy 176 MeV**

**Output at 88 nm ~ 1 μJ  
On-going Experiment Application in Chemistry**

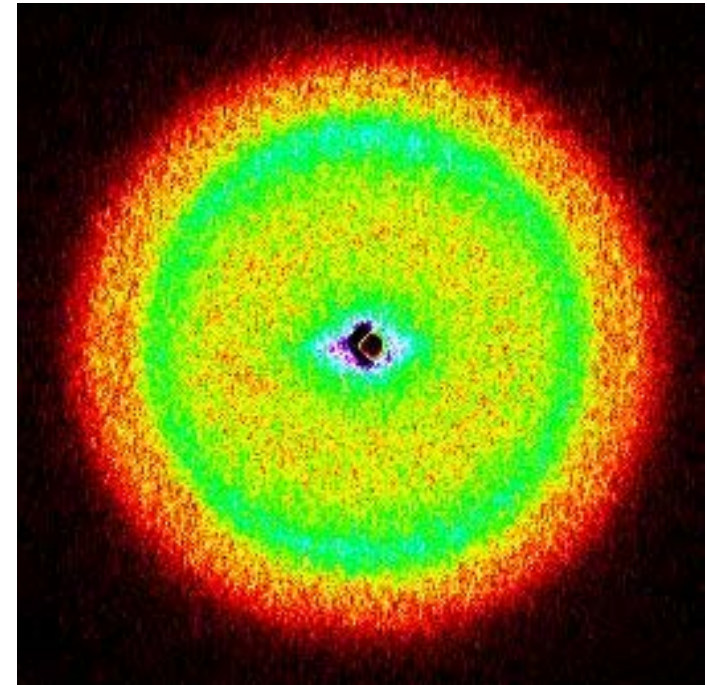
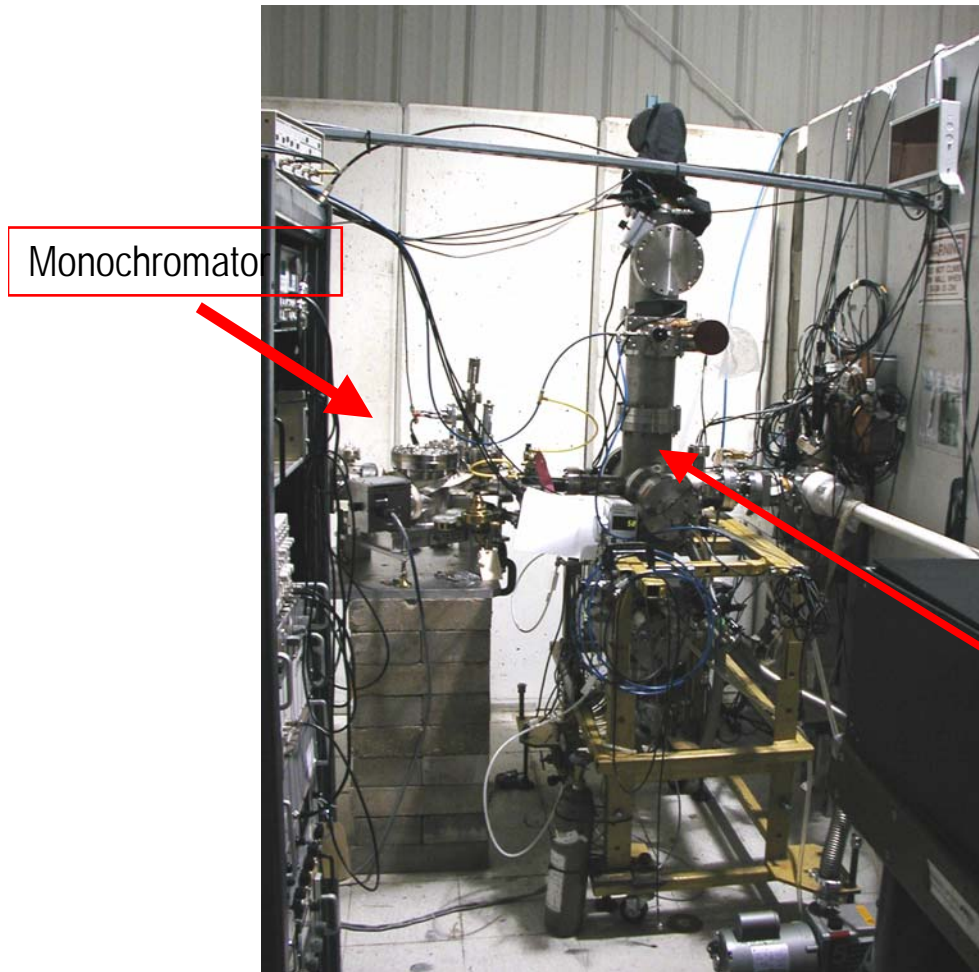
L.H. Yu et al **PRL** 91 074801-1 (2003)

# Spectrum of HGHG and SASE at 266 nm under the same electron beam condition



# Harmonic at 89 nm is used in Ion pair imaging experiment

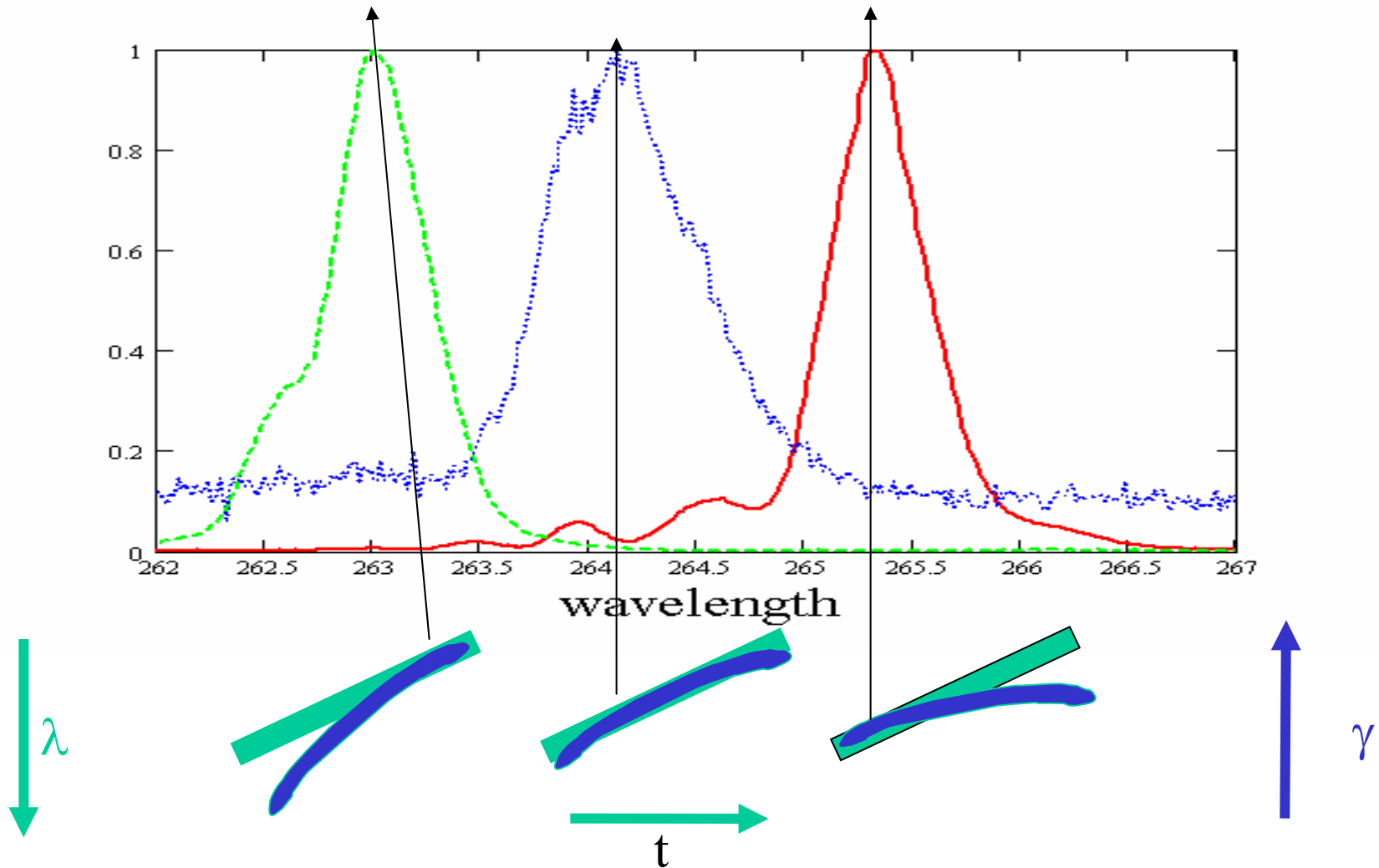
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Ion Pair Imaging station

Brookhaven Science Associates  
U.S. Department of Energy

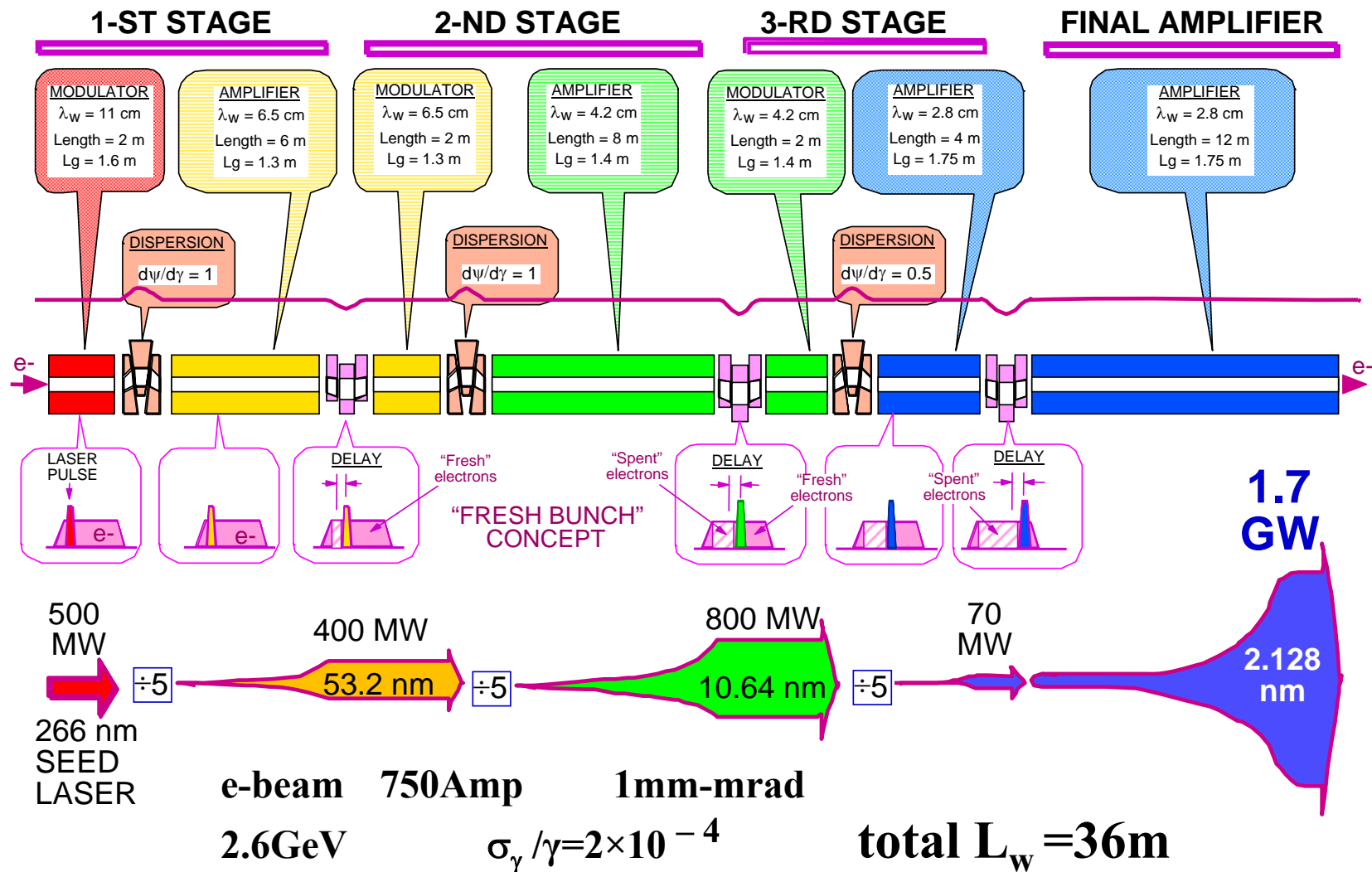
# CPA (Chirped Pulse Amplification) HGHG Spectra



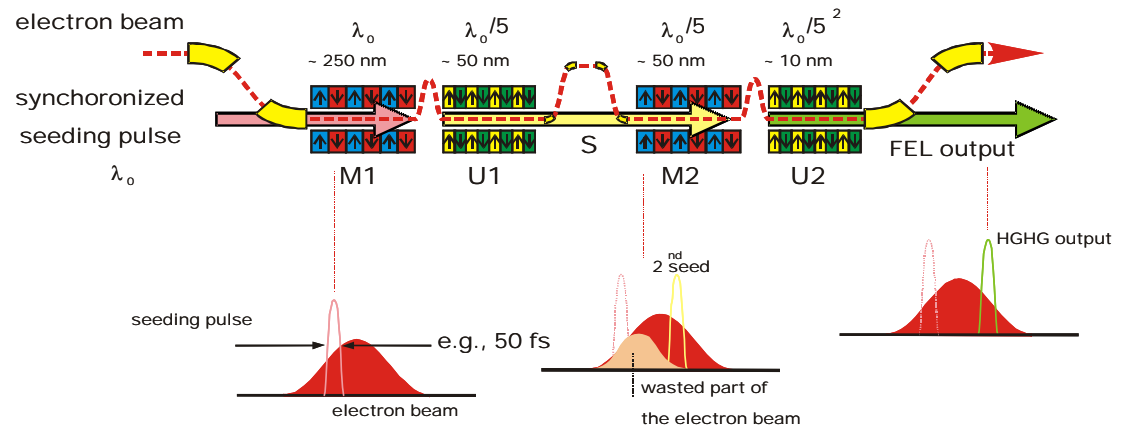
Interpretation of 3 sets of measurements: effect of RF curvature; the second matched e-beam chirp to seed chirp

# R&D Towards Cascading HGHG: One of the Earlier Calculated Schemes

## A Soft X-Ray Free-Electron Laser



# FEL at ELETTRA of Trieste

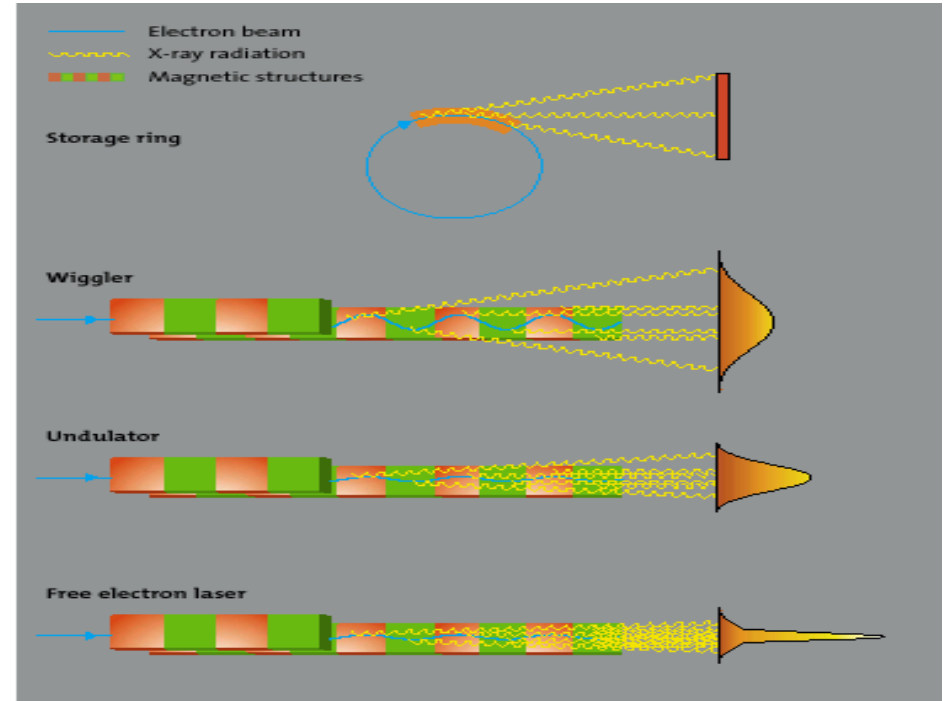
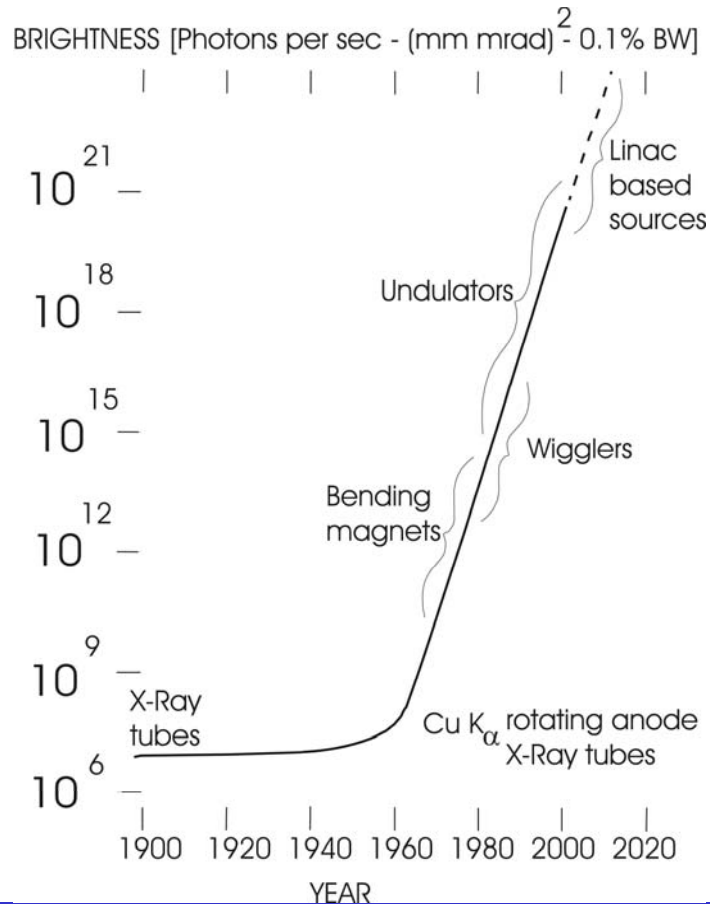


## FEL at 1 GeV

	FEL-1		FEL-2		
Wavelength Target	100	40	40	10	nm
	10	31	25	124	eV
Electron Beam Energy	0.70		0.55	1.00	GeV
Bunch Charge	1.0		1.0		nC
Peak Current	0.8		2.5		kA
Bunch Duration (rms)	500		160		fs
Energy Spread (rms)	0.5		1.0		MeV
Normalized Emittance	2.0		1.5		$\times 10^{-6}$ m
Undulator Period	52		36.6		mm



# Introduction



Coherent time:

$$\tau \approx \hbar / kT \approx 100 \text{ fs}$$

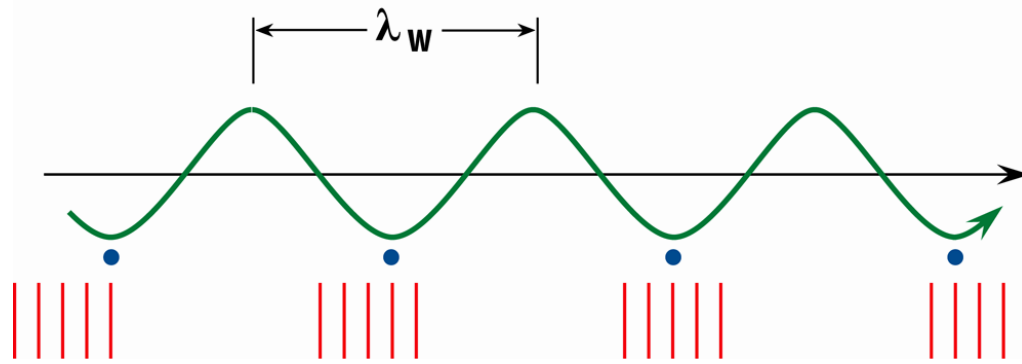
- Ultrahigh Spatial Resolution
- Microscopy
- Coherent X-ray Scattering

## Resonance Condition

$$(c - v_{\parallel}) \frac{\lambda_w}{v_{\parallel}} = \lambda_s$$

$$\frac{1 + a_w^2}{2\gamma_0^2} \lambda_w = \lambda_s$$

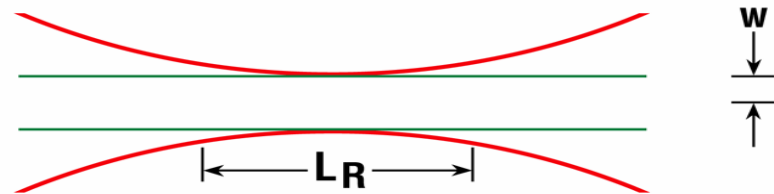
$$a_w = \frac{eA_w}{mC} \sim 1 \quad (\text{scaled vector potential})$$



# Diffraction and Optical Guiding

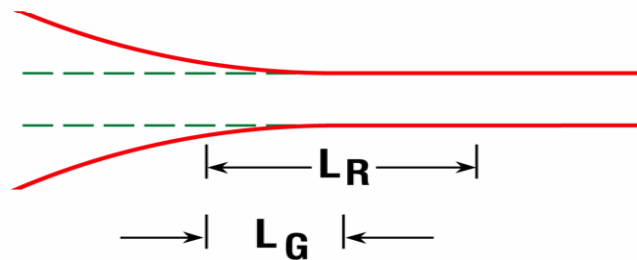
G. Moore, Nucl. Ins. Meth. A239, 119 (1985)

Scharleman, Sessler, Wurteli, PRL, 54, 1925 (1985)

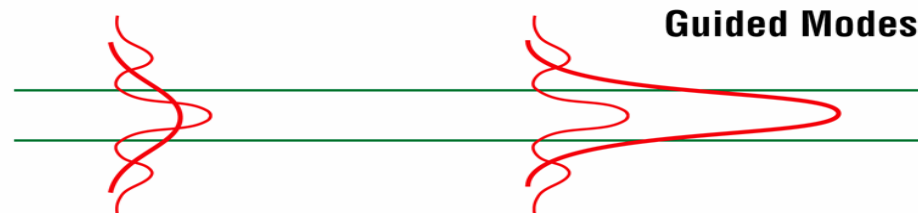


Rayleigh Range

$$L_R = \frac{\pi w^2}{\lambda_s}$$



$$\frac{L_R}{L_G^{1D}} \gg 1 \longrightarrow 1 - D \text{ Theory}$$



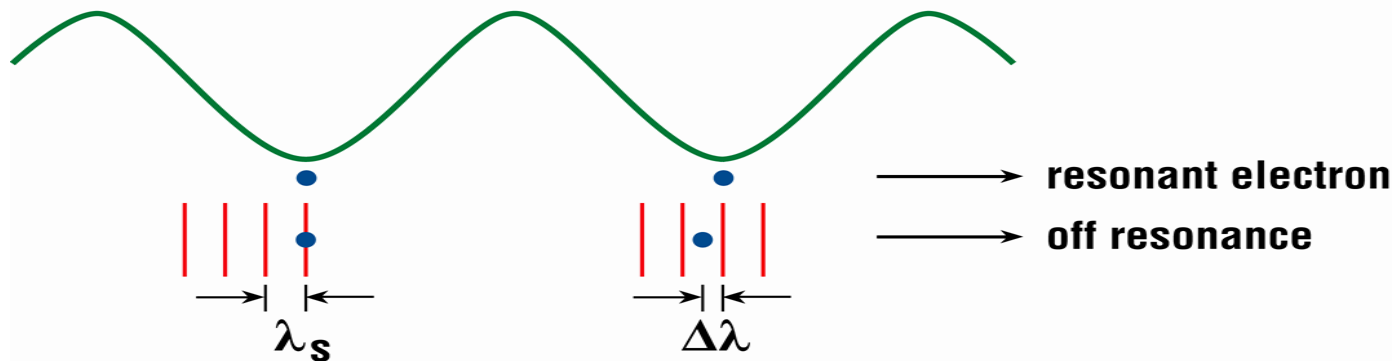
## Effects of Electron Beam Quality

***energy spread degrades coherent bunching***

$$\lambda_s = \frac{\lambda_w}{2\gamma_0^2} (1 + a_w^2)$$

$$\lambda = \frac{\lambda_w}{2\gamma^2} (1 + a_w^2)$$

$$\therefore \frac{\Delta \lambda}{\lambda} = 2 \frac{\Delta \gamma}{\gamma}$$



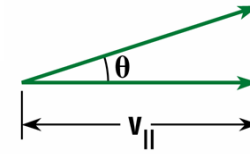
# Emittance

- Angular spread  $\square$  also degrades micro-bunching

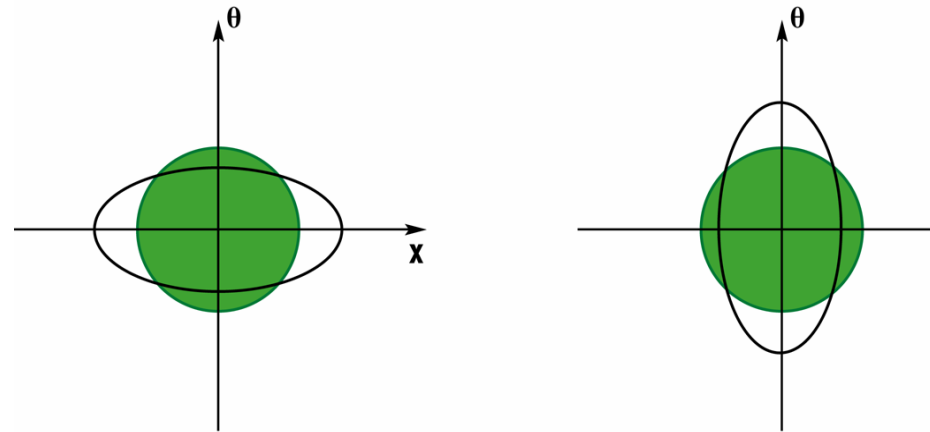
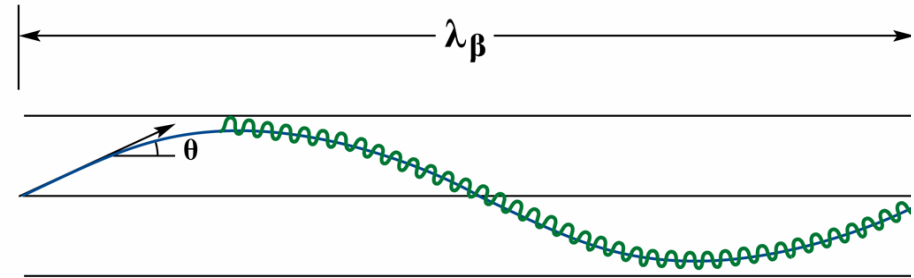
- Strong focusing reduces beam size but increases angular spread

- Emittance  $\mathcal{M}$   
 $\sim$  beam size  $\times$  angular spread

- Requirement on  $\mathcal{M}$ :  $\varepsilon < \frac{\lambda_s}{4\pi}$

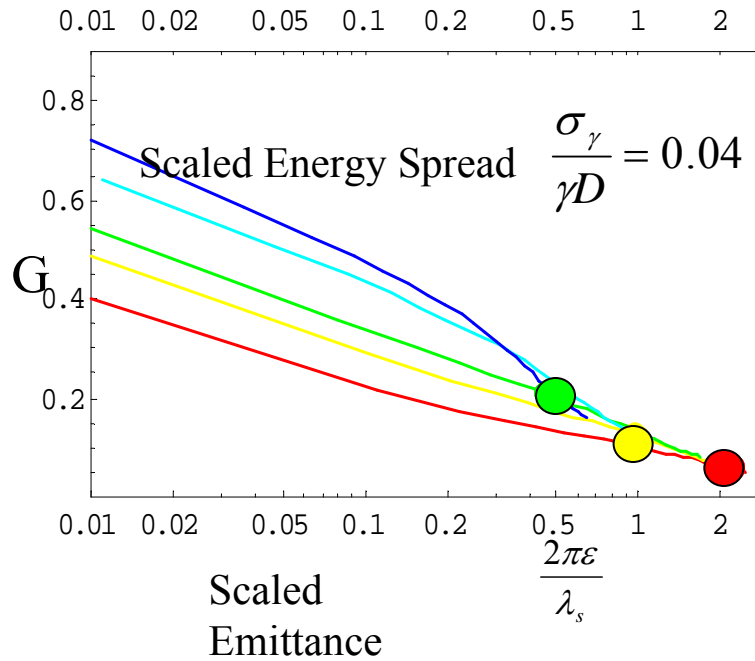


focusing and  $\beta$  motion



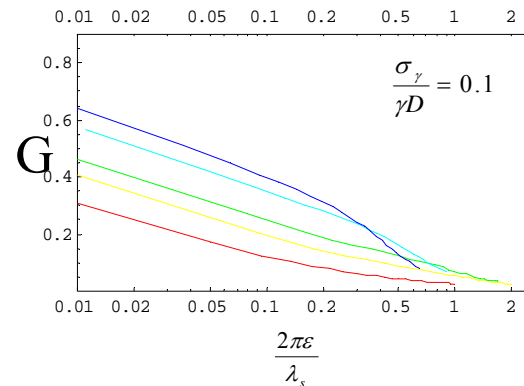
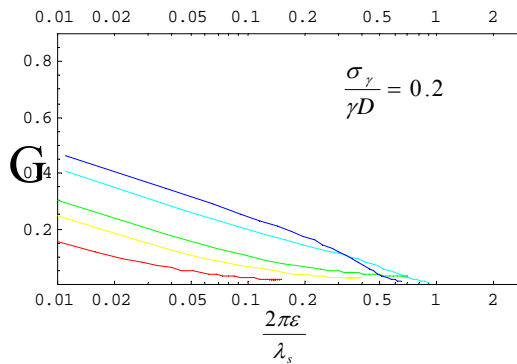
Emittance  $\varepsilon \sim$  phase space area

$$\varepsilon_n = \gamma \varepsilon \text{ (invariant during acceleration)}$$



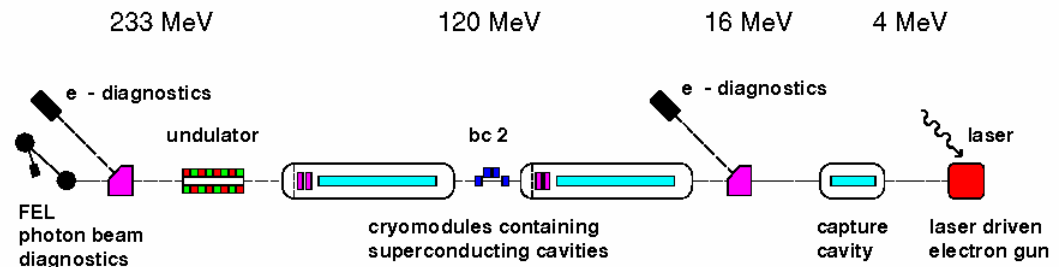
**On the scaled function plot,  
operating points do not change  
very much even though  
wavelength changes several  
orders of magnitudes**

- |                 |                         |                      |
|-----------------|-------------------------|----------------------|
| Scaled Focusing | $\frac{k_\beta}{k_w D}$ |                      |
| <u>Blue</u>     | 1.05                    | ● DUV(BNL)(100nm)    |
| Cyan            | 0.15                    | ● TTF (DESY)(6nm)    |
| Green           | 0.08                    | ● LCLS(SLAC)(0.15nm) |
| Yellow          | 0.08                    |                      |
| Red             | 0.03                    |                      |

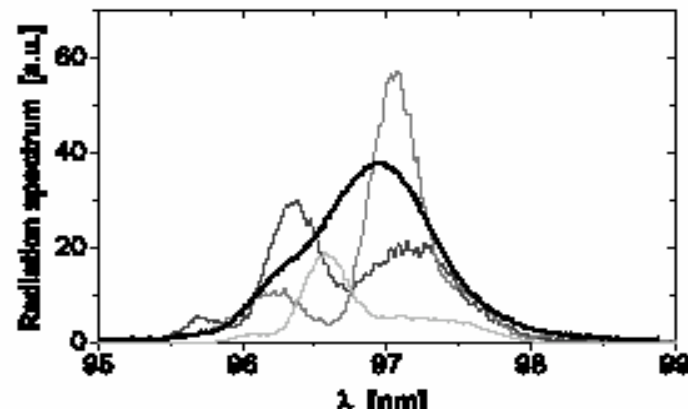


## A series of SASE experiment confirming the Theory

- LANL, UCLA: 15  $\mu\text{m}$  (1999)
- APS: Luetle 0.4  $\mu\text{m}$  (2000-2001)
- BNL, SLAC, UCLA: 0.8  $\mu\text{m}$  (2001)
- DESY: 0.1  $\mu\text{m}$  (2002)
- BNL: UVFEL 0.266  $\mu\text{m}$  (2002)



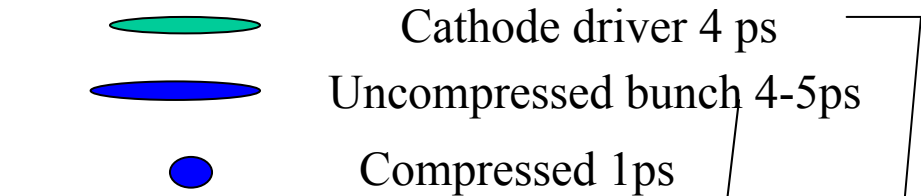
Schematic of DESY SASE experiment at 100 nm



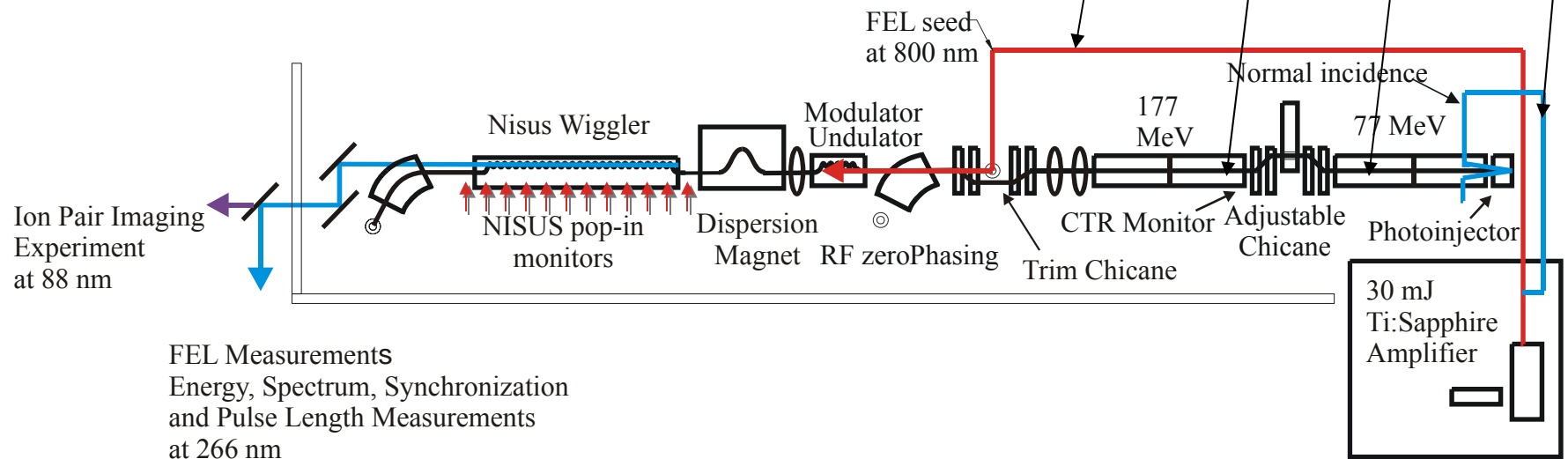
Single shot and average SASE spectrum showing fluctuation and spikes

# Deep UV Free Electron Laser at SDL

Relation of pulse lengths  
During 2002-2003 operation



Seed 9ps



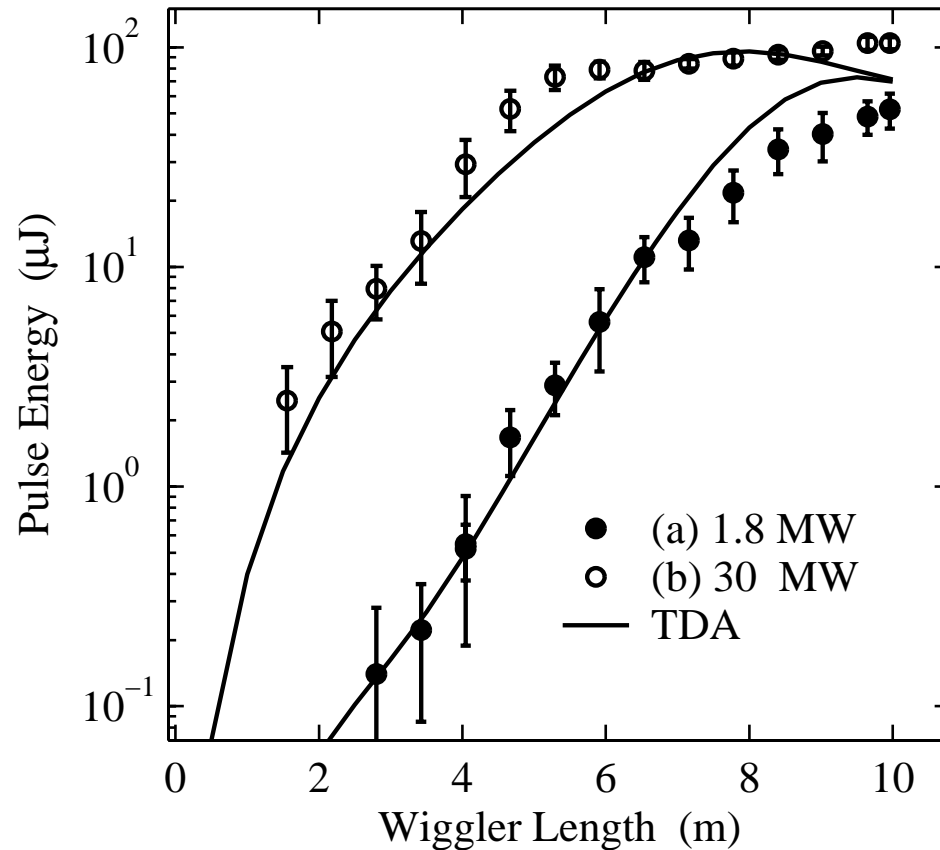
## NISUS Undulator Parameters

- Period  $\lambda_w = 3.9$  cm
- Length 10 m
- Canted poles provide horizontal focusing and reduce vertical focusing
- 4-wire focusing provide tuning ability to reach equal focusing Natural focusing: betatron wavelength  $\lambda_\beta = 20$  m
- Because focusing is not strong, also because period 3.9 cm is large, gain length is far from optimized.

### FEL gain length

- Simulation: HGHG reaches saturation at 400 nm with current  $I_0 = 300$  amp, emittance  $\varepsilon_n = 5$  mm-mrad, energy spread  $\sigma_\gamma / \gamma = 1.5 \times 10^{-3}$ , gain length  $L_G = 1.1$  m
- We do not expect SASE saturation, because we have only 10 gain length
- SASE in February 2002 :  $L_G = 0.9$  m

## Two sets of data compared with Simulation



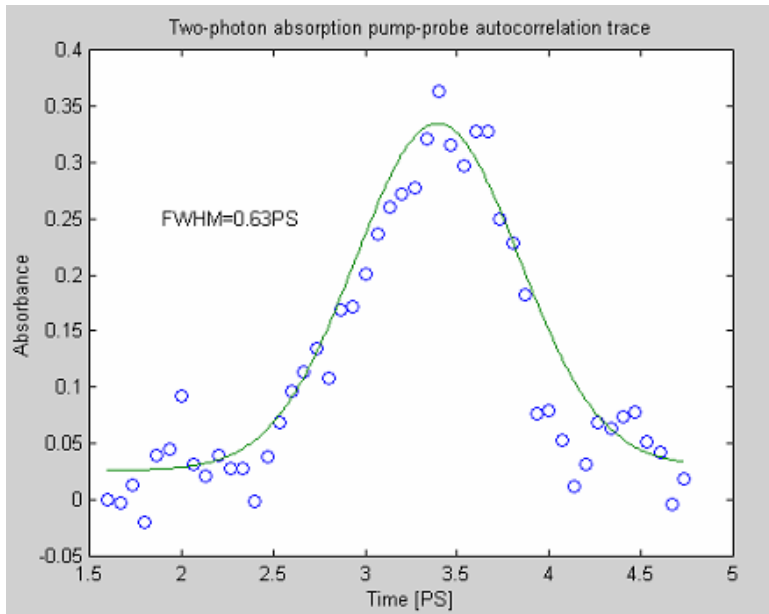
current 300 Amp, energy spread  $\sigma\gamma/\gamma = 1 \times 10^{-4}$ , dispersion  $d\psi/d\gamma = 8.7$

(a) **12/11/02 Model:**  $P_{in} = 1.8 \text{ MW}$  pulse length 0.6 ps, slice emittance 2.7 mm-mrad

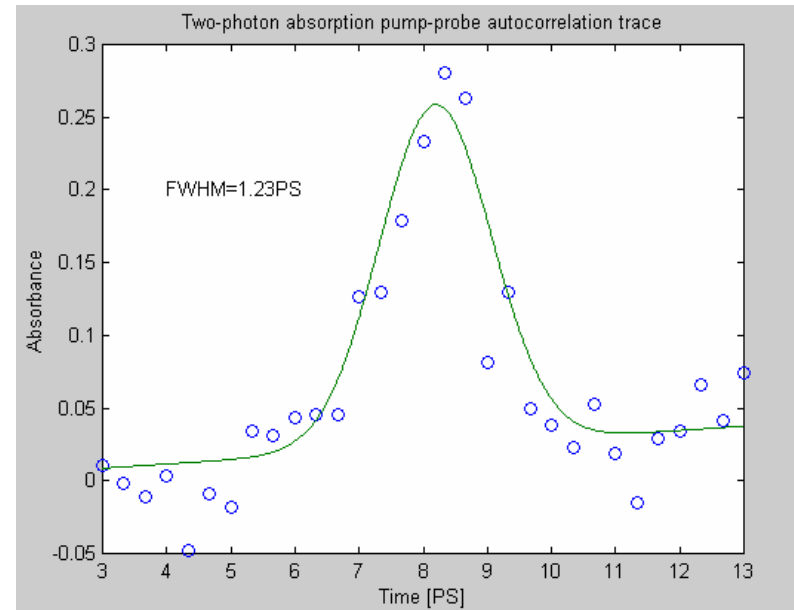
(b) **3/19/03 Model:**  $P_{in} = 30 \text{ MW}$  pulse length 1 ps, projected emittance 4.7 mm-mrad

**Whole bunch contributes to the output**

# Autocorrelation Pulse Length Measurement



(a)



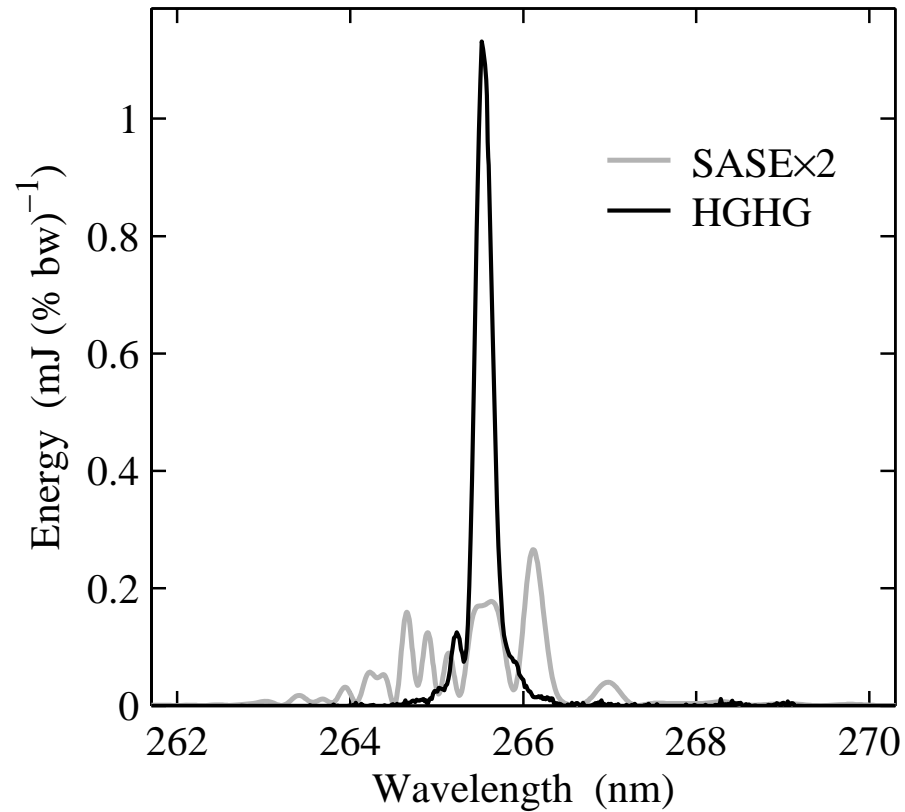
(b)

current 300 Amp, energy spread  $\sigma\gamma/\gamma = 1 \times 10^{-4}$ , dispersion  $d\psi/d\gamma = 8.7$

(a) 12/11/02 Model:  $P_{in} = 1.8\text{MW}$  pulse length 0.6 ps, slice emittance 2.7 mm-mrad

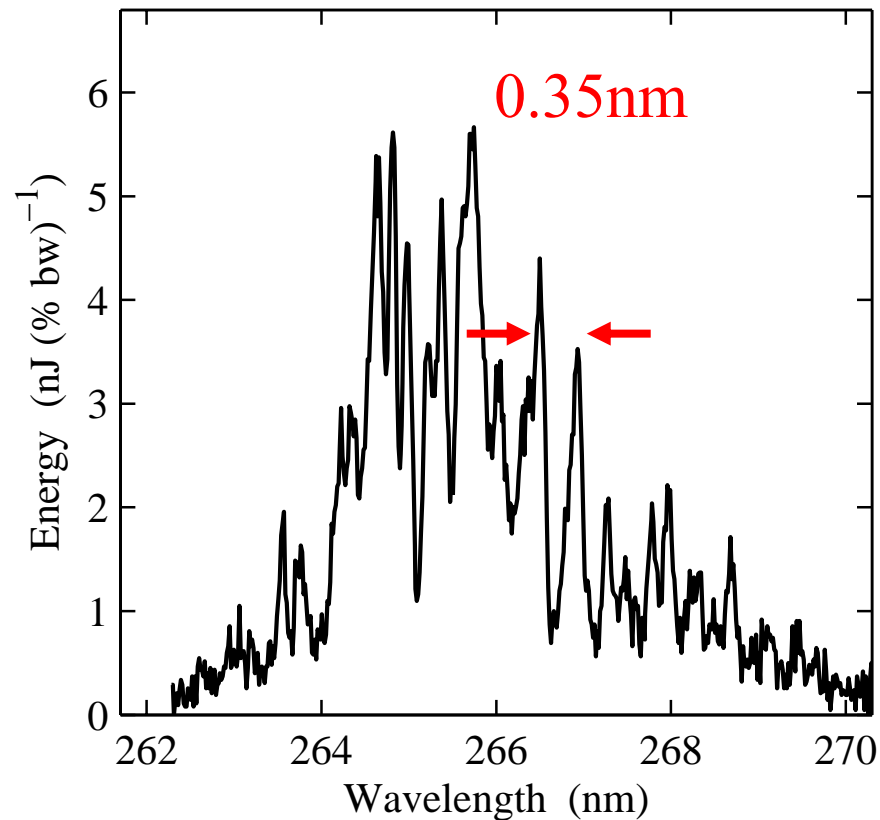
(b) 3/19/03 Model:  $P_{in} = 30\text{MW}$  pulse length 1 ps, projected emittance 4.7 mm-mrad

**Whole bunch contributes to the output**



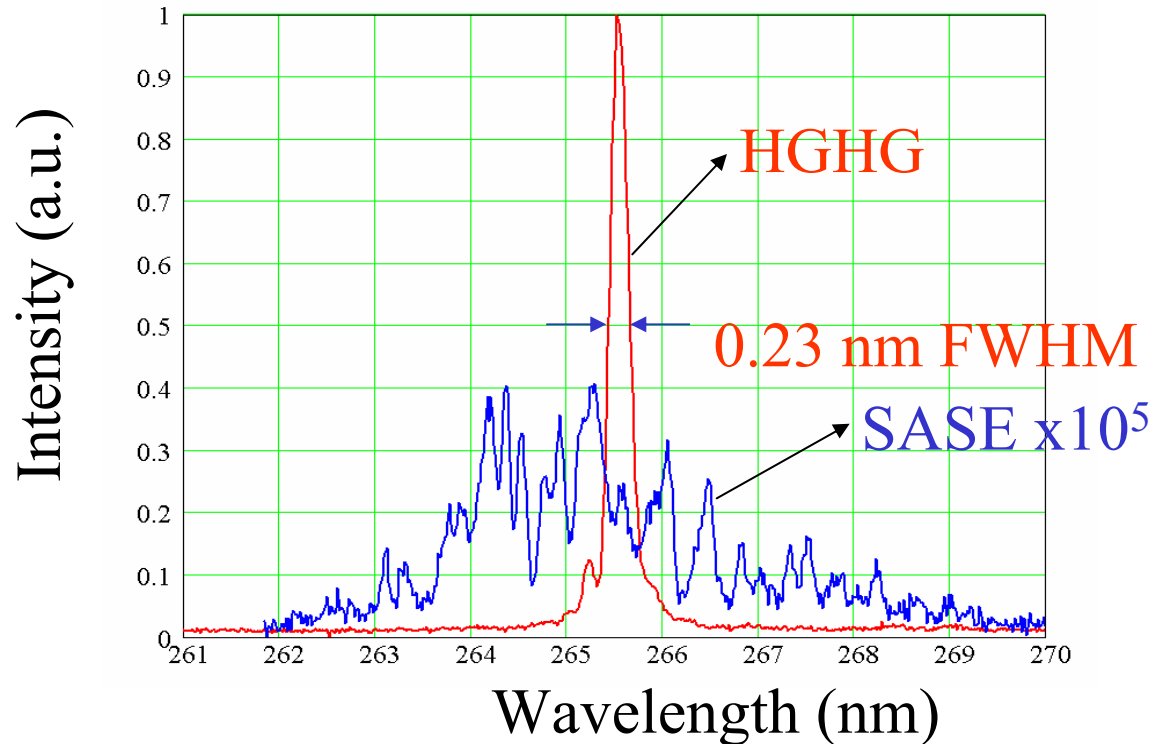
If NISUS were increased to 20 m, the SASE would be saturated. The gray line is a Genesis Simulation.

Average spacing between peaks in SASE spectrum also gives pulse length 1ps



$$T_b = \frac{\lambda^2}{0.64 c \Delta \lambda} = \frac{(266 \text{ nm})^2}{0.64 \times 3 \times 10^8 \text{ m/s} \times 0.35 \text{ nm}} \approx 1 \text{ ps}$$

S.Krinsky  
(2001)



For a flat-top 1ps pulse the bandwidth is

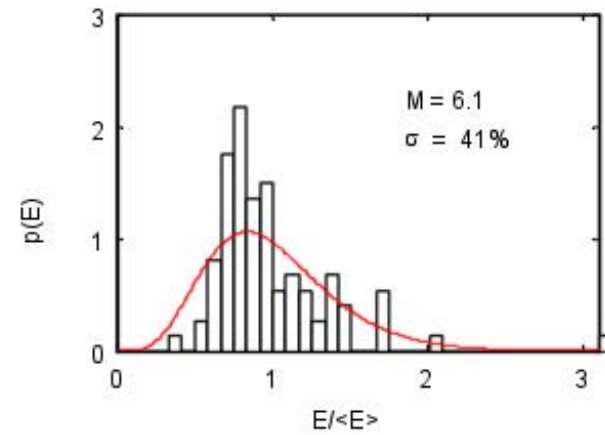
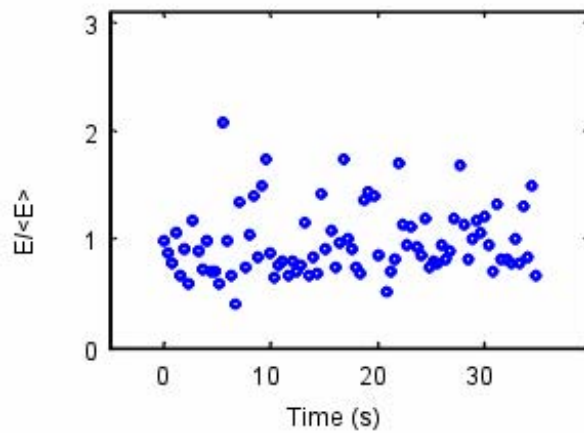
$$\Delta\lambda = \frac{\lambda^2}{\tau c} = \frac{(266\text{nm})^2}{1\text{ps} \times 3 \times 10^8 \text{ m/s}} \approx 0.23\text{nm}$$

HGFG output is nearly Fourier transform limited

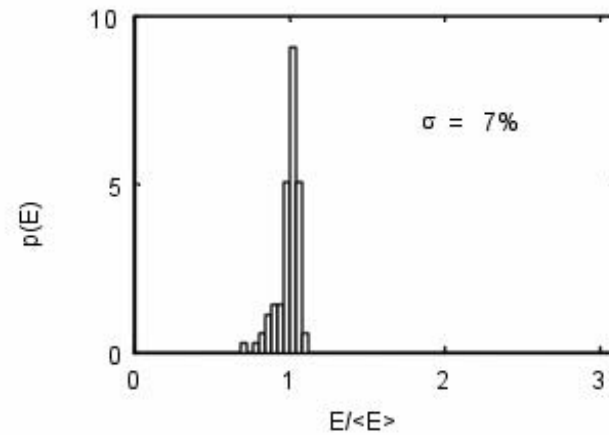
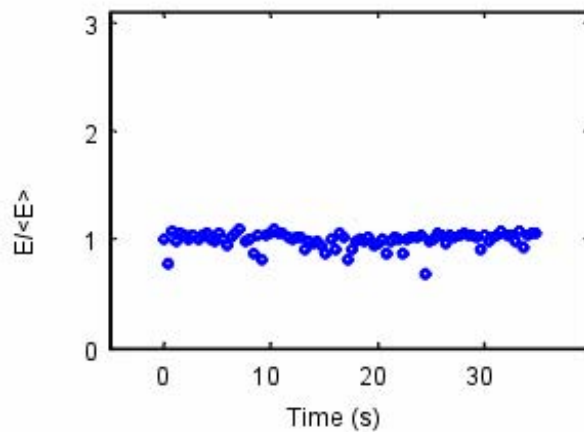
# Shot to Shot Intensity Fluctuation

## Shows High Stability of HGHG output

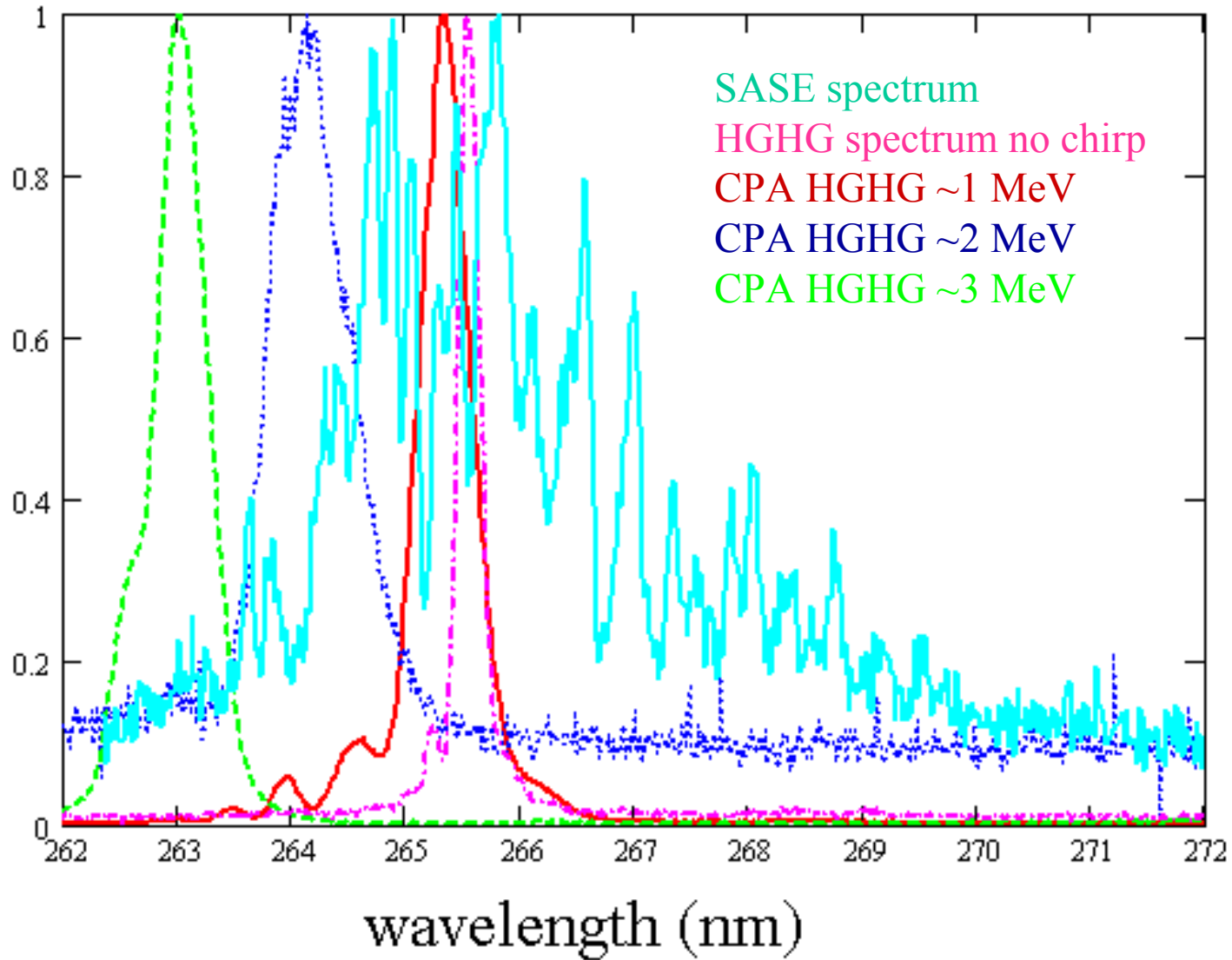
SASE



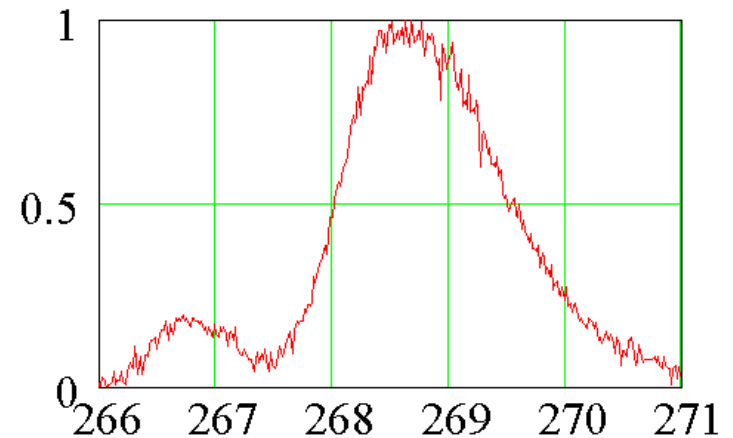
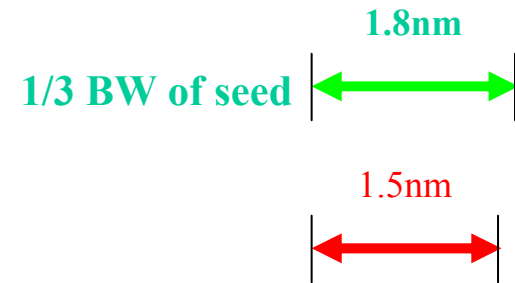
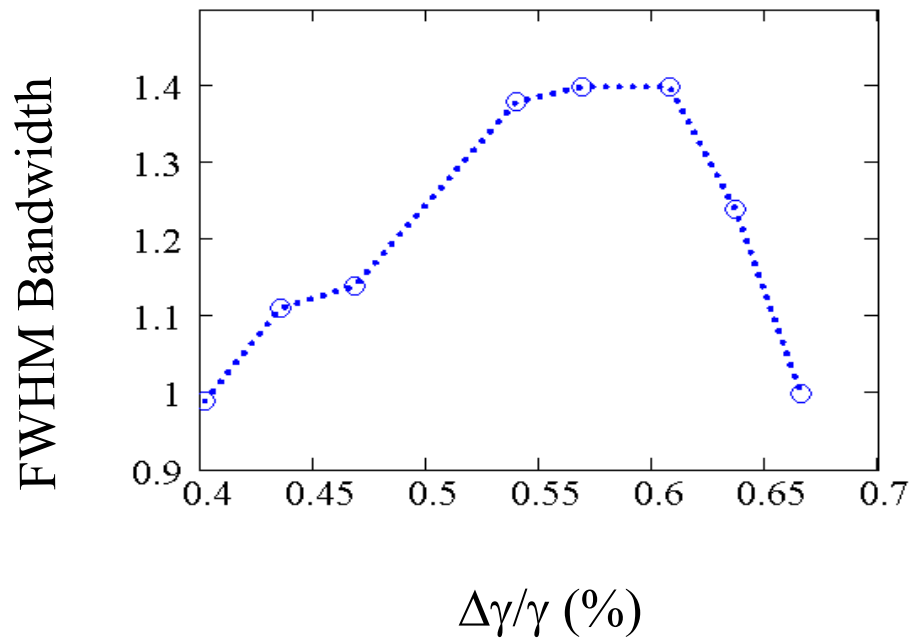
HGHG



# Chirped and Unchirped HGHG spectra



# Bandwidth vs Chirp



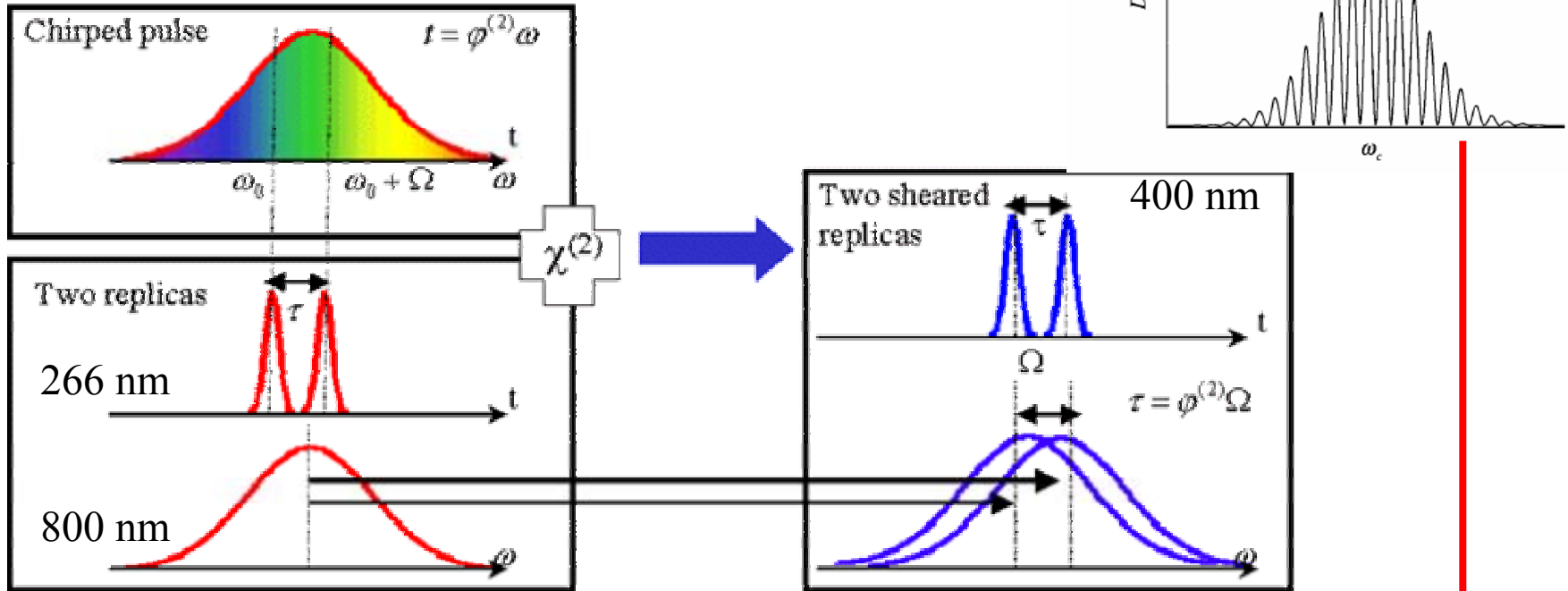
*Seed bandwidth is 5.5 nm currently, thus 1.8 nm is expected at third harmonic.*

*Spectrum at  $\Delta\gamma/\gamma = 0.62$  %*

- **Next: measurement of phase distortion**
- **CPA: expected pulse length > 50fs (B. Sheehy, Z. Wu)**
- **R&D: CPA at 100 nm?**

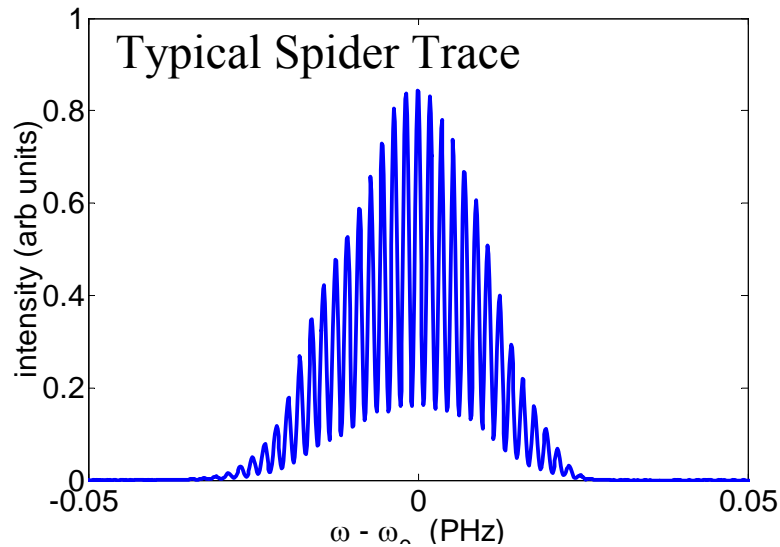
# Measuring the spectral phase: SPIDER (Spectral Interferometry for Direct Electric-Field Reconstruction)

(Walmsley group, Oxford)

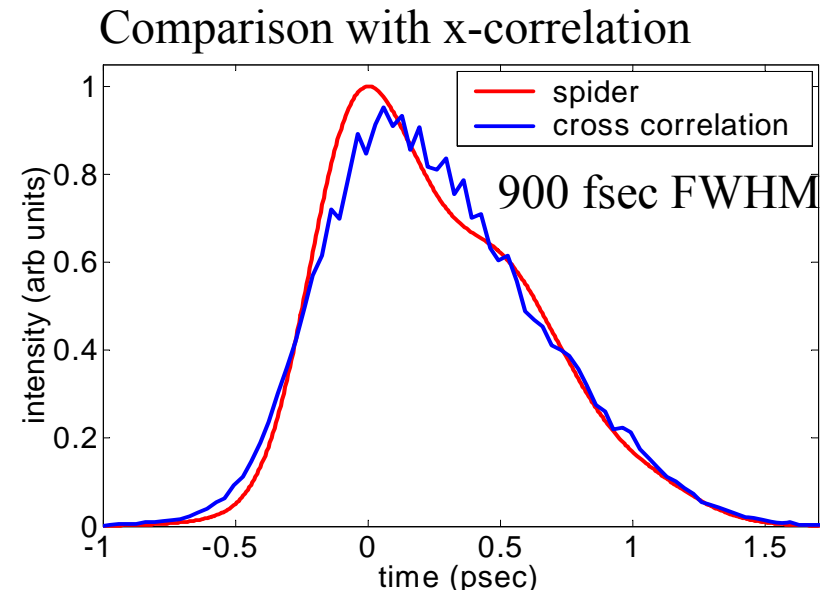
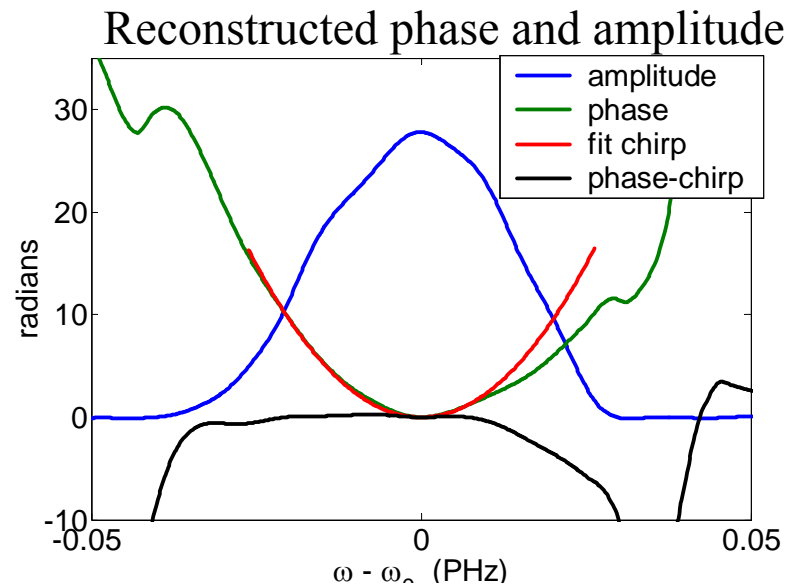


$$D(\omega_c) = |\tilde{E}(\omega_c - \Omega)|^2 + |\tilde{E}(\omega_c)|^2 + 2|\tilde{E}(\omega_c - \Omega)\tilde{E}(\omega_c)| \cos[\phi_\omega(\omega_c - \Omega) - \phi_\omega(\omega_c) - \tau\omega_c].$$

# Spidering a laboratory 266 nm source (B. Sheehy, Z. Wu)

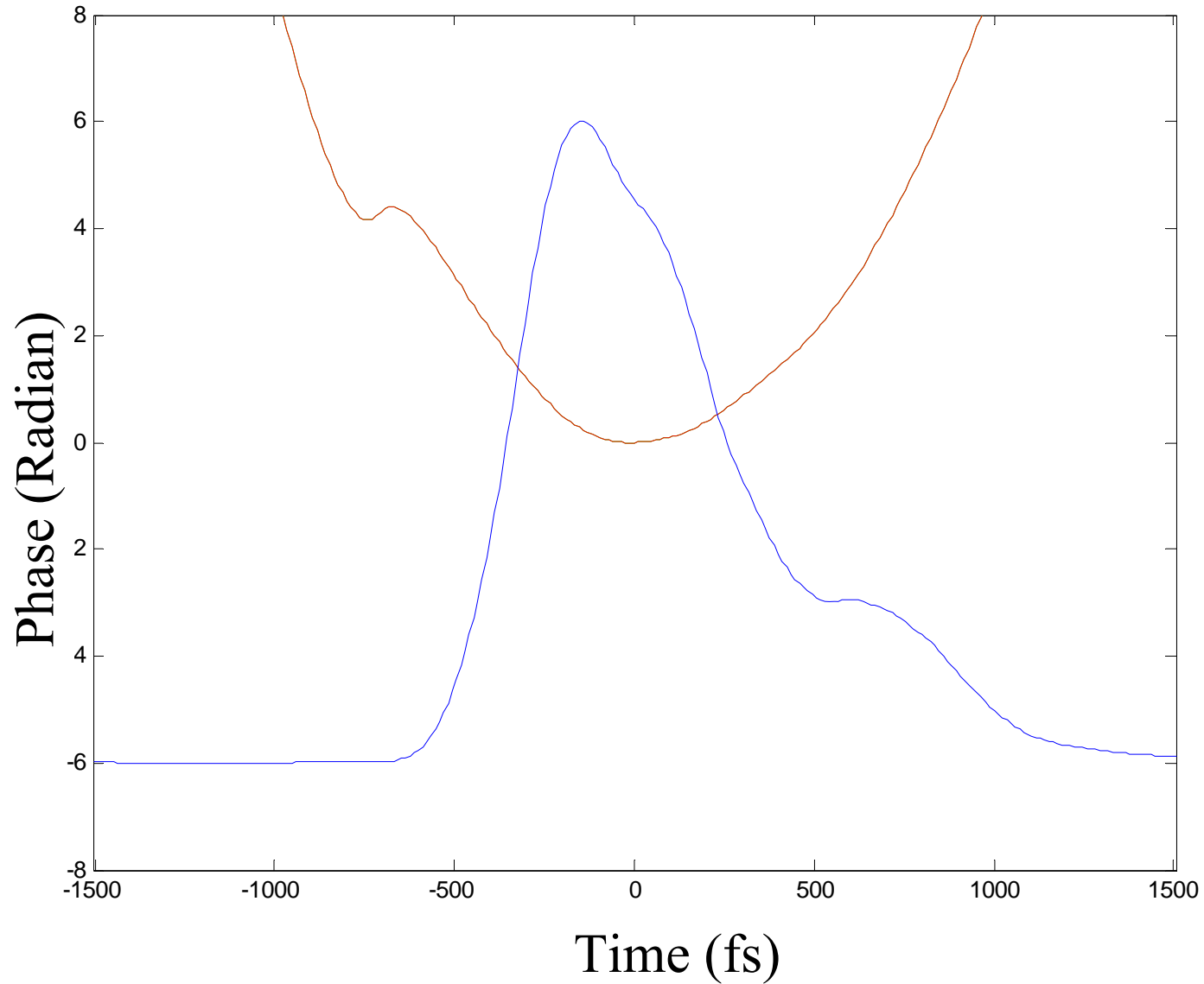


- Undercompress a 100 femtosecond 800 nm Ti:Sapph chirped-pulse-amplification system
- Frequency-triple in BBO to 266 nm (spoil phase matching to create an asymmetry in the time profile)
- Compare scanning multi-shot x-correlation of the 266 nm and a short 800 nm pulse with the average reconstruction, convolved with 250 fsec resolution of the x-correlator

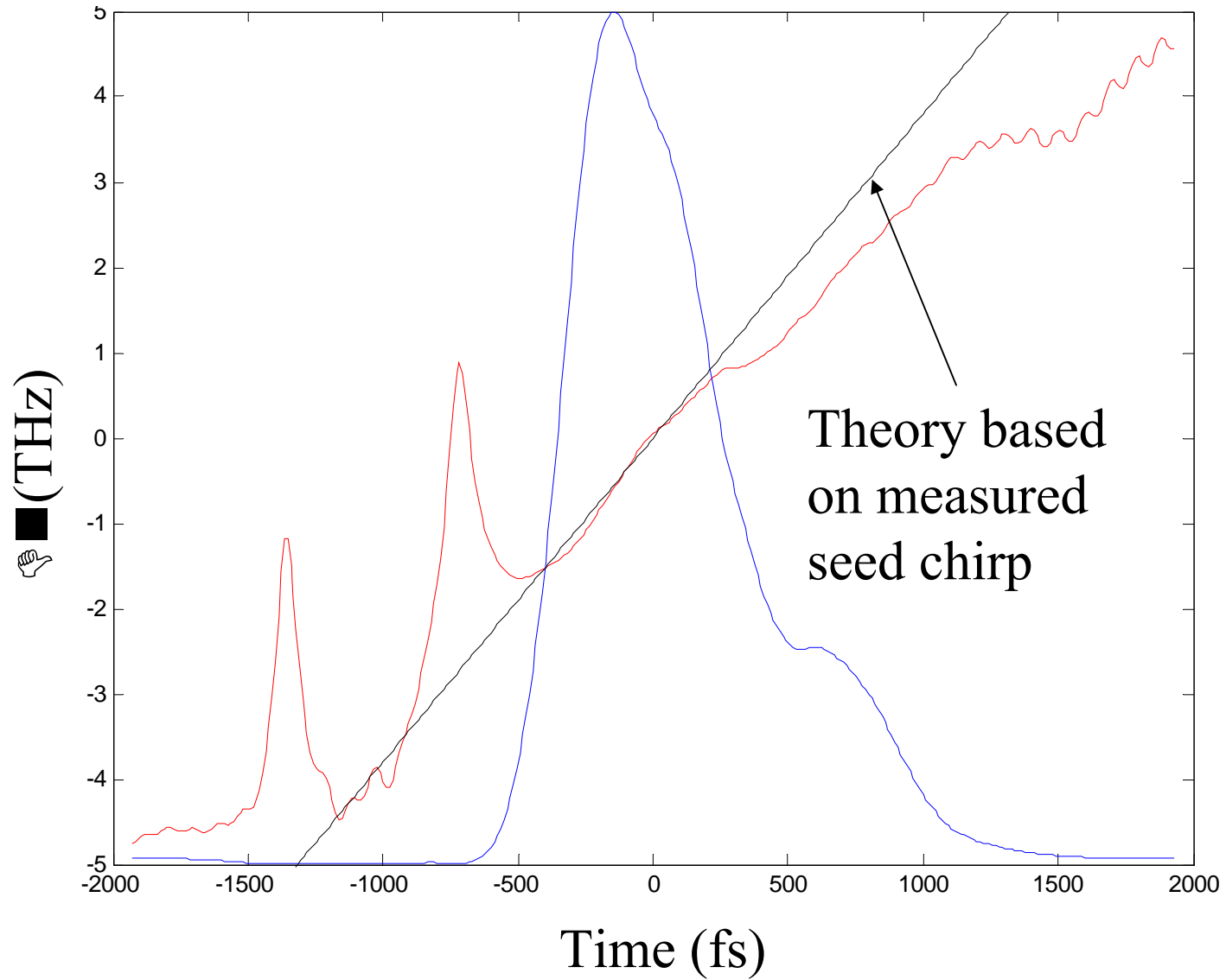


# Preliminary Data of SPIDER for a chirped HGHG

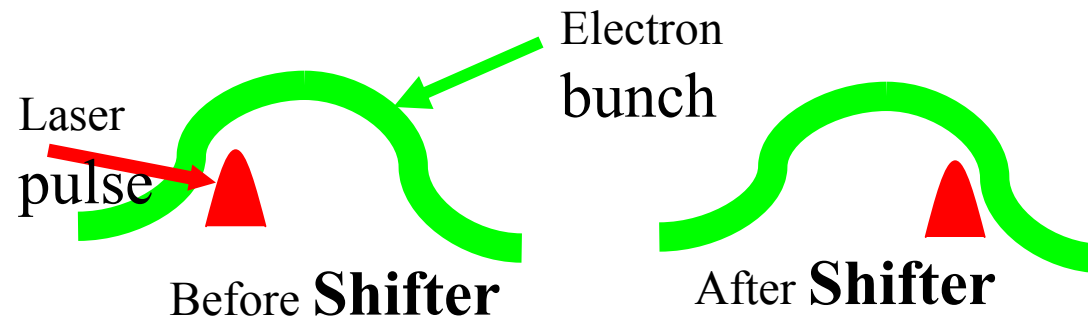
## Phase and amplitude vs. time



# Preliminary Data of SPIDER for a chirped HGHG Frequency change vs. time (electron beam energy chirp unmatched )



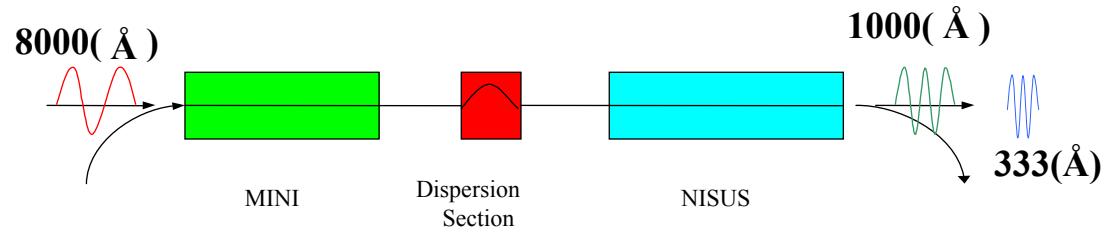
# Fresh Bunch Technique



Technical issues:

- Time jitter of seed  $\ll$  electron bunch length
- Reduce number of stages: higher harmonic, smaller local energy spread
- Shorter seed laser pulse

## Energy Upgrade for 100 nm Output 8'th Harmonic Generation



**HGHG from 800 nm → 100 nm,  
Output at 100 nm: ~0.1 mJ,  
e-beam: 300 Amp, 3 mm-mrad,  
Linac Upgrade: Energy 300 MeV**

# Cascading HGHG for Shorter Wavelength

$P_{in}=1.5$  MW

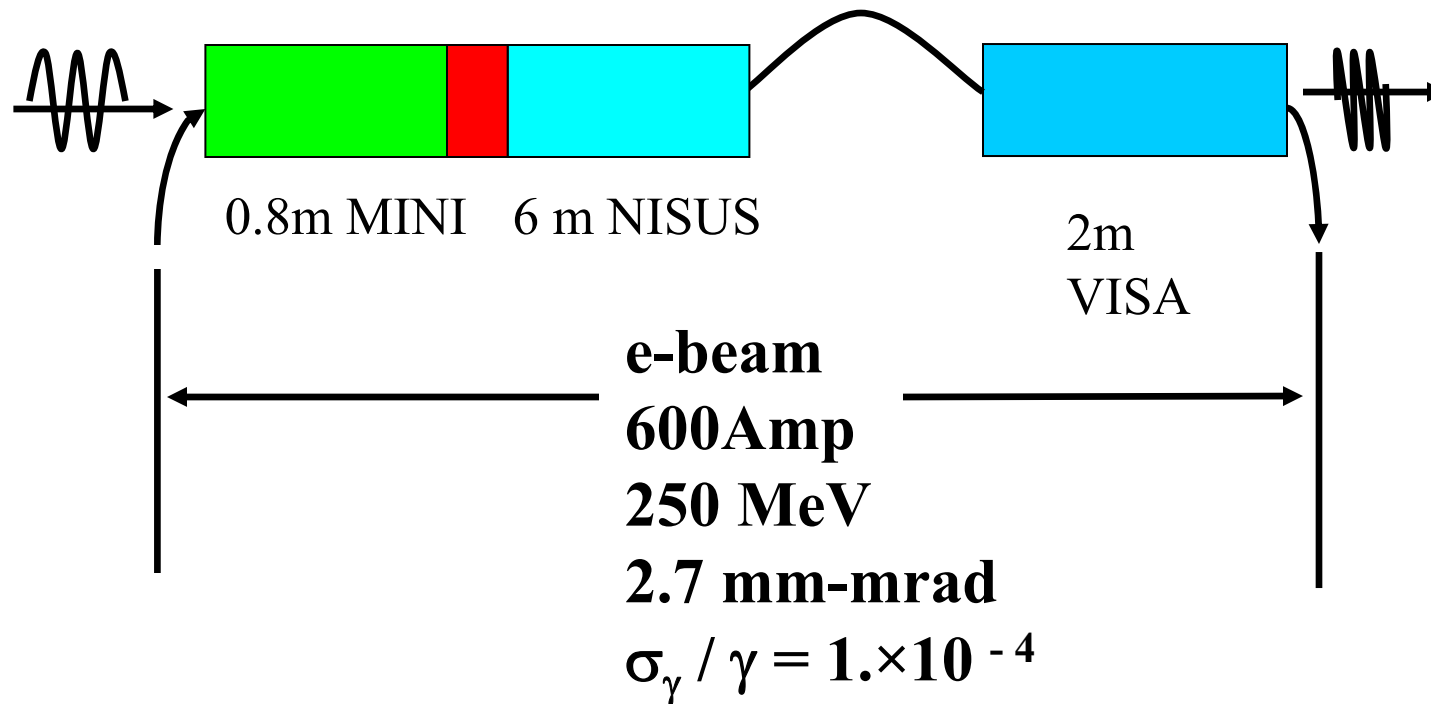
266nm

56 MW

133nm

$P_{out}=140$  MW

66.5 nm



Pulse length  $\sim 0.5$ ps  $\rightarrow 70\mu$ J

Range: 50 nm—100nm