

Compute in H 1s state

$$\begin{aligned}
 7.18 \quad U_{av} &= \int_0^\infty U(r) P(r) dr \\
 &= -\frac{e^2}{4\pi\epsilon_0} \int_0^\infty \frac{1}{r} \cdot r^2 \left(\frac{2}{a_0^{3/2}} e^{-r/a_0}\right)^2 dr \\
 &= -\frac{e^2}{4\pi\epsilon_0} \frac{4}{a_0^3} \int_0^\infty r e^{-2r/a_0} dr \\
 &= -\frac{e^2}{\pi\epsilon_0 a_0^3} \frac{1!}{(2/a_0)^2} \\
 &= -\frac{1}{4\pi\epsilon_0} \frac{1}{a_0}
 \end{aligned}$$

↑ potential for electron in orbit of radius a_0

q. 12 Use the figure
and

$$k = \frac{2(E - E_{\min})}{(R - R_{\text{eq}})^2} \stackrel{\text{define}}{=} \frac{2\Delta E}{\Delta R^2}$$

estimate

$$\Delta E = 0.25 \text{ eV}$$

For

$$\Delta R = 0.5 \text{ nm}$$

or

$$k \approx \frac{0.50 \text{ eV}}{(0.5)^2 \text{ nm}^2} = \frac{5 \times 10^{-1}}{2.5 \times 10^{-3}} \text{ eV nm}^{-2}$$

$$k \approx 2 \times 10^2 \frac{\text{eV}}{\text{nm}^2} \approx 2 \times 10^{20} \frac{\text{eV}}{\text{m}^2}$$

which gives

$$\nu = \frac{1}{2\pi} \sqrt{\frac{k c^2}{m c^2}}$$

$$m = \frac{m_{\text{Na}} m_{\text{Cl}}}{m_{\text{Na}} + m_{\text{Cl}}} = \frac{22 \times 35}{22 + 35} u \approx 14 \text{ u}$$

$$\nu \approx \frac{1}{2\pi} \sqrt{\frac{2 \times 10^{20} \text{ eV}}{\text{m}^2} \times \frac{9 \times 10^{16} \text{ m}^2}{\text{s}^2}}$$

$$\approx \frac{1}{2\pi} \sqrt{\frac{18}{14} \times 10 \times 10^{26}} \text{ Hz}$$

$$\approx \frac{1}{2\pi} 3.5 \times 10^{13} \text{ Hz}$$

$$\approx 5 \times 10^{12} \text{ Hz}$$

$$\lambda = \frac{c}{\nu} \approx 6 \times 10^{-5} \text{ m}$$

infrared (see figure
on p. 6)

$$h\nu \approx (4 \times 10^{-15} \text{ eV.s}) (5 \times 10^{12} \text{ Hz})$$

$$\approx 2 \times 10^{-2} \text{ eV}$$

$$h\nu \approx .004 \text{ eV}$$

at about 0.2 eV,
parabolic approx. fails
corresponds to about n ≈ 10

9.13

The only difference between H_2 and HD is their reduced mass.

$$m_{H_2} = \frac{m_p^2}{2m_p} = \frac{m_p}{2}$$

$$m_{HD} = \frac{m_p(2m_p)}{m_p + 2m_p} = \frac{2}{3} m_p = \frac{4}{3} m$$

R_{eq} is independent of m , since it is determined by electronic states.

The relevant formulas are

$$\gamma = \frac{1}{2\pi} \sqrt{\frac{k}{m}} \quad (k \text{ is also } m\text{-independent})$$

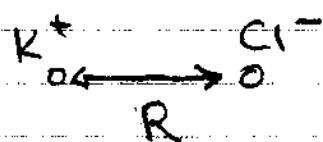
$$\rightarrow \gamma_{HD} = \sqrt{\frac{3}{4}} \gamma_{H_2}$$

Similarly

$$\frac{\hbar^2}{2m_{HD}R_{eq}^2} = \frac{3}{4} \frac{\hbar^2}{2m_{H_2}R_{eq}^2}$$

$$9.3 \quad E_{\text{ion}}^K = 4.34 \text{ eV}$$

$$E_{\text{aff}}^{\text{cl}} = 3.06 \text{ eV}$$



find R such that

$$U(R) = E_{\text{aff}}^K - E_{\text{ion}}^K$$

$$= (3.06 - 4.34) \text{ eV}$$

$$= -1.28 \text{ eV}$$

$$U(R) = -\frac{e^2}{4\pi\epsilon_0} \frac{1}{R} = -1.28 \text{ eV}$$

$$R = (1.28 \text{ eV})^{-1} \left(\frac{e^2}{4\pi\epsilon_0} \right)$$

$$\lambda = 1.440 \text{ nm}$$

$$= 1.13 \text{ nm}$$