

Proton Decay: The Missing Link of Grand Unification

Jogesh C. Pati.

(Ness, 2002).
↳ Conference, YITP, 2002

- 1) K. S. Babu, J. C. Pati & F. Wilczek - "Fermion Masses, Neutrino Oscillations & Proton Decay in the light of Superkamiokande", hep-ph/981538, Nuc. Phys. B 566, 33 (2000)
- 2) J. C. Pati, "Confronting the Conventional Ideas of Grand Unification -----", ICTP, Trieste Lectures (2001), hep-ph/0106082

I. Evidence Favoring SUSY Grand Unification

1) Family Structure - Quantum Nos. } $G(224)/SO(10)$
All members in one multiplet } $\nu_R // B-L$

2) Quantization of Q_{em}

3) Meeting of 3 gauge couplings } SUSY
 $M_X \sim 2 \times 10^{16} \text{ GeV}$ } $SO(10)/SU(5)/$
String $G(224)$

4) $\Delta m^2(\nu_\mu - \nu_\tau) \sim (1/20 \text{ eV})^2$ $\leftrightarrow G(224)/SO(10)$

$$\begin{aligned} &\downarrow \\ &m(\nu^c)_{\text{Dirac}} = m_t(M_X) \approx 120 \text{ GeV} \leftrightarrow SU(4)^c \\ &M(\nu_R^c) \approx M_{\text{GUT}}^2/M_{\text{St}} \approx 5 \times 10^{14} \text{ GeV} \end{aligned}$$

5) $m_b^0 \approx m_\tau^0$

6) $\Theta(\nu_\mu - \nu_\tau) \approx \pi/4 \leftrightarrow V_{cb} \approx 0.04$ $\leftrightarrow G(224)/SO(10)$

7) LEPTOGENESIS $\rightarrow Y_B \approx 5 \times 10^{-11} \leftrightarrow \nu_R // B-L$

Success of All 7 features Seems Nontrivial

\updownarrow
Together make Strong case in favor of
SUSY GRAND UNIFICATION // $G(224)$ or $SO(10)$ in 4D

① Evidence For Grand Unification: Family Multiplet Structure

$$G(213) = SU(2)_L \times U(1)_{Y_W} \times SU(3)^C$$

$$\begin{pmatrix} u_r & u_y & u_b \\ d_r & d_y & d_b \end{pmatrix}_L^{1/3}; \begin{pmatrix} u_r, y, b \end{pmatrix}_R^{4/3}; \begin{pmatrix} d_r, y, b \end{pmatrix}_R^{-2/3}, \begin{pmatrix} \nu_e \\ e^- \end{pmatrix}_L^{-1}, \begin{pmatrix} e^- \\ \nu_e \end{pmatrix}_R^{-2}$$

$$Q_{em} = I_{3L} + Y_W/2$$

5 disconnected multiplets in 1 Family // Y_W ? //

Q_{em} ? // $Q_{e^-} = -Q_p$ // Co-exist of (q, l) // g_1, g_2, g_3 ?



$$G(224) = [SU(2)_L \times SU(2)_R \times SU(4)^C] \otimes (L \leftrightarrow R)$$

$$F_{L,R}^e = \begin{bmatrix} u_r & u_y & u_b & \nu_e \\ d_r & d_y & d_b & e^- \end{bmatrix}_{L,R}$$

All 16 in one L-R Conj. mult / Y_W / Q_{em} quantized /

$Q_{e^-} = -Q_p$ / $\{q, l\}$ unif / (W, E, S) / $\boxed{V_R}$ / $\boxed{B-L}$ Generator

$$Q_{em} = I_{3L} + I_{3R} + \frac{B-L}{2}$$



$$SO(10): 16$$

(1 gauge coupling / $g_1 = g_2 = g_3$ at M_U)

(Georgi // Fritzsch, Minkowski (74/75))

FCP & Salam 72-73

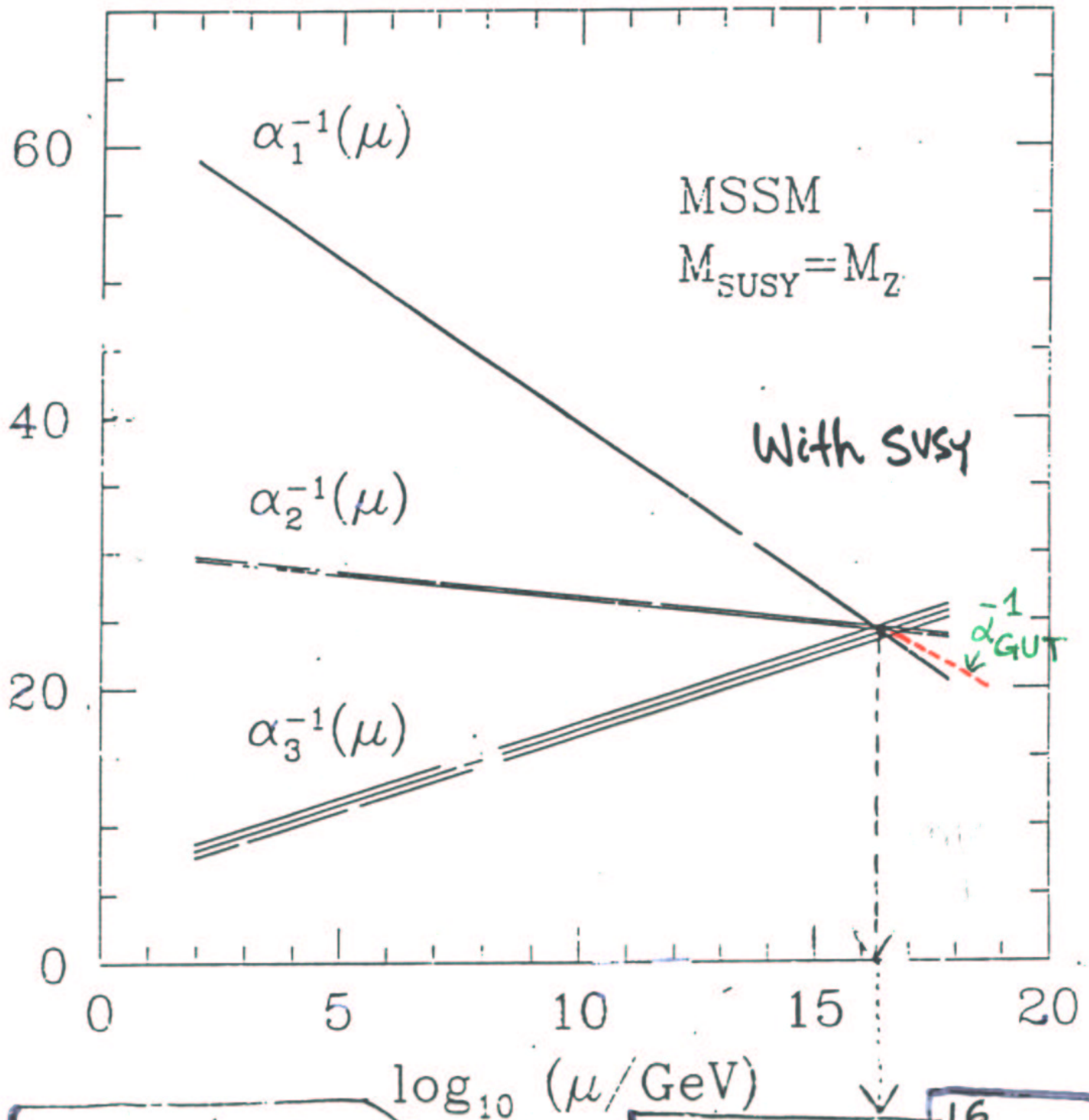
$\bar{5}(213) \rightarrow SU(5)$

$$\bar{5} + 10$$

Georgi & Glashow (74)

NO V_R , NO B-L

② Gauge Coupling Unification



Supports SUSY Unification

$M_{\text{GUT}} \approx 2 \times 10^{16} \text{ GeV}$

$$\sin^2 \Theta_W(m_Z)_{\text{th}} = 0.2315 \pm 0.003$$

$$\alpha_3(m_Z) = 0.125 - 0.13$$

$$\sin^2 \Theta_W(m_Z)_{\text{EXPT}} = 0.23124 \pm 0.00017$$

$$\alpha_3(m_Z)_{\text{obs}} = 0.118 \pm 0.003$$

$$\textcircled{3} \sqrt{\Delta m^2(\nu_\mu \nu_\tau)_{\text{SuperK}}} \approx \frac{1}{20} \text{ eV}$$

SeeSaw

$$m(\nu_L^\tau) \approx \frac{m(\nu_{\text{Dirac}}^\tau)^2}{M(\nu_R^\tau)}$$

Yanagida //
Gell-Mann et al //
Mohapatra &
Senjanovic

$$\textcircled{a} m(\nu_{\text{Dirac}}^\tau) \approx m_t (M_X) \approx 120 \text{ GeV} \leftarrow \begin{matrix} \text{SU(4) - Color} \\ \text{SU(5), [SU(3)]}^3 \end{matrix}$$

(b) Get $M(\nu_R^\tau)$ from SUSY Uni f. Scale: $M_X \approx 2 \times 10^{16} \text{ GeV}$

$$f_{33} \frac{16_3 16_3 \langle \bar{16}_H \rangle \langle \bar{16}_H \rangle}{M_{\text{Pl}} \rightarrow 2 \times 10^{18} \text{ GeV}} \Rightarrow M(\nu_R^\tau) \approx \frac{(2 \times 10^{16} \text{ GeV})^2}{2 \times 10^{18} \text{ GeV}} \approx 2 \times 10^{14} \text{ GeV}$$

(≈ 1)

$$m(\nu_L^\tau) \approx \frac{(120 \text{ GeV})^2}{2 \times 10^{14} \text{ GeV}} \approx \left(\frac{1}{17} \text{ eV}\right) \left(\frac{1}{2} \text{ to } 2\right)$$

$$\text{Also get } m(\nu_L^\mu) \sim \frac{m(\nu_L^\tau)}{10} \Rightarrow \sqrt{\Delta m^2(\nu_\mu \nu_\tau)_{\text{th}}} \sim \left(\frac{1}{17} \text{ eV}\right) \left(\frac{1}{2} - 2\right)$$

Thus SuperK result brings to light the existence of ν_R // reinforces the ideas of

(a) SUSY Unif, (b) SU(4) - Color, & (c) SeeSaw.

In short, just this single piece of information ($\Delta m^2(\nu_\mu \nu_\tau) \sim (\frac{1}{20} \text{ eV})^2$)

Disfavors

- ~~SU(5), [SU(3)]³~~
- ~~Intermediate Scale ($\leq 10^{12}$ GeV) Breaking of B-L~~
- ~~Large Extra Dim.~~

Favors

SU(2)_L x SU(2)_R x SU(4)^C / SO(10) Route
 To Higher Unification with SUSY //
 Single step Breaking of G(224) / SO(10) at $M_X \sim 2 \times 10^{16}$ GeV

④ Fermion Masses & Mixings in SO(10)

Minimal Higgs For SO(10) - Breaking

$45_H, 16_H, \bar{16}_H, 10_H, 126_H, 120_H$

Allowed by String Solns

Too large GUT-scale Threshold Corrections // Also Seems disallowed by strings

SO(10)

$\langle 45_H \rangle \propto B-L$
 $\sim M_{GUT}$

$\langle 16_H \rangle \sim \bar{\nu}_R$
 Breaks (B-L)
 $\sim M_{GUT}$

Babu, Pati & Wilczek (1998/2000)

$SU(2)_L \times U(1)_Y \times SU(3)_C$

$\langle 10_H \rangle \approx v_{EW} \approx 247 \text{ GeV}$

$U(1)_{em} \times SU(3)_C$

- (i) • $16_i 16_j 10_H \rightarrow$ Symmetric in $i \leftrightarrow j$, indep of B-L
- (ii) • But need B-L dep. $\rightarrow m_s^0 \sim m_\mu^0 / 3, \dots$
- (iii) • Need Antisymm. Piece to explain smallness of $V_{bc} \approx 0.04$ & largeness of $Q_{\nu_e \nu_e}^{osc}$
- (iv) Need M_U Not proportional to M_D : $V_{CKM} \neq 1$

But 10_H can't give (ii), (iii) & (iv).

A Concrete Example: Minimal Higgs: $\{45_H, 16_H, \bar{16}_H, 10_H\} + "S"$ 20

only μ & τ Families (Flavor sym: $\mu \neq \tau$)

$$\begin{aligned}
 \mathcal{L}_{\text{mass}} &= h_{33} 16_3 16_3 \langle 10_H \rangle \rightarrow \text{3rd Family: } m_b^0 = m_e^0 \\
 &\propto \textcircled{1} \leftarrow + h_{23} 16_2 16_3 \langle 10_H \rangle \langle S/M \rangle \rightarrow \text{2nd Family} \\
 &\propto \textcircled{5} \leftarrow \text{ANTISYMMETRIC, } \propto B-L \\
 &\propto \textcircled{E} \leftarrow + a_{23} 16_2 16_3 \langle 10_H \rangle \langle 45_H \rangle \rightarrow m_\mu^0 \neq m_s^0 \\
 &\qquad\qquad\qquad M_i \\
 &\propto \textcircled{\eta} \leftarrow + g_{23} 16_2 16_3 \langle 16_H \rangle \langle 16_H \rangle_{EW}^D / M \\
 &\qquad\qquad\qquad \hookrightarrow \text{CKM} \neq 1
 \end{aligned}$$

$10_H \rightarrow (2 \leftrightarrow 2) \ll (2 \leftrightarrow 3) \ll (3, 3)$ Flavor Sym

$$U = \begin{pmatrix} c & t \\ 0 & \epsilon + \sigma \\ -\epsilon + \sigma & 1 \end{pmatrix} m_U^0 \quad \bigg| \quad D = \begin{pmatrix} s & b \\ 0 & \epsilon + \eta \\ -\epsilon + \eta & 1 \end{pmatrix} m_D^0$$

$$N_{\text{Dirac}} = \begin{pmatrix} \nu_\mu & \nu_\tau \\ 0 & -3\epsilon + \sigma \\ +3\epsilon + \sigma & 1 \end{pmatrix} m_U^0 \quad \bigg| \quad L = \begin{pmatrix} \mu & \tau \\ 0 & -3\epsilon + \eta \\ 3\epsilon + \eta & 1 \end{pmatrix} m_D^0$$

Note $q-l$ correlation (SU(4) - color)
 up-down " (SU(2)_L × SU(2)_R)

$\eta \equiv \hat{\eta} + \sigma$

Including $m_U^0 \rightarrow$ 7 param ($\eta, \epsilon, \sigma, \eta', \epsilon', m_U^0, m_D^0$) 22
 describing $9 \times 4 = 36$ entries \rightarrow Will it work?

Input! Assume all param real for a moment

$$m_t^{\text{phys}} = 174 \text{ GeV}; m_c(m_c) = 1.37 \text{ GeV},$$

$$m_s(1 \text{ GeV}) = 116 \text{ MeV}, m_u, m_d, m_u(M_X) = 1.5 \text{ MeV}, m_e$$

$$\sigma \approx 0.110, \eta \approx 0.151, \epsilon \approx -0.095,$$

$$\epsilon' = \sqrt{m_u/m_c} (m_c/m_t) \approx 2 \times 10^{-4}; \eta' = \sqrt{m_e/m_u} (m_c/m_t) \approx 2 \times 10^{-4}$$

$$m_U^0 = m_t(M_X) \approx 110 \text{ GeV}; m_D^0 \approx 1.5 \text{ GeV}$$

Majorana Mass of ν_R^i 's: $f_{ij} 16_i 16_j \bar{16}_H \bar{16}_H / M$

$$\Rightarrow f_{ij} (\nu_R^{iT} \bar{e}^i \nu_R^j) < (16_H)^2 / M$$

$$M_R^\nu = \begin{pmatrix} x & 0 & z \\ 0 & 0 & y \\ \cancel{z} & y & 1 \end{pmatrix}$$

determined by $m_{\nu_2}/m_{\nu_3} \approx 1/8$

M_R

calculated $\approx 5 \times 10^{14} \text{ GeV}$.

6 New observables.

Summary on Fermion Masses & Mixings

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Predictions

$$m_b(m_b) \approx (4.7 - 4.9) \text{ GeV}$$

$$m(\nu_\tau) \approx \left(\frac{1}{20} \text{ eV}\right) \left(\frac{1}{2} - 2\right)$$

$$V_{cb} \approx 0.042$$

$$\sin^2 2\theta_{\nu_\mu \nu_\tau}^{\text{osc}} \approx 0.92 \leftrightarrow \boxed{0.99}$$

$$V_{us} \approx 0.22$$

$$V_{ub} \approx 0.0032$$

$$m_d(1 \text{ GeV}) \approx 8 \text{ MeV}$$

Observations

$$\approx 4.2 \text{ GeV}$$

$$\approx (1/15 - 1/25) \text{ eV} \quad (*)$$

$$\approx 0.04$$

$$\approx 0.92 \leftrightarrow 1$$

$$\approx 0.21$$

$$\approx 0.003 - 0.004$$

$$\approx 8 - 10 \text{ MeV}$$

$$m(\nu_\mu) \approx (2 - 10) \times 10^{-3} \text{ eV} \leftrightarrow \begin{cases} \text{SMA} \sim 3 \times 10^{-3} \text{ eV} \\ \boxed{\text{LMA} \approx 7 \times 10^{-3} \text{ eV}} \end{cases} \quad (*)$$

$$m(\nu_e) \sim (1 \text{ to few}) \times 10^{-3} \text{ eV}$$

Consistent with the framework

$$M(\nu_R^c, \nu_R^u, \boxed{\nu_R^e}) \approx (5 \times 10^{14}, 10^{12}, \boxed{10^9 - 10^{10}}) \text{ GeV}$$

Just right for Lepto/Baryogenesis

SMA or LMA ?

(*) Assuming hierarchical pattern: $m(\nu_L^e) \ll m(\nu_L^u) \ll m(\nu_L^c)$

Predictions

Writing only for 2×2 (for simplicity)

$$U = \begin{pmatrix} c & t \\ 0 & \epsilon + \sigma \\ -\epsilon + \sigma & 1 \end{pmatrix} m_U^0 \quad D = \begin{pmatrix} s & b \\ 0 & \epsilon + \eta \\ -\epsilon + \eta & 1 \end{pmatrix} m_D^0$$

$$N = \begin{pmatrix} 0 & -3\epsilon + \sigma \\ 3\epsilon + \sigma & 1 \end{pmatrix} m_U^0 \quad L = \begin{pmatrix} \mu & \tau \\ 0 & -3\epsilon + \eta \\ 3\epsilon + \eta & 1 \end{pmatrix} m_D^0$$

$$m_b^0 \approx m_\tau^0 (1 - 8\epsilon^2) \Rightarrow m_b(m_b) \approx 4.7 \text{ TeV}$$

$$V_{cb} = \left| \sqrt{\frac{m_s}{m_b}} \left(\frac{\eta + \epsilon}{\eta - \epsilon} \right)^{1/2} - \sqrt{\frac{m_c}{m_t}} \left(\frac{\sigma + \epsilon}{\sigma - \epsilon} \right)^{1/2} \right| = |\sigma - \eta|$$

(0.156) (1/2.2) = 0.042

Suppressed

ENHANCED ≈ 1.8

$$\theta_{\nu_\mu \nu_\tau}^{\text{osc}} = |\theta_{\mu\tau}^l - \theta_{\mu\tau}^{\nu}| = \left| \sqrt{\frac{m_\mu}{m_\tau}} \left(\frac{\eta - 3\epsilon}{\eta + 3\epsilon} \right)^{1/2} + \sqrt{\frac{m_{\nu_2}}{m_{\nu_3}}} \right|$$

$$= 0.437 + \sqrt{m_{\nu_2}/m_{\nu_3}}$$

$$\Rightarrow \sin^2 2\theta_{\nu_\mu \nu_\tau}^{\text{osc}} = 0.92 \leftrightarrow 0.99 \quad \text{Expt } 0.92 \leftrightarrow 1$$

$$m_{\nu_2}/m_{\nu_3} = \left(\frac{1}{15} \right) \leftrightarrow \left(\frac{1}{8} \right) \text{ LMA.}$$

Summary on Fermion Masses & Mixings in the $G(224)/SO(10)$ Framework

Given the bizarre pattern of masses & mixings of quarks, charged leptons and neutrinos, it seems remarkable that the simple pattern of fermion mass matrices^{*}, motivated in large part by the group th of $G(224)/SO(10)$ and the assumption of minimality of the system of \wedge Higgs, makes γ predictions in agreement with observation.

- Study Proton Decay, CP within this framework
- Extend to ESSM

^{*} Need to understand the origin of flavor symmetries.
→ Hierarchical entries.

SMA or LMA?

$f_{ij} 16_i 16_j \bar{16}_H \bar{16}_H / M$

Just with M_R^ν (superheavy Maj. Masses of ν_R 's)

SMA rather generic

$m(\nu_L^e) \sim 2 \times 10^{-5} - 2 \times 10^{-6} \text{ eV} // m(\nu_L^\mu) \sim 3 \times 10^{-3} \text{ eV}$

$\Theta_{\nu_e \nu_\mu}^{osc} = \Theta_{e\mu}^L - \Theta_{e\mu}^\nu \approx 0.05$

situation alters once allow for direct Maj. masses of ν_L 's - Most likely to arise through Higher Dim. op. involving GUT & EW VEV's - Through tiny $\sim 10^{-3} \text{ eV}$ entries \rightarrow IMPORTANT For $(\nu_e - \nu_\mu)$

$W \supset g_{12} 16_1 16_2 16_H 16_H 10_H 10_H / M_{GUT}^3$

$g_{12} (\nu_L^e \nu_L^\mu) (\langle \nu_{RH} \rangle / M_{GUT})^2 \rightarrow (\nu_\mu^2) (\sim 175 \text{ GeV})^2$

$\sim g_{12} (\nu_L^e \nu_L^\mu) (1.5 - 6) \times 10^{-3} \text{ eV} (\langle 16_H \rangle \approx (1-2) M_{GUT})$

$\begin{bmatrix} \nu_L^e & \nu_L^\mu \\ \approx 0 & (3-4) \end{bmatrix} \times 10^{-3} \text{ eV} \Rightarrow \left[\begin{array}{l} \Theta_{\nu_e \nu_\mu}^\nu \approx 1/2 \\ \sin^2 2\Theta_{\nu_e \nu_\mu}^{osc} \approx 0.7 \end{array} \right] \text{ Quite Plausible}$

Thus LMA not strictly a prediction, but perfectly plausible within the framework.

$$m(\nu_L^e \nu_L^e)_{\text{Non-Seesaw}} \sim (2-6) \times 10^{-3} \text{ eV}$$

$$\Theta_{13} \sim \frac{(2-6) \times 10^{-3} \text{ eV}}{5 \times 10^{-2} \text{ eV}}$$

$$\sim 0.03 - 0.1$$

⑤ Leptogenesis $\xleftrightarrow{\text{Sphalerons}}$ Baryogenesis

Kuzmin, Rubakov, Shaposhnikov // Fukugita, Yanagida

Use the Dirac & Majorana Mass-Matrices
of the $G(224)/SO(10)$ -framework

$$\nu_R \rightarrow l + \Phi_H, \quad \bar{l} + \bar{\Phi}_H \quad \Delta(B-L) \neq 0$$

↓
LEPTON - ASYMMETRY

↓ EW Sphalerons

$$(Y_B)_{Th} / (\sin 2\phi_{eff}) \approx (4-37) \times 10^{-11}$$

$$(Y_B)_{obs} \approx (3-9) \times 10^{-11} \quad (\text{BBN})$$

Agrees with observation for natural values
of $\phi_{eff} \approx 1/3 - 1/8$ in accord with
Gravitino Constraint Pati/hep-ph/0209160

Supports the Exist. of ν_R // Spont. Viol of
B-L at High Temp // $G(224)/SO(10)$ Mass Matrices,

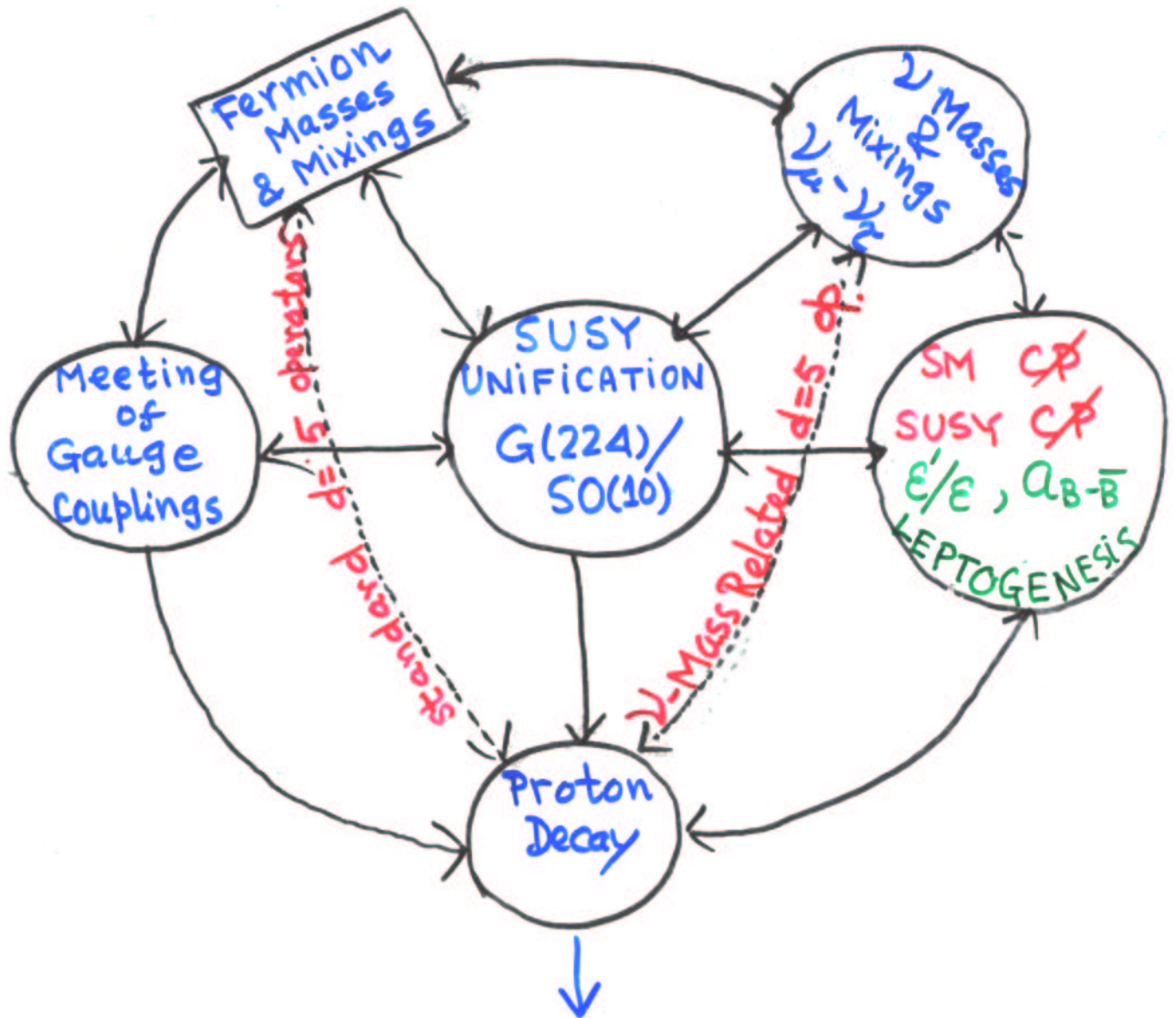
Exist of ν_R // Non-zero ν -Masses &

ν -Oscillation \rightarrow

A NECESSARY SUBTLETY

② UNIFICATION LINKS

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The Missing Link

III Proton Decay: The Hall Mark of Grand Unification

One can argue that a gauge
unification of $\{q \text{ and } l\}$

\Rightarrow (B, L) Non-Conservation

\Downarrow Simplest

Proton Decay

1973

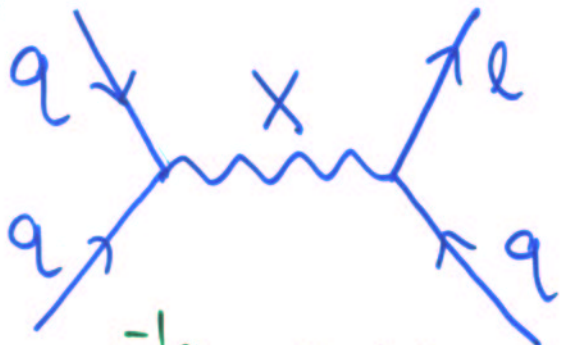
$$\begin{array}{l}
 \text{Proton} \rightarrow e^+ \pi^0, \bar{\nu} K^+, \bar{\nu} \pi^+ \\
 \left. \begin{array}{l} B = +1 \\ L = 0 \end{array} \right\} \begin{array}{l} 0 \quad 0 \\ -1 \quad 0 \end{array} \left. \begin{array}{l} \\ \\ \end{array} \right\} \begin{array}{l} \Delta B \neq 0 \\ \Delta L \neq 0 \end{array}
 \end{array}$$

Expt. (SuperK) $\tau^{-1}(p \rightarrow e^+ \pi^0) \gtrsim 6 \times 10^{33} \text{ yrs}$
 $\tau^{-1}(p \rightarrow \bar{\nu} K^+) \gtrsim 2 \times 10^{33} \text{ yrs}$

Minimal Non-SUSY SU(5) $\left. \begin{array}{l} \tau^{-1}(p \rightarrow e^+ \pi^0)_{\text{Theory}} < 10^{31} \text{ yrs} \\ \text{Excluded by expt} \\ \text{Also excluded by Gauge Coupling Unif or Prediction of } \sin^2 \theta_w \end{array} \right\}$

d=6 Gauge Mediated: SUSY SU(5)/SUSY

①



Amp ($p \rightarrow e^+ \pi^0$)
 $\propto g^2 / M_X^2$

$\Gamma(p \rightarrow e^+ \pi^0) \approx 10^{35.3 \pm 1.3}$ yrs (Theory)

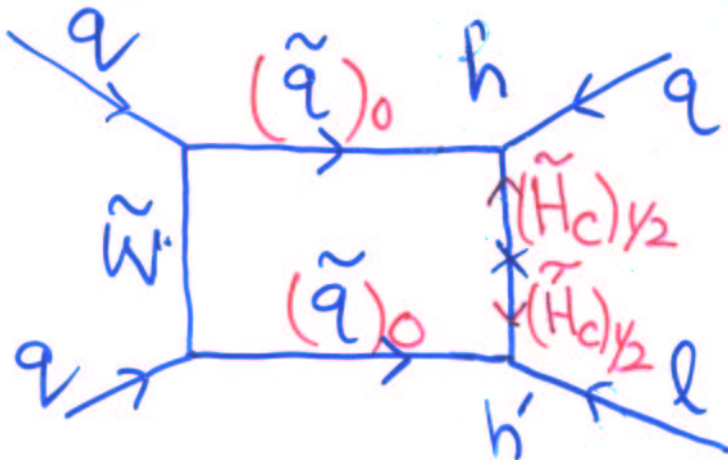
Can be as short as $\approx 10^{34}$ yrs

②

In Supersymmetry

Color triplet Higgsino Mediated

Sakai, Yanagida // Weinberg.



$qqq \rightarrow l$

$\tilde{H}_c \subset 10_H$
 $= (2, 2, 1)_H$
 $+ (1, 1, 6)$

Standard
d=5
operator

Amp $\propto \frac{h h'}{M_{H_c}} \left(\frac{m_{\tilde{W}}}{m_{\tilde{q}}} \right) \alpha_2$

$p \rightarrow \bar{\nu} K^+$

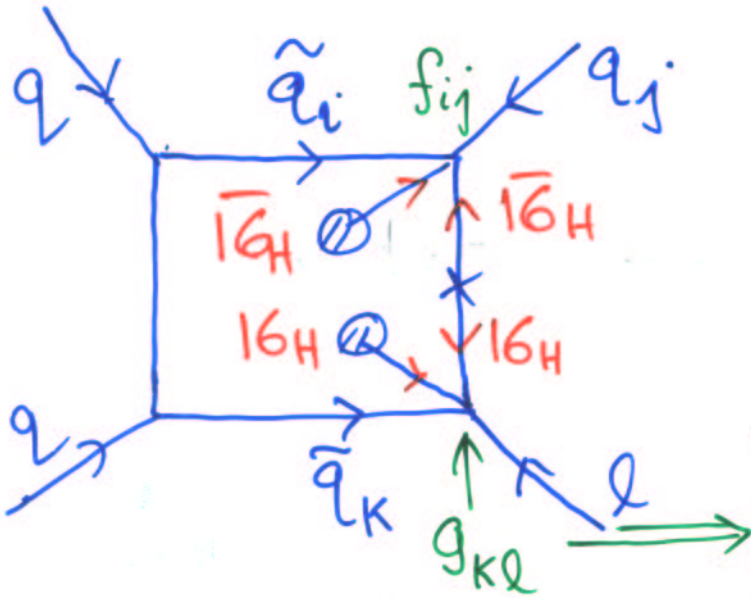
dominant

$\rightarrow \mu^+ K^0$

(For "Standard" d=5, suppressed)

$BR(p \rightarrow \mu^+ K^0)_{std, d=5} \sim 10^{-3}$

③ New d=5 operators \leftrightarrow Majorana Masses of ν_R 's



$$f_{ij} 16_i 16_j \bar{16}_H \bar{16}_H / M$$

Generically $16_i \bar{16}_H$ in 45 & 1 of $SO(10)$

ENTERS INTO DIRAC MASS MATRICES $\leftrightarrow V_{CKM} \neq \mathbb{1}$

Babu, Pati & Wilczek (1997)

$$\Gamma^{-1}(p \rightarrow \bar{\nu} K^+) \text{ New } d=5 \approx 10^{-33} - 10^{-34} \text{ yrs.}$$

$$BR(p \rightarrow \mu^+ K^0) \text{ New } d=5 \approx (10 - 30)\%$$

SUSY
SO(10)
OR
G(224)

Note - that these contribs would generically be present, even if "standard" d=5 operators med. by color triplets $\subset 10_H$ are absent, as in SUSY G(224)

The $\mu^+ K^0$ mode A Signature of this Mechanism

d=5 Rate : Update : JCP-ph/0204240

$$\Gamma(p \rightarrow \bar{\nu} K^+) \approx (0.15 \times 10^{31} \text{ yrs}) \times (0.32/A_L)^2 \times (0.93/A_S)^2 \\ \times \left(\frac{0.014 \text{ GeV}^3}{\beta_H} \right)^2 \times \left(\frac{1/6}{m_{\tilde{W}}/m_{\tilde{Q}}} \right)^2 \times \left(\frac{m_{\tilde{Q}}}{1.2 \text{ TeV}} \right)^2 \times \left[\frac{2 \times 10^{-24} \text{ GeV}^{-1}}{\hat{A}(\bar{\nu})} \right]^2$$

calculate in correlation with

a) fermion Masses & Mixings

b) Unif. Scale Threshold Corr. (Right sign & magnitude)
 $\sim 5\%$ in $\alpha_3(m_Z)$

c) New d=5 op related to ν Masses

① Matrix Element $\beta_H = (0.014 \text{ GeV}^3) (\frac{1}{2} - 2)$

② $\tan\beta \gg 3$

③ $m_{\tilde{Q}} \approx (1.2 \text{ TeV}) (\frac{1}{\sqrt{2}} - 2)$ (Focus point SUSY)

④ $(\frac{m_{\tilde{W}}}{m_{\tilde{Q}}}) \approx (1/6) (\frac{1}{2} - 2)$

⑤ $A_L \approx 0.32$ (2 loop)

⑥ $A_S \approx 0.93$

⇓

$\Gamma_{\text{now}} \approx 16 \Gamma_{\text{before}}$

Summary on proton Decay : d = 5

$\frac{\text{SUSY SU(5)}}{\text{MSSM}}$ } $\bar{\Gamma}^1(p \rightarrow \bar{\nu} K^+) \leq 2 \times 10^{31} \text{ yrs}$ Excluded by SuperK // IMB

$\frac{\text{SUSY SO(10)}}{\text{MSSM}}$ } $\bar{\Gamma}^1(p \rightarrow \bar{\nu} K^+) \leq 1.9 \times 10^{33} \text{ yrs}$ Tightly Constrained (std. d=5)

$\frac{\text{SUSY SO(10)}}{\text{ESSM}}$ } $\bar{\Gamma}^1(p \rightarrow \bar{\nu} K^+) \leq 2 \times 10^{34} \text{ yrs}$ Std. d=5

$\frac{\text{SUSY G(224)/SO(10)}}{\text{MSSM or ESSM}}$ } $\bar{\Gamma}^1(p \rightarrow \bar{\nu} K^+) \approx 10^{33} - 10^{34} \text{ yrs}$ New d=5 (ν Masses)

Last Two fully Compatible with present limits, Must be seen with improvement by factor of 5-10, or else SUSY SO(10) or string G(224), which is otherwise so successful, would have a major set back.

Proton decay : d = 6 : SUSY SU(5)/SO(10) //

and Flipped SU(5) } $\bar{\Gamma}^1(p \rightarrow e^+ \pi^0) \approx 10^{35.3 \pm 1.3} \text{ yrs}$

I) PROBING INTO NEW PHYSICS

1

<u>Process</u>	<u>Tests</u>	<u>Dist. Scale</u>
<u>PROTON DECAY</u> $p \rightarrow \bar{\nu} K^+, \mu^+ K^0, e^+ \pi^0$ $\lambda \sim \frac{Q^2}{M^2}$	$q-l$ unif. $q-\bar{q}, q-\bar{l}, l-\bar{l}$ unif SUSY // Und. Fermion Masses // UNITY OF FORCES	$\sim 10^{-30}$ cm
<u>NEUTRINO MASS</u> $m(\nu_e) \approx \frac{m(\nu_e)_{\text{DIRAC}}^2}{M_R}$ $\approx 1/20$ eV	$M_R \sim 10^{15}$ GeV $\Delta L = 2$; See Saw $q-l$ unif \leftrightarrow SU(4)-col SUSY unif. scale	$\sim 10^{-29}$ cm
<u>$n - \bar{n}$ Oscillation</u>	$\Delta B = 2$	$10^{-19} - 10^{-20}$ cm
<u>$\mu \rightarrow e \gamma, \tau \rightarrow \mu \gamma, \dots$</u>	Flavor Viol, SUSY	$10^{-17} - 10^{-18}$ cm
<u>$\delta(g-2)_\mu$</u>	SUSY // Heavy Fermions	$10^{-16} - 10^{-18}$ cm
<u>$K^0 - \bar{K}^0, B^0 - \bar{B}^0$</u>	Flavor, CP	$10^{-17} - 10^{-18}$ cm
<u>$(edm)_{n,e}$</u>	CP & P	$10^{-17} - 10^{-18}$ cm

Proton Decay is the Hallmark of Grand Unification.

It Provides a unique window to view physics at truly short distances $\lesssim 10^{-30}$ cm.

$P \rightarrow \bar{\nu} K^+$ if seen \Rightarrow $\left\{ \begin{array}{l} q-l, q-\bar{q}, q-\bar{l} \text{ unif} // \\ \text{SUSY unif} // \\ M_G \sim M_X \sim 10^{16} - 10^{17} \text{ GeV} \end{array} \right.$

$P \rightarrow \mu^+ K^0$ if seen \Rightarrow $\left\{ \begin{array}{l} \oplus \text{ Relevance of } \\ \nu\text{-Mass Related operator} \end{array} \right.$

THUS PROTON DECAY IF SEEN WOULD PROVIDE A WEALTH OF KNOWLEDGE THAT CAN NOT BE GAINED BY ANY OTHER MEANS

Discovery Potential High

THIS IS WHY NEED NEXT-GENERATION DETECTOR FOR PROTON DECAY & NEUTRINO-oscillation \rightarrow UNO // HyperKamiokande.