

# OVERVIEW OF SUSY GUT

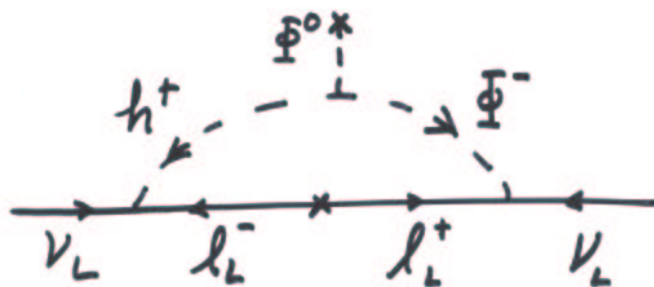
## MODELS OF NEUTRINO MIXING

WHERE DOES  $\nu$  MASS COME FROM?  
 GUTs (SEE-SAW MECHANISM)  
 or NEW PHYSICS AT LOW ENERGY?  
 (or EXTRA DIMENSIONS?)

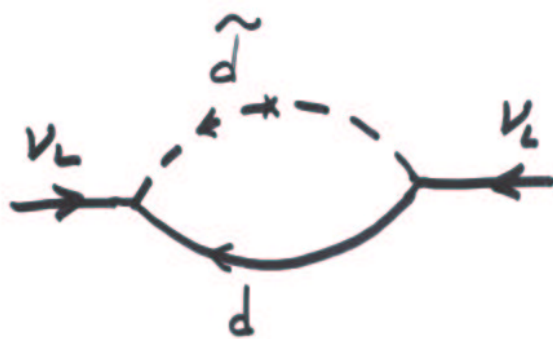
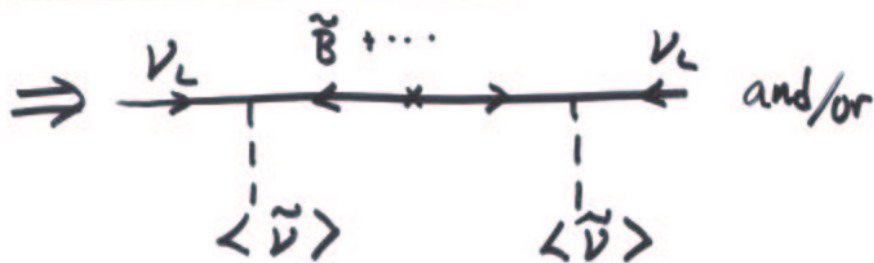
EXAMPLES OF POSSIBLE LOW ENERGY NEW PHYSICS THAT PRODUCE  $m_\nu$ .

### ZEE MODEL

$$h^+ L_i L_j + h^+ \Phi_\alpha \Phi_\beta \Rightarrow$$



### SUSY with $\mathcal{R}$



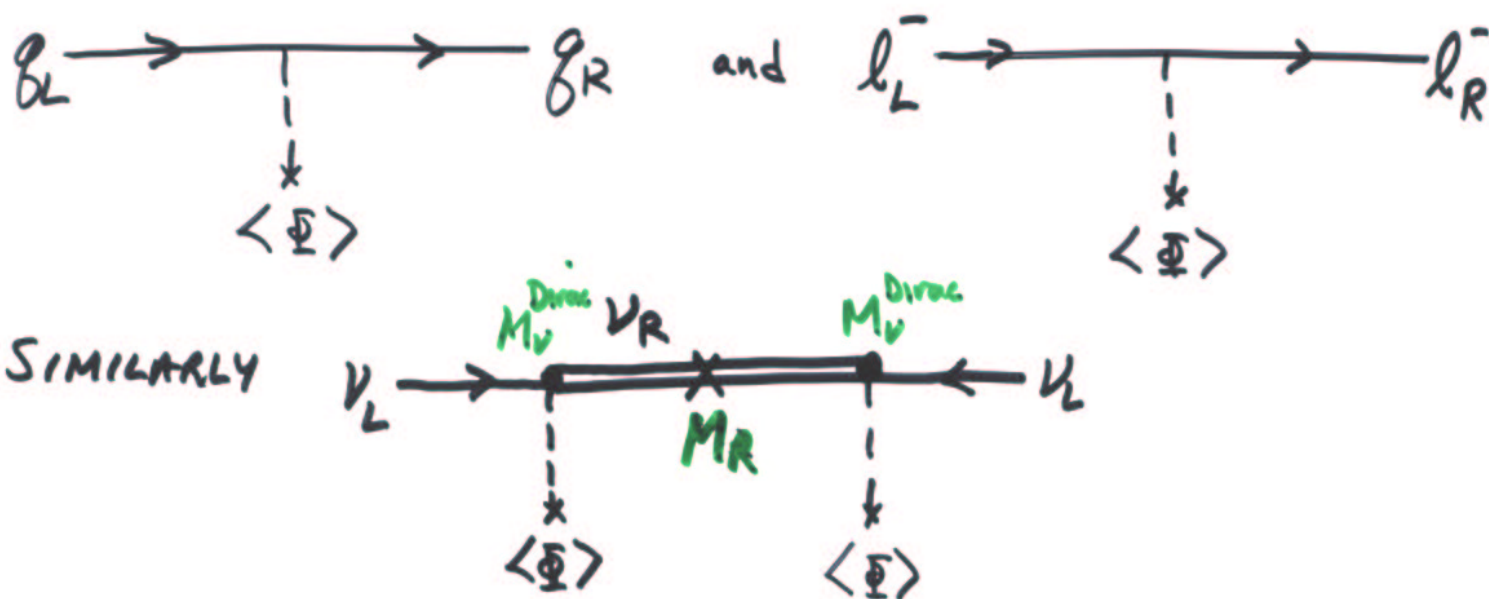
•  $m_\nu^{\text{loop}} \sim \frac{\pi(\lambda)}{16\pi^2} M_w \Rightarrow \pi(\lambda) \approx \frac{10^{-12}}{\text{WHY?}}$

- NEW PHYSICS NOT STRONGLY MOTIVATED ON OTHER GROUNDS.

# IN SUSY GUTs

2.

- NEW PHYSICS IS VERY WELL MOTIVATED
- EXPLANATION OF ORIGIN OF  $m_\nu$  SIMILAR TO EXPLANATION OF OTHER LIGHT FERMION MASSES. i.e. EXISTENCE OF RIGHT-HANDED  $\nu$ .
- MAGNITUDE OF  $m_\nu$  EXPLAINED.



$$M_\nu = - M_{\nu}^{\text{Dirac}T} M_R^{-1} M_{\nu}^{\text{Dirac}}$$

$$m_{\nu_\tau} \sim m_t M_R^{-1} m_t$$

$$M_R \sim m_t^2 / m_{\nu_\tau} \sim \frac{(175 \text{ GeV})^2}{(6 \times 10^{-11} \text{ GeV})}$$

$$M_R \sim \frac{1}{2} \times 10^{15} \text{ GeV}$$

- GRATIFYINGLY CLOSE TO  $M_{\text{GUT}}$  FROM RGE.
- STRONG PRESUMPTION IN FAVOR OF SUSY GUT/SEE SAW EXPLANATION OF  $m_\nu$ .
- GUTs ALSO ALLOW POSSIBILITY OF PREDICTIVE MODELS

HOWEVER, DISCOVERY OF LARGE  $\theta_{atm}$  CAME AS A SURPRISE FROM GUT POINT OF VIEW. 3.

## PUZZLING FEATURES

- GUTS RELATE QUARKS AND LEPTONS  
eg. MINIMAL  $SU(5) \Rightarrow M_D = M_L^T$   
MINIMAL  $SO(10) \Rightarrow M_D = M_L \propto M_U = M_\nu^{Dirac}$   
= SYMMETRIC.

SO WHY AREN'T LEPTONIC ANGLES SMALL LIKE QUARK (CKM) ANGLES?

- THERE HAD SEEMED TO BE A RELATION, FOR QUARKS, BETWEEN MASS RATIOS AND ANGLES. SO WHY, FOR LEPTONS, MASS RATIOS SMALL ( $m_e \ll m_\mu \ll m_\tau$  AND  $\Delta M_{sol}^2 \ll \Delta M_{atm}^2$ ) WHEREAS ANGLES LARGE ( $\sin^2 2\theta_{atm} = 1.0$ ,  $\tan^2 \theta_{sol} \approx 0.4$  FOR LMA, BUT  $\nu_{e3} < 0.16$ )?

Old empirical relation  $\tan \theta_c \approx \sqrt{m_d/m_s}$   
explained by "texture":

$$\left[ \begin{array}{cc} 0 & \epsilon \\ \epsilon & 1 \end{array} \right]_m \rightarrow \left. \begin{array}{l} \tan 2\theta = 2\epsilon \Rightarrow \theta \approx \epsilon \\ |m_d/m_s| \approx \epsilon^2 \end{array} \right\}$$

[Weinberg; Wilczek and Zee; Fritzsch (1979)]

This was generalized to 3 families by Fritzsche, and many models since then predict relations between angles and mass ratios. 4.

eg Babu and Nandi (1999)

$$M_D \approx \begin{bmatrix} \epsilon^6 & \epsilon^6 & \epsilon^6 \\ \epsilon^6 & \epsilon^4 & \epsilon^4 \\ \epsilon^6 & \epsilon^4 & \epsilon^2 \end{bmatrix}_m; \quad M_u \approx \begin{bmatrix} \epsilon^6 & \epsilon^4 & \epsilon^4 \\ \epsilon^4 & \epsilon^2 & \epsilon^2 \\ \epsilon^4 & \epsilon^2 & 1 \end{bmatrix}_m$$

Note characteristic "hierarchical" form  
 $\equiv$  elements get smaller up and to left  
of diagonal elements.  
 $\Rightarrow$  small mass ratios and angles,  
related.

However, for neutrinos

$$\Delta M_{\text{sol}}^2 / \Delta M_{\text{atm}}^2 \lesssim 10^{-2}$$

$$\text{But } \theta_{\text{atm}} \cong \pi/4$$

$$\theta_{\text{sol}} \approx \pi/6$$

(On other hand  $|U_{e3}| \leq 0.16$ )

Two ways to explain large leptonic angle:  
From  $M_\nu$  or from  $M_L$ .

5.

Recall:

$$U_{\text{leptonic}} = U_L^\dagger U_\nu = \begin{bmatrix} U_{e1} & U_{e2} & U_{e3} \\ U_{\mu 1} & U_{\mu 2} & U_{\mu 3} \\ U_{\tau 1} & U_{\tau 2} & U_{\tau 3} \end{bmatrix}$$

$\swarrow$  solar  
 $\leftarrow$  atm

$$\cong \underbrace{\begin{bmatrix} 1 & & \\ & C_a & S_a \\ & -S_a & C_a \end{bmatrix}}_{R_{23}(\theta_{\text{atm}})} \underbrace{\begin{bmatrix} C_s & S_s & \\ -S_s & C_s & \\ & & 1 \end{bmatrix}}_{R_{12}(\theta_{\text{solar}})} = \begin{bmatrix} C_s & S_s & 0 \\ -S_s C_a & -C_s C_a & S_a \\ S_s S_a & -C_s S_a & C_a \end{bmatrix}$$

Three simple possibilities

A)  $U_L \cong I, \quad U_\nu \cong R_{23}(\theta_{\text{atm}}) R_{12}(\theta_{\text{solar}})$

i.e. both large mixings from  $\nu$  sector  
[first 23, then 12 rotation to diagonalize  $M_\nu$ ]

B)  $U_\nu \cong I, \quad U_L \cong R_{12}(\theta_{\text{solar}}) R_{23}(\theta_{\text{atm}})$

i.e. both large mixings from  $l^-$  sector  
[first 12, then 23 rotation to diagonalize  $M_L$ ]

C)  $U_L \cong R_{23}(\theta_{\text{atm}}) \quad U_\nu \cong R_{12}(\theta_{\text{solar}})$

$\theta_{\text{atm}}$  from  $l^-$  sector,  $\theta_{\text{solar}}$  from  $\nu$  sector.

How could large  $\theta_{23}$  come from  $M_\nu$  (case A) in GUT/see-saw models? 6.

**Obvious Problem:**

- $M_U, M_D$  presumably hierarchical.
- GUTs would tend to give  $M_\nu^{\text{Dirac}}$  hierarchical also (for  $SO(10)$ , for example).

- Then see-saw formula  $M_\nu = -M_\nu^{\text{Dirac}T} M_R^{-1} M_\nu^{\text{Dirac}}$  would tend to give  $M_\nu$  hierarchical.

eg. Let  $M_\nu^{\text{Dirac}} \propto \begin{bmatrix} \epsilon' & & \\ & \epsilon & \\ & & 1 \end{bmatrix}, \quad \epsilon' \ll \epsilon \ll 1$

$$M_R^{-1} = \begin{bmatrix} a_{11} & a_{12} & a_{13} \\ a_{12} & a_{22} & a_{23} \\ a_{13} & a_{23} & a_{33} \end{bmatrix}$$

$$\Rightarrow M_\nu = \begin{bmatrix} \epsilon'^2 a_{11} & \epsilon' \epsilon a_{12} & \epsilon' a_{13} \\ \epsilon' \epsilon a_{12} & \epsilon^2 a_{22} & \epsilon a_{23} \\ \epsilon' a_{13} & \epsilon a_{23} & a_{33} \end{bmatrix} = \text{hierarchical (typically)}$$

$$\theta_{23} \approx \frac{\epsilon a_{23}}{a_{33} - \epsilon^2 a_{22}}$$

$\ll 1$  unless, say,  $a_{33}/a_{23} \sim \epsilon$

(Conspiracy between Majorana, Dirac matrices)  
 $M_R$   $M_\nu^{\text{Dirac}}$

However, large  $\theta_{atm}$  can come from  $M_\nu$ , even with  $M_\nu^{Dirac}$  = hierarchical and with no "conspiracy". 7

EXAMPLE 1 (Teebäck and Sumino, 1998)

$$M_\nu^{Dirac} = \begin{bmatrix} \epsilon' \epsilon^2 & & \\ & \epsilon & \epsilon \\ & \mathcal{O}(\epsilon^2) & 1 \end{bmatrix} m, \quad M_R = \begin{bmatrix} & & A \\ & 1 & \\ A & & \end{bmatrix} m_R$$

$$\Rightarrow M_\nu = \begin{bmatrix} 0 & \mathcal{O}(\epsilon' \epsilon^4 / A) & \epsilon' \epsilon^2 / A \\ \mathcal{O}(\epsilon' \epsilon^4 / A) & \epsilon^2 & \epsilon^2 \\ \epsilon' \epsilon^2 / A & \epsilon^2 & \epsilon^2 \end{bmatrix} \frac{m^2}{m_R}$$

$$R_{23}(\pi/4)^T M_\nu R_{23}(\pi/4) \cong \begin{bmatrix} 0 & -\frac{\epsilon' \epsilon^2}{\sqrt{2}A} & \frac{\epsilon' \epsilon^2}{\sqrt{2}A} \\ -\frac{\epsilon' \epsilon^2}{\sqrt{2}A} & 0 & 0 \\ \frac{\epsilon' \epsilon^2}{\sqrt{2}A} & 0 & 2\epsilon^2 \end{bmatrix} \frac{m^2}{m_R}$$

"Pseudo-Dirac" (after rotating away 13, 31 elements)  $\Rightarrow \theta_{sol} \cong \pi/4$

EXAMPLE 2 (Altarelli, Feruglio + Masina, 1999)

$$M_\nu = \begin{bmatrix} \epsilon' & \epsilon' & \epsilon' \\ \epsilon' & \epsilon & \epsilon \\ \epsilon' & \epsilon & 1 \end{bmatrix} \begin{bmatrix} - & & \\ & 1 & \\ & & - \end{bmatrix} \begin{bmatrix} \epsilon' & \epsilon' & \epsilon' \\ \epsilon' & \epsilon & \epsilon \\ \epsilon' & \epsilon & 1 \end{bmatrix} \frac{m^2}{m_R}$$

Order  $\epsilon' \ll \epsilon \ll 1$

$M_R^{-1} \cong$  projection one very light  $\nu_R$  (second family)

$$\Rightarrow M_\nu = \begin{bmatrix} \epsilon'^2 & \epsilon' \epsilon & \epsilon' \epsilon \\ \epsilon' \epsilon & \epsilon^2 & \epsilon^2 \\ \epsilon' \epsilon & \epsilon^2 & \epsilon^2 \end{bmatrix} \frac{m^2}{m_R}$$

Lesson: Require special forms.

Also need further conditions satisfied to get LMA solar solution from  $M_\nu$ .

However, there is a simpler way to get large  $\theta_{atm}$ : from  $M_L$  (case B, or case C) 8.

Suppose  $M_D = \text{hierarchical}$  (as is typical in GUTs)

$$\text{but } M_L = \begin{bmatrix} - & - & - \\ - & - & \rho \\ - & - & 1 \end{bmatrix}^m, \quad \rho \sim 1$$

$$\Rightarrow U_L^\dagger \cong \begin{bmatrix} 1 & & \\ & C_a + S_a & \\ & -S_a & C_a \end{bmatrix} \quad \tan \theta_{atm} \cong -\rho$$

Then small  $m_2/m_3$  comes from  $M_\nu$   
whereas large  $\theta_{23} = \theta_{atm}$  comes from  $M_L$   
(Resolves a puzzle)

But then, why isn't there also a large quark angle(s); since  $M_D$  related to  $M_L$ ?

There is, but it is harmless

$$\text{minimal SU(5)} \Rightarrow M_D = \begin{bmatrix} - & - & - \\ - & - & - \\ - & \rho & 1 \end{bmatrix}^m$$

produces large mixing of right handed b, s.

Idea of "lopsided" models.

- ✱ large  $\theta_{atm}$  from  $M_L$  (not  $M_\nu$ ),  $M_L$  highly asymmetric
- ✱  $M_L$  related to  $M_D^T$  by SU(5)

$\theta_{sol}$  can come either from  $M_L$  (B) or  $M_\nu$  (C)

Example of a highly predictive SO(10) model based on "lopsided" idea (Albright, Babu, Barr) 9.

$$M_U = \begin{bmatrix} \eta & 0 & 0 \\ 0 & 0 & -\epsilon/3 \\ 0 & \epsilon/3 & 1 \end{bmatrix} m_U, \quad M_D = \begin{bmatrix} 0 & \delta & \delta' \\ \delta & 0 & -\epsilon/3 \\ \delta' & \rho + \epsilon/3 & 1 \end{bmatrix} m_D$$

$$M_\nu^{\text{Dirac}} = \begin{bmatrix} \eta & 0 & 0 \\ 0 & 0 & \epsilon \\ 0 & -\epsilon & 1 \end{bmatrix} m_U, \quad M_L = \begin{bmatrix} 0 & \delta & \delta' \\ \delta & 0 & \rho + \epsilon \\ \delta' & -\epsilon & 1 \end{bmatrix} m_D$$

With six terms, fits well all quark and charged lepton masses, CKM parameters, and  $\nu$  mixing (atm).

In this model, large  $\theta_{sol}$  can also arise through  $M_\nu$ . (Case C). (Albright, Barr 1999)

eg.  $M_R = \begin{bmatrix} & & A \\ & 1 & \\ A & & \end{bmatrix} m_R$  or  $\begin{bmatrix} & A \\ A & & 1 \end{bmatrix} m_R$

$$M_\nu = M_\nu^{\text{Dirac}T} M_R^{-1} M_\nu^{\text{Dirac}} = \begin{bmatrix} 0 & \epsilon\eta/A \\ \epsilon\eta/A & 0 \\ \eta/A & 0 \end{bmatrix} \begin{bmatrix} \eta/A \\ 0 \\ \epsilon^2 \end{bmatrix} \frac{m_U^2}{m_R}$$

After rotating away (3,3) elements  $(M_\nu)_{11} = -(\eta/\epsilon A)^2$

$\Rightarrow$  12 block "pseudo Dirac"  $\Rightarrow \theta_{sol} \approx \pi/4$

Note, no difficulty getting large  $\theta_{sol}$  from  $M_\nu$  since  $m_1/m_2$  not known to be small, could be  $\sim 1$ .

Another simple way to get large  $\theta_{sol}$  in lopsided model: from  $M_L$  also (case B) "doubly lopsided". (Babu, Barr 1996, 2002)

Consider  $M_\nu = \text{hierarchical}$

$$M_L = \begin{bmatrix} - & - & \rho' \\ - & - & \rho \\ - & - & 1 \end{bmatrix} m, \quad \rho, \rho' \sim 1$$

Rotate first from left by  $R_{12}(\theta_{sol})$ , where  $\tan \theta_{sol} = \rho'/\rho$

$$M_L \rightarrow \begin{bmatrix} - & - & - \\ - & - & \sqrt{\rho^2 + \rho'^2} \\ - & - & 1 \end{bmatrix} m$$

Then rotate from left by  $R_{23}(\theta_{atm})$ , where  $\tan \theta_{atm} = \sqrt{\rho^2 + \rho'^2}$

$$M_L \rightarrow \begin{bmatrix} - & - & - \\ - & - & - \\ - & - & \sqrt{1 + \rho^2 + \rho'^2} \end{bmatrix} m$$

$$\text{So } U_{\text{leptonic}} = U_L^\dagger U_\nu \cong \begin{bmatrix} 1 & & \\ c_a & s_a & \\ -s_a & c_a & \end{bmatrix} \begin{bmatrix} c_s & s_s \\ -s_s & c_s \\ 1 & \end{bmatrix} = \begin{bmatrix} c_s & s_s & 0 \\ -s_s c_a & c_s c_a & s_a \\ s_s s_a & -s_s c_a & c_a \end{bmatrix}$$

$$\tan \theta_{atm} \cong \sqrt{\rho^2 + \rho'^2} \sim 1$$

$$\tan \theta_{sol} \cong \rho'/\rho \sim 1$$

Question: Why is  $\theta_{atm}$  so close to maximal, i.e.  $\pi/4$ ?

This requires some non-abelian flavor symmetry

CASE A If  $\theta_{atm}$  primarily from  $M_\nu$

then need  $M_\nu \approx \begin{bmatrix} - & - & - \\ - & A & B \\ - & B & A \end{bmatrix}$

$\Rightarrow \theta_{atm} \approx \pi/4$   
 $|m_2/m_3| = \left| \frac{A-B}{A+B} \right| \ll 1 \Rightarrow A \approx B$

CASES B,C

Let us take lopsided (or doubly lopsided) form, and by non-abelian flavor symmetry (eg  $\mu_L^- \leftrightarrow \tau_L^-$ ) make

$\rho = 1$

$M_L = \begin{bmatrix} - & - & \rho' \\ - & - & 1 \\ - & - & 1 \end{bmatrix} m$

(while angles from  $U_\nu$  small)

$\Rightarrow \tan^2 \theta_{sol} \approx (\rho'/\rho)^2 = \rho'^2$   
 $\tan^2 \theta_{atm} \approx \frac{\rho^2 + \rho'^2}{\rho^2} = 1 + \rho'^2$

$\Rightarrow \tan^2 \theta_{atm} = 1 + \tan^2 \theta_{sol} \Rightarrow \sin^2 2\theta_{atm} = \frac{1 + \tan^2 \theta_{sol}}{(1 + \frac{1}{2} \tan^2 \theta_{sol})^2}$   
 eg.  $\tan^2 \theta_{sol} = 0.4 \Rightarrow \sin^2 2\theta_{atm} = 0.97$

# Lopsided mass matrices and b- $\tau$ unification.

Minimal SU(5) :  $M_D = M_L^T \Rightarrow$

$$\begin{aligned} m_b^0 &= m_\tau^0 \\ m_s^0 &= m_\mu^0 \\ m_d^0 &= m_e^0 \end{aligned}$$

actual  
 $\checkmark$   
 $\approx \times \frac{1}{3}$  (approx)  
 $\approx \times 3$  (approx)  


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 $\det M_D^0 \cong \det M_L^0$

Old Georgi-Jarlskog idea

$$M_D^0 = \begin{bmatrix} b & b \\ b & -\frac{1}{3}a \\ & & 1 \end{bmatrix} m, \quad M_L^0 = \begin{bmatrix} b & b \\ b & a \\ & & 1 \end{bmatrix} m$$

So  $\det M_D^0 = \det M_L^0$ .

$10_2 \bar{5}_2 \langle \bar{45}_H \rangle$

others:  $10_i \bar{5}_j \langle \bar{5}_H \rangle$

How good is  $m_\tau^0 = m_b^0$  ?

<u><math>\tan \beta</math></u>	<u><math>m_\tau^0 / m_b^0</math></u>
1.85	1.03
2.0	1.09
2.5	1.15
3.0	1.17
5.0	1.20
10.0	1.21
20.0	1.20
40.0	1.16
50.0	1.11
55.0	1.06
58.0	1.00

- $(\alpha_s(M_Z)) = 0.118$
- $M_b(m_b) = 4.25 \text{ GeV}$
- $M_\tau(m_\tau) = 1.777 \text{ GeV}$
- All SUSY masses taken to be degenerate at  $M_t$ .
- From  $M_t$  to  $M_G$  use 2-loop MSSM RGE.
- From 1 GeV to  $M_Z$  use 3-loop QCD, 1-loop QED RGE.

Discrepancy can be explained by

$$\tan\beta \approx 1, \text{ or } \tan\beta \approx m_t/m_b$$

or by one-loop corrections to  $m_b$  at low-energy

gluino loop



$$W \supset h_b (b_L b_L^c) H_d + \mu H_u H_d$$

$$\rightarrow \mu h_b \tilde{b}_L \tilde{b}_L^c \langle H_u \rangle^* = \mu m_b \underline{\tan\beta} \tilde{b}_L \tilde{b}_L^c$$

$$\delta m_b / m_b \Big|_{\text{gluino}} = \frac{2\alpha_3}{3\pi} (\mu M_3) \mathcal{I}(m_{\tilde{b}_1}^2, m_{\tilde{b}_2}^2, M_3^2) \underline{\tan\beta}$$

(also chargino loop)

Can give 15% correction (for large  $\tan\beta$ ) needed

Need  $\text{sgn}(\mu M_3) < 0$  generally to get  $\delta m_b < 0$  and explain discrepancy.

However, in mSUGRA typically that makes chargino loop contribution to  $b \rightarrow s\gamma$  add constructively to SM + charged Higgs loops and  $b \rightarrow s\gamma$  too big unless sparticles heavy.

[See S. Komine and M. Yamaguchi, PRD 65, 075013 (2001).]

Perhaps  $M_T^0/M_b^0$  really is  $> 1$ .

[To know, we would need to know  $\tan\beta$ ,  
and  $\mu, M_3, M_{H_u}^2$ .]

This could be result of same large  
offdiagonal elements in  $M_L$  that give  
large  $\theta_{atm}$ . i.e. effect of lopsidedness.  
(Barr + Dorsner 2002)

Consider

$$M_L = \begin{bmatrix} - & - & \rho' \\ - & - & \rho \\ - & - & 1 \end{bmatrix}_m$$

$$M_D = \begin{bmatrix} - & - & - \\ - & - & - \\ c\rho' & c\rho & 1 \end{bmatrix}_m$$

↙ "Clebsch"

$$\frac{M_T^0}{M_b^0} \approx \frac{\sqrt{1 + \rho^2 + \rho'^2}}{\sqrt{1 + c^2(\rho^2 + \rho'^2)}}$$

If angles in  $U_\nu$  negligible:

$$\tan \theta_{atm} = \sqrt{\rho^2 + \rho'^2} \quad (\text{page 10})$$

$$\frac{M_T^0}{M_b^0} \approx \frac{\sqrt{1 + \tan^2 \theta_{atm}}}{\sqrt{1 + c^2 \tan^2 \theta_{atm}}} \approx \sqrt{\frac{2}{1 + c^2}}$$

$$m_{\tau}^0/m_b^0 \cong \sqrt{\frac{2}{1+c^2}}$$

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(a) If  $\rho, \rho'$  come from  $10_3 \bar{5}_{2,1} \langle \bar{45}_H \rangle$   
(a la Georgi Jarlskog)

$$\Rightarrow \boxed{c = -1/3} \quad m_{\tau}^0/m_b^0 \cong \sqrt{\frac{9}{5}} = \underline{1.34}$$

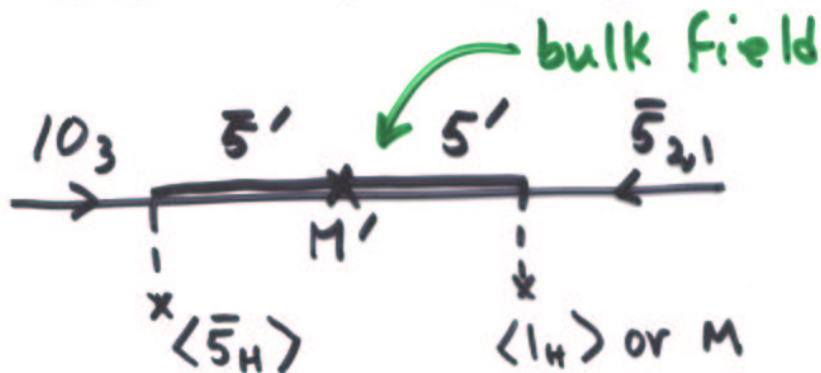
(b) If  $\rho, \rho'$  come from  $10_3 \bar{5}_{2,1} \langle \bar{5}_H \rangle \langle \frac{24_H}{M} \rangle$

by diagram:



$$\Rightarrow \boxed{c = -2/3} \quad m_{\tau}^0/m_b^0 \cong \sqrt{18/13} = \underline{1.17}$$

(c)  $\rho, \rho'$  come from similar diagram in 5d SUSY GUT ( $5'/z_2 \times z_2'$  orbifold)



With certain assumptions ( $M, M' \ll 1/R$ )

$$\text{one finds } \boxed{c \ll 1} \quad m_{\tau}^0/m_b^0 \cong \sqrt{2} = \underline{1.41}$$

## 'LOPSIDED MODELS'

- RELATIVELY SIMPLE WAY TO GET LARGE  $\nu$  ANGLES
- CAN MAKE HIGHLY PREDICTIVE (AND SUCCESSFUL) MODELS.
- MAY HAVE SOMETHING TO DO WITH  $m_{\nu}^0 / m_b^0 \approx 1$   
(INTERESTING PREDICTIONS FROM SPECIAL CASES)
- MAY ALLOW RELATIVELY SIMPLE WAY TO EXPLAIN  $\theta_{atm} \cong \text{MAXIMAL}$ .